

# A VERSATILE THREE-DIMENSIONAL RAY TRACING COMPUTER PROGRAM FOR RADIO WAVES IN THE IONOSPHERE

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## PREFACE

This report documents the latest version of the three-dimensional ray tracing program originally described in "A Three-Dimensional Ray Tracing Computer Program," by R. M. Jones, ESSA Technical Report IER 17-ITSA 17, and later modified in "Modifications to the Three-Dimensional Ray Tracing Program Described in IER 17-ITSA 17," by R. M. Jones, ESSA Technical Memorandum ERLTM-ITS 134. This report replaces all of the material contained in the above two reports.





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A VERSATILE THREE-DIMENSIONAL RAY TRACING  
COMPUTER PROGRAM FOR RADIO WAVES  
IN THE IONOSPHERE

R. Michael Jones\* and Judith J. Stephenson\*\*

This report describes an accurate, versatile FORTRAN computer program for tracing rays through an anisotropic medium whose index of refraction varies continuously in three dimensions. Although developed to calculate the propagation of radio waves in the ionosphere, the program can be easily modified to do other types of ray tracing because of its organization into subroutines.

The program can represent the refractive index by either the Appleton-Hartree or the Sen-Wyller formula, and has several ionospheric models for electron density, perturbations to the electron density (irregularities), the earth's magnetic field, and electron collision frequency.

For each path, the program can calculate group path length, phase path length, absorption, Doppler shift due to a time-varying ionosphere, and geometrical path length. In addition to printing these parameters and the direction of the wave normal at various points along the ray path, the program can plot the projection of the ray path on any vertical plane or on the ground and punch the main characteristics of each ray path on cards.

The documentation includes equations, flow charts, program listings with comments, definitions of program variables, deck set-ups, descriptions of input and output, and a sample case.

Key words: Ray tracing, computer program, radio waves, ionosphere, three-dimensional, Appleton-Hartree formula, Sen-Wyller formula.

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## 1. INTRODUCTION

This report describes a three-dimensional ray tracing program written in FORTRAN language for the CDC-3800 computer. Copies of the program deck are available from the Institute for Telecommunication Sciences.

Earlier versions of this program have been in use now for over nine years, both by us and by people scattered all over the world. During that time we have improved and modified the program to the extent that we now need to document these changes so that the present program will be easier to use. We have included the input parameter forms that we use to request ray path calculations because they give nearly all the necessary input data and describe the electron density, collision frequency, and magnetic field models.

## 2. GENERAL DESCRIPTION

This computer program traces the path of radio wave through a user-specified model of the ionosphere when given the transmitter location (longitude, latitude, and height above the ground), the frequency of the wave, the direction of transmission (both elevation and azimuth), the receiver height, and the maximum number of hops wanted.

## 3. RAY TRACING EQUATIONS

The program calculates ray paths by numerically integrating Hamilton's equations. Lighthill (1965) gives Hamilton's equations in four dimensions (three spatial and one time) for Cartesian coordinates. Haselgrove (1954) gives Hamilton's equations in three dimensions for spherical polar coordinates. Combining the two gives Hamilton's equations in four dimensions in which the three spatial coordinates are spherical polar (see Table 1 for a definition of the symbols):

$$\frac{dr}{d\tau} = \frac{\partial H}{\partial k_r}, \quad (1)$$

Table 1. List of the More Important Symbols

---

A	In section 3, absorption in decibels.
$B_o$	Magnetic induction of earth's magnetic field.
c	Speed of electromagnetic waves in free space.
C	Cosine of the angle of incidence on the ionosphere.
e	Charge of the electron (a negative number).
$F(w)$	$F(w) = w C_{3/2}(w) + i \frac{5}{2} C_{5/2}(w) = \frac{1}{3/2!} \int_0^{\infty} \frac{t^{3/2} \exp(-t) dt}{w - it}$ (Davies, 1965, p. 86)
f	Wave frequency.
$\Delta f$	Frequency shift of a wave due to a time varying ionosphere (sometimes called Doppler shift).
$f_H$	Gyro frequency for electrons, $ e B_o/2\pi m$ .
$f_N$	Plasma frequency, $(Ne^2/4\pi^2 \epsilon_o m)^{1/2}$ .
$G(w)$	$wF(w)$ .
H	Hamiltonian.
$k_r, k_\theta, k_\varphi$	Components of the propagation vector in the r, $\theta$ , $\varphi$ directions -- a vector perpendicular to the wave front having a magnitude $2\pi/\lambda = \omega/v$ .
m	Mass of electron.
N	Number of electrons per unit volume.
n	Phase refractive index (in general complex).
$n'$	Group refractive index (in general complex).
P	Phase path length, phase of wave divided by free space wave number $2\pi/\lambda_o$ .
$P'$	Group path length ct.
r, $\theta$ , $\varphi$	Coordinates of a point in spherical polar coordinates.
s	Geometric ray path length.
S	Sine of the angle of incidence on the ionosphere.

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Table 1. (Continued)

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t	Time, travel time of a wave packet.
U	$1 - iZ$ in the Appleton-Hartree formula or $Z/F(1/Z) = 1/G(1/Z)$ in the Sen-Wyller formula.
$V_r, V_\theta, V_\varphi$	Components of the wave normal direction in the r, $\theta$ , and $\varphi$ directions, normalized so that $V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real}\{n^2\}$ .
v	Phase velocity.
X	$\omega_N^2/\omega^2 = f_N^2/f^2 = Ne^2/(\epsilon_0 m\omega^2)$ .
Y	$\omega_H/\omega = f_H/f$ .
$Y_L$	$Y \cos \psi$ .
$Y_T$	$Y \sin \psi$ .
Z	$u/\omega$ or $u_m/\omega$ .
$\epsilon_0$	Electric permittivity of free space.
$\theta$	Colatitude in spherical polar coordinates.
$\lambda$	Wavelength.
$\lambda_0$	Wavelength in free space.
$\nu$	Electron collision frequency.
$\nu_m$	Mean electron collision frequency.
$\rho$	Characteristic wave polarization (definition in Table 6).
$\rho_L$	Longitudinal polarization (definition in Table 6).
$\tau$	Independent variable in Hamilton's equations.
$\varphi$	Longitude in spherical polar coordinates.
$\psi$	Angle between wave normal and $-B_0$ .
$\omega$	$2\pi f$ , angular wave frequency.
$\Delta\omega$	$2\pi\Delta f$ , angular frequency shift.
$\omega_H$	$2\pi f_H =  e B_0/m$ , angular gyrofrequency.
$\omega_N$	$2\pi f_N = (Ne^2/\epsilon_0 m)^{\frac{1}{2}}$ , angular plasma frequency.

---

$$\frac{d\theta}{d\tau} = \frac{1}{r} \frac{\partial H}{\partial k_{\theta}} , \quad (2)$$

$$\frac{d\varphi}{d\tau} = \frac{1}{r \sin\theta} \frac{\partial H}{\partial k_{\varphi}} , \quad (3)$$

$$\frac{dt}{d\tau} = - \frac{\partial H}{\partial \omega} , \quad (4)$$

$$\frac{dk_r}{d\tau} = - \frac{\partial H}{\partial r} + k_{\theta} \frac{d\theta}{d\tau} + k_{\varphi} \sin\theta \frac{d\varphi}{d\tau} , \quad (5)$$

$$\frac{dk_{\theta}}{d\tau} = \frac{1}{r} \left( - \frac{\partial H}{\partial \theta} - k_{\theta} \frac{dr}{d\tau} + k_{\varphi} r \cos\theta \frac{d\varphi}{d\tau} \right) , \quad (6)$$

$$\frac{dk_{\varphi}}{d\tau} = \frac{1}{r \sin\theta} \left( - \frac{\partial H}{\partial \varphi} - k_{\varphi} \sin\theta \frac{dr}{d\tau} - k_{\theta} r \cos\theta \frac{d\theta}{d\tau} \right) , \quad (7)$$

$$\frac{d\omega}{d\tau} = \frac{\partial H}{\partial t} , \quad (8)$$

The variables  $r$ ,  $\theta$ ,  $\varphi$  are the spherical polar coordinates of a point on the ray path;  $k_r$ ,  $k_{\theta}$ , and  $k_{\varphi}$  are the components of the propagation vector (wave normal direction normalized so that in free space

$$k_r^2 + k_{\theta}^2 + k_{\varphi}^2 = \frac{\omega^2}{c^2} , \quad (9)$$

where  $\omega = 2\pi f$  is the angular frequency of the wave and  $c$  is the speed of propagation of electromagnetic waves in free space);  $t$  is time, in (4) it is the propagation time of a wave packet, in (8) it expresses the variation with time of a time varying medium;  $\tau$  is a parameter whose value depends on the choice of the Hamiltonian  $H$ .

For actual calculation, the ray tracing program uses group path  $P' = ct$  as the independent variable because the derivatives with respect to  $P'$  are independent of the choice of Hamiltonian, allowing the program to switch Hamiltonians in the middle of a path. This choice automatically causes the program to take smaller steps in real path length near reflection where the calculations are more critical. The resulting equations obtained by dividing (1) through (8) by  $c$  times (4) are:

$$\frac{dr}{dP'} = -\frac{1}{c} \frac{\partial H / \partial k_r}{\partial H / \partial \omega} , \quad (9)$$

$$\frac{d\theta}{dP'} = -\frac{1}{rc} \frac{\partial H / \partial k_\theta}{\partial H / \partial \omega} , \quad (10)$$

$$\frac{d\varphi}{dP'} = -\frac{1}{rc \sin\theta} \frac{\partial H / \partial k_\varphi}{\partial H / \partial \omega} , \quad (11)$$

$$\frac{dk_r}{dP'} = \frac{1}{c} \frac{\partial H / \partial r}{\partial H / \partial \omega} + k_\theta \frac{d\theta}{dP'} + k_\varphi \sin\theta \frac{d\varphi}{dP'} , \quad (12)$$

$$\frac{dk_\theta}{dP'} = \frac{1}{r} \left( \frac{1}{c} \frac{\partial H / \partial \theta}{\partial H / \partial \omega} - k_\theta \frac{dr}{dP'} + k_\varphi r \cos\theta \frac{d\varphi}{dP'} \right) , \quad (13)$$

$$\frac{dk_\varphi}{dP'} = \frac{1}{r \sin\theta} \left( \frac{1}{c} \frac{\partial H / \partial \varphi}{\partial H / \partial \omega} - k_\varphi \sin\theta \frac{dr}{dP'} - k_\theta r \cos\theta \frac{d\theta}{dP'} \right) , \quad (14)$$

$$\frac{d(\Delta f)}{dP'} = \frac{1}{2\pi} \frac{d\Delta\omega}{dP'} = \frac{1}{2\pi} \frac{d\omega}{dP'} \quad (15)$$

$$= -\frac{1}{2\pi} \frac{\partial H / \partial t}{\partial H / \partial \omega} .$$

Equation (15) for the frequency shift of a wave propagating through a time varying medium follows directly from Hamilton's equations (4) and (8). An alternative derivation is given by Bennett (1967). For large frequency shifts, the frequency shift should be accumulated along the ray path and the shifted frequency used in calculations at each point on the ray path. Equations (1) through (8) imply that all eight dependent variables vary along the path, and that at each point on the path the instantaneous value of all parameters (including frequency) is used in further evaluations of the equations. However, the time variation of the ionosphere due to natural causes (such as solar flares) is so slow that the resulting frequency shifts are small enough (less than one part in  $10^5$ ) to have negligible effect on the propagation. For this reason, the program calculates frequency shift to compare with frequency shift measurements, but does not adjust the carrier frequency of the wave used in the propagation calculations.

The first six differential equations (9) through (14) are always integrated. The user can choose whether to have the program integrate (15) to calculate the frequency shift.

There are three other quantities that can be calculated by integration along the ray path. The phase path  $P$  (phase divided by the free space wavenumber  $2\pi/\lambda_0 = \omega/c$ ) is calculated by integrating

$$\begin{aligned} \frac{dP}{dP'} &= \frac{c}{\omega} \left( k_r \frac{dr}{dP'} + k_\theta r \frac{d\theta}{dP'} + k_\varphi r \sin\theta \frac{d\varphi}{dP'} \right) \\ &= - \frac{1}{\omega} \frac{k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi}}{\partial H / \partial \omega} . \end{aligned} \quad (16)$$

If the absorption per wavelength is small (as it must be for this type of ray tracing to be valid), then an approximate formula can be integrated

to give the absorption in decibels

$$\begin{aligned} \frac{dA}{dP'} &= -\frac{10}{\log_e 10} \frac{\omega}{c} \frac{\text{imag} \left( \frac{\omega^2}{c^2} n^2 \right)}{k_r^2 + k_\theta^2 + k_\varphi^2} \frac{dP}{dP'} \\ &= \frac{10}{\log_e 10} \frac{\text{imag} \left( \frac{\omega^2}{c^2} n^2 \right)}{k_r^2 + k_\theta^2 + k_\varphi^2} \frac{k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi}}{c \partial H / \partial \omega}, \end{aligned} \quad (17)$$

where  $n$  is the (complex) phase refractive index. The geometrical path length of the ray can be calculated by integrating

$$\begin{aligned} \frac{ds}{dP'} &= \sqrt{\left( \frac{dr}{dP'} \right)^2 + r^2 \left( \frac{d\theta}{dP'} \right)^2 + r^2 \sin^2 \theta \left( \frac{d\varphi}{dP'} \right)^2} \\ &= \frac{\sqrt{\left( \frac{\partial H}{\partial k_r} \right)^2 + \left( \frac{\partial H}{\partial k_\theta} \right)^2 + \left( \frac{\partial H}{\partial k_\varphi} \right)^2}}{c \partial H / \partial \omega}. \end{aligned} \quad (18)$$

The user can choose to have frequency shift, phase path, absorption, or path length calculated using equations (15), (16), (17), or (18) and printed by setting the appropriate value in the input W array. (W59, W57, W58, W60 in Table 2.)

If the user wants to add differential equations to the program, he can do so by modifying subroutine HAMLTN, which evaluates Hamilton's equations.

The Hamiltonian and its derivatives are calculated by one of the versions of subroutine RINDEX, which also calculates the phase refractive index and its derivatives.

#### 4. CHOOSING AND CALCULATING THE HAMILTONIAN

Because Hamilton's equations guarantee that the Hamiltonian is constant along the ray path and because it is desirable to have the dispersion



relation satisfied at each point on the ray path, it is usual to write the dispersion relation in the form  $H = \text{constant}$  and choose that  $H$  as the Hamiltonian. Two problems arise. First, in a lossy medium the dispersion relation is complex, so that the resulting complex Hamiltonian gives ray paths having complex coordinates when used in Hamilton's equations. Second, in some cases some forms of the dispersion relation have computational advantages over others when used as a Hamiltonian.

Allowing the coordinates of the ray path to assume complex values is called ray tracing in complex space (Budden and Jull, 1964; Jones, 1970; Budden and Terry, 1971) which is the extension to three dimensions of the phase integral method (Budden, 1961). Ray tracing in complex space is necessary to calculate the propagation of LF radio waves in the D region of the ionosphere (Jones, 1970), and it may also be needed for some medium frequencies.

However, the effect of losses on the ray path of HF radio waves in the ionosphere is probably small, so that the only effect of losses is to attenuate the signal. For this case, then, it is desirable to find a prescription for calculating ray paths having real coordinates. Several methods exist for doing this, and except for computational difficulties, one is probably as good as another. One should recognize that along the ray path:

- (1) the dispersion cannot be exactly satisfied, or
- (2) Hamilton's equations cannot be satisfied, or
- (3) both of the above.

In our program, we have chosen to keep Hamilton's equations and require only the real part of the dispersion relation to be satisfied, neglecting the imaginary part. Another approach (Suchy, 1972) is to alter Hamilton's equations so that the full complex dispersion relation is still satisfied along a ray path having real coordinates. We are

reasonably certain that for any situation in which Suchy's method gives significantly different answers from ours, neither method is valid; ray tracing in complex space or an equivalent method would then be required.

Three choices for the Hamiltonian illustrate the computational difficulties involved. Haselgrove (1954) used the following Hamiltonian

$$H = \frac{c}{\omega} \frac{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}}{\text{real}(n)} - 1, \quad (19)$$

which, except for the effects of errors in the numerical integration and the value of the independent variable, is equivalent to

$$\begin{aligned} H &= 1 - \frac{\omega}{c} \frac{\text{real}(n)}{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}} \\ &= \text{real} \left\{ 1 - \frac{\omega}{c} \frac{n}{(k_r^2 + k_\theta^2 + k_\varphi^2)^{\frac{1}{2}}} \right\}. \end{aligned} \quad (20)$$

There are eight versions of the subroutine RINDEX which calculate the Hamiltonian and its partial derivatives. (Eight versions allow the user to choose the Appleton-Hartree formula or the Sen-Wyller formula, and to include or ignore the earth's magnetic field and collisions.) Six of these versions (subroutines AHFWFC, AHWFNC, AHNFWC, AHNFNC, SWWF, and SWNF) use the following Hamiltonian:

$$\begin{aligned} H &= \frac{1}{2} \left( \frac{c^2}{\omega^2} (k_r^2 + k_\theta^2 + k_\varphi^2) - \text{real}(n^2) \right) \\ &= \text{real} \left\{ \frac{1}{2} \left( \frac{c^2}{\omega^2} (k_r^2 + k_\theta^2 + k_\varphi^2) - n^2 \right) \right\}. \end{aligned} \quad (21)$$

The other two versions (subroutines BQFWFC and BQWFNC) use as a Hamiltonian the real part of the quadratic equation whose solution is the Appleton-Hartree formula (Budden, 1961)

$$\begin{aligned}
H = \text{real} \left\{ & [(U - X) U^2 - Y^2 U] c^4 k^4 + X(k \cdot Y)^2 c^4 k^2 + \right. \\
& + [-2U(U - X)^2 + Y^2(2U - X)] c^2 k^2 \omega^2 - X(k \cdot Y)^2 c^2 \omega^2 + \\
& \left. + [(U - X)^2 - Y^2] (U - X) \omega^4 \right\} \tag{22}
\end{aligned}$$

except in or near free space (defined by  $X < 0.1$ ) where they also use (21) as the Hamiltonian. In (22),  $U = 1 - iZ$ , and  $X$ ,  $Y$ , and  $Z$  are the usual magnetoionic parameters.

In a lossy medium, the Hamiltonians in (20), (21), and (22) determine slightly different ray paths, but the differences are significant only when it is no longer valid to represent ray paths with coordinates that are real rather than complex. In fact, this is a weak criterion. The ray paths determined by these three Hamiltonians will become invalid before there are noticeable differences between the three ray paths. In a lossless medium, the above three Hamiltonians determine identical ray paths (except for integration errors).

For either a lossy or lossless medium, some of the above three Hamiltonians have computational difficulties. Special care must be taken in using (19) or (20) in an evanescent region (which is frequently necessary at or near vertical incidence because the numerical integration subroutine usually requires the evaluation of the differential equations not only on the ray path, but also at points near the ray path). For instance, in a lossless medium,  $\text{real}(n)$  is zero in an evanescent region, which leads to problems in (19) and (20). This problem will not arise in (21) because  $\text{real}(n^2)$  is well behaved in or at the boundary of an evanescent region, nor will it occur in using (22).

Neither (20) nor (21) (nor any other Hamiltonian based on the refractive index) will work for a ray passing through a spitze (Davies, 1965, p. 202) because the refractive index is indeterminate at a spitze, and

some of the derivatives of  $n$  diverge. So far, we have had no problems using (22) to calculate ray paths through a spitze with or without collisions.

However, the Hamiltonian in (22) will not work in or near free space because all of its derivatives are zero in free space. This problem is related to (22) not being able to distinguish between ordinary and extraordinary waves. To get started, the program uses (21) until the electron density is large enough that  $X$  is equal or greater than  $1/10$ .

As far as we can tell, the AHWFWC (Appleton-Hartree, with field, with collisions) version of subroutine RINDEX has been made obsolete by the BQWFWC (Booker quartic, with field, with collisions) version. The latter will do everything the AHWFWC version will do and in addition it will calculate rays through spitzes. A few trial runs, however, indicate that AHWFWC runs about 30 percent faster than BQWFWC.

Similarly, the AHWFNC (Appleton-Hartree, with field, no collisions) version has been made obsolete by the BQWFNC version, which apparently runs just as fast as the AHWFNC version. We are continuing to include the AHWFNC version just in case there are undiscovered problems with the BQWFNC version.

In addition to the Appleton-Hartree formula, which is based on a constant collision frequency, the program also includes the generalized formula of Sen and Wyller (1960), which assumes a Maxwell-Boltzman distribution of electron energy and a collision frequency proportional to energy. Two versions of subroutine RINDEX use the Sen-Wyller formula for calculating the refractive index and the resulting Hamiltonian with its derivatives. These are SWWF, which includes the effects of the earth's magnetic field, and SWNF, which neglects the Earth's magnetic field. The SWWF version will probably not work for calculating rays through a spitze. It would be possible to make a version which used as

its Hamiltonian the quadratic equation whose solution is the Sen-Wyller formula for calculating rays through a spitze, but it is unlikely that we will ever do that.

The versions of subroutine RINDEX that use (21) for a Hamiltonian use the following formulas for calculating the derivatives of that Hamiltonian.

$$\frac{\partial H}{\partial t} = -n \frac{\partial n}{\partial t} , \quad (23)$$

$$\frac{\partial H}{\partial r} = -n \frac{\partial n}{\partial r} , \quad (24)$$

$$\frac{\partial H}{\partial \theta} = -n \frac{\partial n}{\partial \theta} , \quad (25)$$

$$\frac{\partial H}{\partial \varphi} = -n \frac{\partial n}{\partial \varphi} , \quad (26)$$

$$\frac{\partial H}{\partial \omega} = -\frac{n n'}{\omega} , \quad (27)$$

$$\frac{\partial H}{\partial k_r} = \frac{c^2}{\omega^2} k_r - \frac{c}{\omega} n \frac{\partial n}{\partial V_r} , \quad (28)$$

$$\frac{\partial H}{\partial k_\theta} = \frac{c^2}{\omega^2} k_\theta - \frac{c}{\omega} n \frac{\partial n}{\partial V_\theta} , \quad (29)$$

$$\frac{\partial H}{\partial k_\varphi} = \frac{c^2}{\omega^2} k_\varphi - \frac{c}{\omega} n \frac{\partial n}{\partial V_\varphi} , \quad (30)$$

$$\vec{k} \cdot \frac{\partial H}{\partial \vec{k}} = k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi} = n^2 , \quad (31)$$

where  $n'$  is the group refractive index defined by

$$n' = n + f \frac{dn}{df} = n + \omega \frac{dn}{d\omega} , \quad (32)$$

and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions normalized so that

$$V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real} \{n^2\} . \quad (33)$$

The derivatives of the Hamiltonian in (22) are given in section 5. 5.

## 5. REFRACTIVE INDEX EQUATIONS

The refractive index equations used in this ray tracing program are based either on the Appleton-Hartree formula (Budden, 1961) or on the generalized formula of Sen and Wyller (1960). There are eight versions of SUBROUTINE RINDEX, the subroutine that calculates the refractive index and its gradient:

- (1) Appleton-Hartree formula with field, with collisions.
- (2) Appleton-Hartree formula with field, no collisions.
- (3) Appleton-Hartree formula with collisions, no field.
- (4) Appleton-Hartree formula no field, no collisions.
- (5) Booker quartic with field, with collisions.
- (6) Booker quartic with field, no collisions.
- (7) Sen-Wyller formula with field.
- (8) Sen-Wyller formula, no field.

Each of these eight versions calculates  $n^2$ ,  $nn'$ ,  $n \partial n / \partial r$ ,  $n \partial n / \partial \theta$ ,  $n \partial n / \partial \varphi$ ,  $n \partial n / \partial V_r$ ,  $n \partial n / \partial V_\theta$ ,  $n \partial n / \partial V_\varphi$ ,  $n \partial n / \partial t$ , and the polarization, where  $n$  is the complex phase refractive index;  $n'$  is the complex group refractive index;  $r$ ,  $\theta$ , and  $\varphi$  are the spherical polar coordinates of a point on the ray path, and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions. The quantities

$$n' = n + f \frac{dn}{df} = n + \omega \frac{dn}{d\omega} , \quad (32)$$

and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions normalized so that

$$V_r^2 + V_\theta^2 + V_\varphi^2 = \text{Real} \{n^2\} . \quad (33)$$

The derivatives of the Hamiltonian in (22) are given in section 5. 5.

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- (1) Appleton-Hartree formula with field, with collisions.
- (2) Appleton-Hartree formula with field, no collisions.
- (3) Appleton-Hartree formula with collisions, no field.
- (4) Appleton-Hartree formula no field, no collisions.
- (5) Booker quartic with field, with collisions.
- (6) Booker quartic with field, no collisions.
- (7) Sen-Wyller formula with field.
- (8) Sen-Wyller formula, no field.

Each of these eight versions calculates  $n^2$ ,  $nn'$ ,  $n \partial n / \partial r$ ,  $n \partial n / \partial \theta$ ,  $n \partial n / \partial \varphi$ ,  $n \partial n / \partial V_r$ ,  $n \partial n / \partial V_\theta$ ,  $n \partial n / \partial V_\varphi$ ,  $n \partial n / \partial t$ , and the polarization, where  $n$  is the complex phase refractive index;  $n'$  is the complex group refractive index;  $r$ ,  $\theta$ , and  $\varphi$  are the spherical polar coordinates of a point on the ray path, and  $V_r$ ,  $V_\theta$ , and  $V_\varphi$  are the components of the wave normal direction in the  $r$ ,  $\theta$ , and  $\varphi$  directions. The quantities

$X$ ,  $\partial X/\partial r$ ,  $\partial X/\partial \theta$ ,  $\partial X/\partial \varphi$ , and  $\partial X/\partial t$  are supplied by one of the versions of subroutine ELECTX which defines the electron density model. The quantities  $Y$ ,  $\partial Y/\partial r$ ,  $\partial Y/\partial \theta$ ,  $\partial Y/\partial \varphi$ ,  $Y_r$ ,  $\partial Y_r/\partial r$ ,  $\partial Y_r/\partial \theta$ ,  $\partial Y_r/\partial \varphi$ ,  $Y_\theta$ ,  $\partial Y_\theta/\partial r$ ,  $\partial Y_\theta/\partial \theta$ ,  $\partial Y_\theta/\partial \varphi$ ,  $Y_\varphi$ ,  $\partial Y_\varphi/\partial r$ ,  $\partial Y_\varphi/\partial \theta$ , and  $\partial Y_\varphi/\partial \varphi$  are supplied by one of the versions of subroutine MAGY which defines the magnetic field model. The quantities  $Z$ ,  $\partial Z/\partial r$ ,  $\partial Z/\partial \theta$ , and  $\partial Z/\partial \varphi$  are supplied by one of the versions of subroutine COLFRZ which defines the collision frequency model.

In our formulation, we have tried to avoid using multivalued functions, such as the square root or  $\cos^{-1}$ , wherever possible. Only twice do we use the square root. One instance is the square root in the Appleton-Hartree formula, unavoidable without adding more differential equations to the system. The second instance is a square root used to calculate polarization. This latter use is unimportant because the polarization is not used in the ray tracing equations.

It is desirable to avoid multivalued functions because, unless extreme care is used, the value of such a function can change discontinuously from one point on the ray path to the next. A particularly troublesome case occurs at reflection for vertical incidence. At that point, the real part of  $n^2$  goes through zero, and  $n$  changes from approximately purely real to approximately purely imaginary. Since the numerical integration subroutine usually requires the evaluation of the differential equations not only on the ray path, but also at points near the ray path, it is necessary to be able to evaluate the differential equations above the reflection height, that is, in an evanescent region.

We have found that it is possible to regroup the variables in the equations to avoid this problem: we calculate the real part of  $n^2$  and its derivatives instead of the real part of  $n$  and its derivatives. And we calculate  $nn'$  instead of  $n'$ .



It was not easy to avoid using multivalued functions, however. Many of the usual parameters used to compute the refractive index require the use of multivalued functions in their calculation. Thus, we couldn't calculate

$$V = \sqrt{V_r^2 + V_\theta^2 + V_\phi^2}$$

nor  $\cos \psi$ , nor  $\sin \psi$ , where  $\psi$  is the angle between the wave normal direction and the earth's magnetic field. Thus, we also could not calculate

$$Y_L = Y \cos \psi$$

nor

$$Y_T = Y \sin \psi.$$

The most difficult part of avoiding the use of multivalued functions was in calculating the derivatives.

The following is a list of the equations calculated by the eight versions of subroutine RINDEX.

### 5.1 Appleton-Hartree Formula with Field, with Collisions

The square of the complex phase refractive index is given by

$$n^2 = 1 - 2X \frac{1 - iZ - X}{2(1 - iZ)(1 - iZ - X) - Y_T^2 \pm \sqrt{Y_T^4 + 4Y_L^2(1 - iZ - X)^2}}, \quad (34)$$

where

$$X = \frac{f^2 N}{f^2}, \quad (35)$$

$$Y = \frac{f_H}{f}, \quad (36)$$

$$Z = \frac{\nu}{2\pi f}, \quad (37)$$

$$Y_T = Y \sin \psi, \quad (38)$$

$$Y_L = Y \cos \psi, \quad (39)$$

$f_N$  is the plasma frequency,  $f_H$  is the electron gyrofrequency,  $\nu$  is the electron collision frequency,  $f$  is the wave frequency, and  $\psi$  is the angle between the wave normal direction and the earth's magnetic field.

The following equations parallel the formulas in this version of RINDEX.

$$V^2 = V_r^2 + V_\theta^2 + V_\phi^2 \quad (40)$$

$$V \cdot Y = V_r Y_r + V_\theta Y_\theta + V_\phi Y_\phi \quad (41)$$

$$\frac{Y_L}{V} = \frac{V \cdot Y}{V^2} \quad (42)$$

$$Y_L^2 = \frac{(V \cdot Y)^2}{V^2} \quad (43)$$

$$Y_T^2 = Y^2 - Y_L^2 \quad (44)$$

$$Y_T^4 = \left( Y_T^2 \right)^2 \quad (45)$$

$$U = 1 - iZ \quad (46)$$

$$RAD = \pm \sqrt{Y_T^4 + 4Y_L^2 (U-X)^2} \quad (47)$$

$$D = 2U(U-X) - Y_T^2 + RAD \quad (48)$$

$$n^2 = 1 - \frac{2X(U-X)}{D} \quad (49)$$

$$\frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} = \frac{2X(U-X) \left( -1 + \frac{Y_T^2 - 2(U-X)^2}{RAD} \right)}{D^2} \quad (50)$$

$$Y_T Y_L \frac{\partial \psi}{\partial r} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial r} - \left( V_r \frac{\partial Y_r}{\partial r} + V_\theta \frac{\partial Y_\theta}{\partial r} + V_\varphi \frac{\partial Y_\varphi}{\partial r} \right) \left( \frac{Y_L}{V} \right) \quad (51)$$

$$Y_T Y_L \frac{\partial \psi}{\partial \theta} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \theta} - \left( V_r \frac{\partial Y_r}{\partial \theta} + V_\theta \frac{\partial Y_\theta}{\partial \theta} + V_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) \left( \frac{Y_L}{V} \right) \quad (52)$$

$$Y_T Y_L \frac{\partial \psi}{\partial \varphi} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \varphi} - \left( V_r \frac{\partial Y_r}{\partial \varphi} + V_\theta \frac{\partial Y_\theta}{\partial \varphi} + V_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right) \left( \frac{Y_L}{V} \right) \quad (53)$$

$$n \frac{\partial n}{\partial X} = - \frac{\left( 2U(U-X)^2 - Y_T^2(U-2X) + \frac{Y_T^4(U-2X) + 4Y_L^2(U-X)^3}{\text{RAD}} \right)}{D^2} \quad (54)$$

$$n \frac{\partial n}{\partial Y} = \frac{2X(U-X)}{D^2 Y} \left( -Y_T^2 + \frac{Y_T^4 + 2Y_L^2(U-X)^2}{\text{RAD}} \right) \quad (55)$$

$$n \frac{\partial n}{\partial Z} = \frac{iX}{D^2} \left( -2(U-X)^2 - Y_T^2 + \frac{Y_T^4}{\text{RAD}} \right) \quad (56)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_L Y_T \frac{\partial \psi}{\partial r} \quad (57)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_T Y_L \frac{\partial \psi}{\partial \theta} \quad (58)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} + \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} Y_T Y_L \frac{\partial \psi}{\partial \varphi} \quad (59)$$

$$n \frac{\partial n}{\partial V_r} = \frac{n}{Y_T Y_L} \frac{\partial n}{\partial \psi} \left( \frac{V_r Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_r \right) \quad (60)$$

$$n \frac{\partial n}{\partial V_\theta} = \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \left( \frac{V_\theta Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_\theta \right) \quad (61)$$

$$n \frac{\partial n}{\partial V_\varphi} = \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \left( \frac{V_\varphi Y_L^2}{V^2} - \left( \frac{Y_L}{V} \right) Y_\varphi \right) \quad (62)$$

$$nn' = n^2 - \left( 2Xn \frac{\partial n}{\partial X} + Yn \frac{\partial n}{\partial Y} + Zn \frac{\partial n}{\partial Z} \right) \quad (63)$$

$$\text{Polarization} = \rho = -i \frac{(-Y_T^2 + \text{RAD}) \sqrt{V^2}}{2(U-X) V \cdot Y} \quad (64a)$$

(Budden, 1961, page 49).

$$\text{Longitudinal polarization} = \frac{iX\sqrt{Y_T^2}}{(U-X)(U+i \frac{V \cdot Y}{\sqrt{V^2}} \rho)} \quad (64b)$$

(Budden, 1961, page 54).

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (65)$$

## 5.2 Appleton-Hartree Formula with Field, no Collisions

The equations are the same as with field with collisions except for:

$$Z = \frac{\partial Z}{\partial r} = \frac{\partial Z}{\partial \theta} = \frac{\partial Z}{\partial \varphi} = 0, \quad (66)$$

$$U = 1. \quad (67)$$

### 5.3 Appleton-Hartree Formula no Field, With Collisions

$$U = 1 - iZ \quad (68)$$

$$n^2 = 1 - \frac{X}{U} \quad (69)$$

$$n \frac{\partial n}{\partial X} = -\frac{.5}{U} \quad (70)$$

$$n \frac{\partial n}{\partial Z} = -\frac{.5iX}{U^2} \quad (71)$$

$$nn' = n^2 - \left( 2Xn \frac{\partial n}{\partial X} + Zn \frac{\partial n}{\partial Z} \right) \quad (72)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} \quad (73)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} \quad (74)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} \quad (75)$$

$$n \frac{\partial n}{\partial V_r} = 0 \quad (76)$$

$$n \frac{\partial n}{\partial V_\theta} = 0 \quad (77)$$

$$n \frac{\partial n}{\partial V_\varphi} = 0 \quad (78)$$

$$\text{Polarization} = i \quad (79a)$$

$$\text{Longitudinal polarization} = 0 \quad (79b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (80)$$

#### 5.4 Appleton-Hartree Formula no Field, no Collisions

$$n^2 = 1 - X \quad (81)$$

$$n \frac{\partial n}{\partial X} = -.5 \quad (82)$$

$$n \frac{\partial n}{\partial r} = \left( n \frac{\partial n}{\partial X} \right) \left( \frac{\partial X}{\partial r} \right) \quad (83)$$

$$n \frac{\partial n}{\partial \theta} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \theta} \quad (84)$$

$$n \frac{\partial n}{\partial \varphi} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \varphi} \quad (85)$$

$$n \frac{\partial n}{\partial V_r} = 0 \quad (86)$$

$$n \frac{\partial n}{\partial V_\theta} = 0 \quad (87)$$

$$n \frac{\partial n}{\partial V_\varphi} = 0 \quad (88)$$

$$nn' = 1 \quad (89)$$

$$\text{Polarization} = i \quad (90a)$$

$$\text{Longitudinal polarization} = 0 \quad (90b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (91)$$

#### 5.5 Booker Quartic with Field, with Collisions

The form of the dispersion relation used for the Hamiltonian in this version of subroutine RINDEX is the quadratic equation whose solution is the Appleton-Hartree formula. This Hamiltonian, given by

(22), is also a special case of the Booker quartic for  $S = 0$  and  $C = 1$  (Budden, 1961). This version uses the Hamiltonian in (22) only when the electron density is large enough that  $X$  is greater than or equal to  $1/10$ . For  $X$  less than  $1/10$ , the Hamiltonian in (21) is used.

Below are the equations for the derivatives of the Hamiltonian in (22). The equations for the derivatives of the Hamiltonian in (21) are the same as those in section 4.1.

$$k^2 = k_r^2 + k_\theta^2 + k_\varphi^2 , \quad (92)$$

$$k \cdot Y = k_r Y_r + k_\theta Y_\theta + k_\varphi Y_\varphi , \quad (93)$$

$$U = 1 - iZ , \quad (94)$$

$$A = (U - X) U^2 - UY^2 , \quad (95)$$

$$B = -2U(U - X)^2 + Y^2(2U - X) , \quad (96)$$

$$\alpha = A c^4 k^4 + X(k \cdot Y)^2 c^4 k^2 , \quad (97)$$

$$\beta = B c^2 k^2 \omega^2 - X(k \cdot Y)^2 c^2 \omega^2 , \quad (98)$$

$$\gamma = ((U - X)^2 - Y^2)(U - X) \omega^4 , \quad (99)$$

$$H = \alpha + \beta + \gamma , \quad (100)$$

$$\begin{aligned} \frac{\partial H}{\partial X} = & -U^2 c^4 k^4 + (k \cdot Y)^2 c^4 k^2 + (4U(U - X) - Y^2) c^2 k^2 \omega^2 + \\ & -(k \cdot Y)^2 c^2 \omega^2 + (-3(U - X)^2 + Y^2) \omega^4 , \end{aligned} \quad (101)$$

$$\frac{\partial H}{\partial(Y^2)} = -U c^4 k^4 + (2U - X) c^2 k^2 \omega^2 - (U - X) \omega^4 , \quad (102)$$

$$\frac{\partial H}{\partial((k \cdot Y)^2)} = X c^2 (c^2 k^2 - \omega^2) , \quad (103)$$

$$\begin{aligned} \frac{\partial H}{\partial U} &= (2 U(U - X) + U^2 - Y^2) c^4 k^4 + \\ &+ 2(Y^2 - (U - X)^2 - 2U(U - X)) c^2 k^2 \omega^2 + \\ &+ (3(U - X)^2 - Y^2) \omega^4 , \end{aligned} \quad (104)$$

$$\frac{\partial H}{\partial Z} = -i \frac{\partial H}{\partial U} , \quad (105)$$

$$\frac{\partial H}{\partial(k^2)} = 2 A c^4 k^2 + X(k \cdot Y)^2 c^4 + B c^2 \omega^2 , \quad (106)$$

$$\frac{\partial H}{\partial t} = \frac{\partial H}{\partial X} \frac{\partial X}{\partial t} , \quad (107)$$

$$\begin{aligned} \frac{\partial H}{\partial r} &= \frac{\partial H}{\partial X} \frac{\partial X}{\partial r} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial r} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial r} + \\ &+ 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial r} + k_\theta \frac{\partial Y_\theta}{\partial r} + k_\varphi \frac{\partial Y_\varphi}{\partial r} \right) , \end{aligned} \quad (108)$$

$$\begin{aligned} \frac{\partial H}{\partial \theta} &= \frac{\partial H}{\partial X} \frac{\partial X}{\partial \theta} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial \theta} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial \theta} + \\ &+ 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial \theta} + k_\theta \frac{\partial Y_\theta}{\partial \theta} + k_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) , \end{aligned} \quad (109)$$



$$\begin{aligned} \frac{\partial H}{\partial \varphi} = & \frac{\partial H}{\partial X} \frac{\partial X}{\partial \varphi} + 2 \frac{\partial H}{\partial(Y^2)} Y \frac{\partial Y}{\partial \varphi} + \frac{\partial H}{\partial Z} \frac{\partial Z}{\partial \varphi} + \\ & + 2 \frac{\partial H}{\partial((k \cdot Y)^2)} (k \cdot Y) \left( k_r \frac{\partial Y_r}{\partial \varphi} + k_\theta \frac{\partial Y_\theta}{\partial \varphi} + k_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right), \end{aligned} \quad (110)$$

$$\begin{aligned} \frac{\partial H}{\partial \omega} = & (2\beta + 4\gamma)/\omega - 2 \frac{\partial H}{\partial X} \frac{X}{\omega} - 2 \frac{\partial H}{\partial(Y^2)} \frac{Y^2}{\omega} + \\ & - 2 \frac{\partial H}{\partial((k \cdot Y)^2)} \frac{(k \cdot Y)^2}{\omega} - \frac{\partial H}{\partial Z} \frac{Z}{\omega}, \end{aligned} \quad (111)$$

$$\frac{\partial H}{\partial k_r} = 2 \frac{\partial H}{\partial(k^2)} k_r + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_r, \quad (112)$$

$$\frac{\partial H}{\partial k_\theta} = 2 \frac{\partial H}{\partial(k^2)} k_\theta + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_\theta, \quad (113)$$

$$\frac{\partial H}{\partial k_\varphi} = 2 \frac{\partial H}{\partial(k^2)} k_\varphi + 2(k \cdot Y) \frac{\partial H}{\partial((k \cdot Y)^2)} Y_\varphi, \quad (114)$$

$$k^2(\text{calculated}) = k^2 \frac{-\beta \pm \sqrt{\beta^2 - 4\alpha\gamma}}{2\alpha}, \quad (115)$$

$$k \cdot \frac{\partial H}{\partial k} = k_r \frac{\partial H}{\partial k_r} + k_\theta \frac{\partial H}{\partial k_\theta} + k_\varphi \frac{\partial H}{\partial k_\varphi} = 4\alpha + 2\beta, \quad (116)$$

$$\text{Polarization} = i \frac{\sqrt{k^2}}{k \cdot Y} \left( U + \frac{X\omega^2}{c^2 k^2 \text{calculated} - \omega^2} \right), \quad (117)$$

$$\begin{aligned} \text{Longitudinal} \\ \text{Polarization} = & i \frac{\sqrt{Y^2 - \frac{(k \cdot Y)^2}{k^2}}}{U - X} \left( 1 - \frac{c^2 k^2 \text{calculated}}{\omega^2} \right). \end{aligned} \quad (118)$$

### 5.6 Booker Quartic with Field, no Collisions

All the equations here are the same as for the Booker quartic version with collisions (section 5.5) except for:

$$Z = \frac{\partial Z}{\partial r} = \frac{\partial Z}{\partial \theta} = \frac{\partial Z}{\partial \varphi} = 0 , \quad (119)$$

$$U = 1. \quad (120)$$

All the variables except polarization are real; polarization is pure imaginary.

### 5.7 Sen-Wyller Formula with Field

It is possible to write the generalized Appleton-Hartree formula of Sen and Wyller (1960) in the following form:

$$n^2 = 1 - \frac{2X(U-X) + 2 AUX \sin^2 \psi}{2U(U-X)(1+A) + 2 AUX \sin^2 \psi - U((1-BC)U + A(U+X)) \sin^2 \psi + RAD} , \quad (121)$$

where

$$RAD = \pm \sqrt{U^2((1-BC)U + A(U+X))^2 \sin^4 \psi + U^2(U-X)^2(C-B)^2 \cos^2 \psi} , \quad (122)$$

$$A = \frac{C+B}{2} - 1 , \quad (123)$$

$$B = \frac{F\left(\frac{1}{Z}\right)}{F\left(\frac{1-Y}{Z}\right)} , \quad (124)$$

$$C = \frac{F\left(\frac{1}{Z}\right)}{F\left(\frac{1+Y}{Z}\right)}, \quad (125)$$

$$U = \frac{Z}{F\left(\frac{1}{Z}\right)} \quad (126)$$

$$F(w) = w C_{3/2}(w) + i \frac{5}{2} C_{5/2}(w) = \frac{1}{\frac{3}{2}!} \int_0^{\infty} \frac{t^{3/2} e^{-t}}{w-it} dt, \quad (127)$$

$$X = \frac{f_N^2}{f^2}, \quad (128)$$

$$Y = \frac{f_H}{f}, \quad (129)$$

$$Z = \frac{\nu_m}{2\pi f}, \quad (130)$$

$f_N$  is the plasma frequency,  $f_H$  is the electron gyrofrequency,  $\nu_m$  is the mean electron collision frequency,  $f$  is the wave frequency, and  $\psi$  is the angle between the wave normal direction and the earth's magnetic field. (Note that if we would use

$$F(w) = \frac{1}{w-i}, \quad (131)$$

then (121) would reduce to the usual Appleton-Hartree formula.)

This version of RINDEX calls subroutine FSW. The first argument in the calling sequence is the argument  $w$  of  $F(w)$ . The second argument in the calling sequence is the value of the function  $F(w)$  calculated by FSW. The third argument in the calling sequence is the derivative of the function,  $F'(w)$ .

The following equations parallel the formulas in this version of RINDEX.

$$\alpha = F\left(\frac{1}{Z}\right) \quad (132)$$

$$\alpha' = F'\left(\frac{1}{Z}\right) \quad (133)$$

$$\beta = F\left(\frac{1-Y}{Z}\right) \quad (134)$$

$$\beta' = F'\left(\frac{1-Y}{Z}\right) \quad (135)$$

$$\gamma = F\left(\frac{1+Y}{Z}\right) \quad (136)$$

$$\gamma' = F'\left(\frac{1+Y}{Z}\right) \quad (137)$$

$$U = \frac{Z}{\alpha} \quad (138)$$

$$\frac{\partial U}{\partial Z} = \left(1 + \frac{\alpha'}{\alpha Z}\right) / \alpha \quad (139)$$

$$B = \frac{\alpha}{\beta} \quad (140)$$

$$\frac{\partial B}{\partial Y} = \frac{B}{Z} \frac{\beta'}{\beta} \quad (141)$$

$$\frac{\partial B}{\partial Z} = -\frac{B}{Z^2} \left(\frac{\alpha'}{\alpha} - (1-Y) \frac{\beta'}{\beta}\right) \quad (142)$$

$$C = \frac{\alpha}{\gamma} \quad (143)$$

$$\frac{\partial C}{\partial Y} = -\frac{C}{Z} \frac{\gamma'}{\gamma} \quad (144)$$

$$\frac{\partial C}{\partial Z} = -\frac{C}{Z^2} \left( \frac{\alpha'}{\alpha} - (1+Y) \frac{\gamma'}{\gamma} \right) \quad (145)$$

$$A = .5(B+C) - 1 \quad (146)$$

$$\frac{\partial A}{\partial Y} = .5 \left( \frac{\partial B}{\partial Y} + \frac{\partial C}{\partial Y} \right) \quad (147)$$

$$\frac{\partial A}{\partial Z} = .5 \left( \frac{\partial B}{\partial Z} + \frac{\partial C}{\partial Z} \right) \quad (148)$$

$$V^2 = V_r^2 + V_\theta^2 + V_\varphi^2 \quad (149)$$

$$V \cdot Y = V_r Y_r + V_\theta Y_\theta + V_\varphi Y_\varphi \quad (150)$$

$$Y_L^2 = \frac{(V \cdot Y)^2}{V^2} \quad (151)$$

$$Y_T^2 = Y^2 - Y_L^2 \quad (152)$$

$$\sin^2 \psi = \frac{Y_T^2}{Y^2} \quad (153)$$

$$\cos^2 \psi = \frac{Y_L^2}{Y^2} \quad (154)$$

$$T_1 = [(1-BC) U^2 + A U(U+X)] \sin^2 \psi \quad (155)$$

$$\frac{\partial T_1}{\partial X} = +AU \sin^2 \psi \quad (156)$$

$$\frac{\partial T_1}{\partial Y} = \left( U(U+X) \frac{\partial A}{\partial Y} - U^2 \left( B \frac{\partial C}{\partial Y} + C \frac{\partial B}{\partial Y} \right) \right) \sin^2 \psi \quad (157)$$

$$\begin{aligned} \frac{\partial T_1}{\partial Z} = & \left( 2U \frac{\partial U}{\partial Z} (1-BC+A) + A X \frac{\partial U}{\partial Z} - U^2 \left( B \frac{\partial C}{\partial Z} + C \frac{\partial B}{\partial Z} \right) \right. \\ & \left. + U(U+X) \frac{\partial A}{\partial Z} \right) \sin^2 \psi \end{aligned} \quad (158)$$

$$\frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} = \frac{2T_1}{Y_T^2} \quad (159)$$

$$T_2 = U^2 (C-B)^2 (U-X)^2 \cos^2 \psi \quad (160)$$

$$\frac{\partial T_2}{\partial X} = -2(U-X) U^2 (C-B)^2 \cos^2 \psi \quad (161)$$

$$\frac{\partial T_2}{\partial Y} = 2U^2 (U-X)^2 \cos^2 \psi (C-B) \left( \frac{\partial C}{\partial Y} - \frac{\partial B}{\partial Y} \right) \quad (162)$$

$$\frac{\partial T_2}{\partial Z} = 2U^2 (U-X)^2 (C-B) \left( \frac{\partial C}{\partial Z} - \frac{\partial B}{\partial Z} \right) \cos^2 \psi + 2T_2 \left( \frac{1}{U} + \frac{1}{U-X} \frac{\partial U}{\partial Z} \right) \quad (163)$$

$$\frac{1}{Y_L Y_T} \frac{\partial T_2}{\partial \psi} = -\frac{2T_2}{Y_L^2} \quad (164)$$

$$RAD = \pm \sqrt{T_1^2 + T_2} \quad (165)$$

$$\frac{\partial \text{RAD}}{\partial X} = \frac{T_1 \frac{\partial T_1}{\partial X} + \frac{1}{2} \frac{\partial T_2}{\partial X}}{\text{RAD}} \quad (166)$$

$$\frac{\partial \text{RAD}}{\partial Y} = \frac{T_1 \frac{\partial T_1}{\partial Y} + \frac{1}{2} \frac{\partial T_2}{\partial Y}}{\text{RAD}} \quad (167)$$

$$\frac{\partial \text{RAD}}{\partial Z} = \frac{T_1 \frac{\partial T_1}{\partial Z} + \frac{1}{2} \frac{\partial T_2}{\partial Z}}{\text{RAD}} \quad (168)$$

$$\frac{1}{Y_L Y_T} \frac{\partial \text{RAD}}{\partial \psi} = \frac{T_1 \left( \frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} \right) + \frac{1}{2} \left( \frac{1}{Y_L Y_T} \frac{\partial T_2}{\partial \psi} \right)}{\text{RAD}} \quad (169)$$

$$D = 2U(U-X) (1 + A) - T_1 + \text{RAD} + 2AUX \sin^2 \psi \quad (170)$$

$$\frac{\partial D}{\partial X} = -2U - \frac{\partial T_1}{\partial X} + \frac{\partial \text{RAD}}{\partial X} + 2AU \sin^2 \psi \quad (171)$$

$$\frac{\partial D}{\partial Y} = 2U(U-X) \frac{\partial A}{\partial Y} - \frac{\partial T_1}{\partial Y} + \frac{\partial \text{RAD}}{\partial Y} + 2U \sin^2 \psi \frac{\partial A}{\partial Y} \quad (172)$$

$$\begin{aligned} \frac{\partial D}{\partial Z} = & 2(1+A) \frac{\partial U}{\partial Z} (2U-X) + 2U(U-X) \frac{\partial A}{\partial Z} - \frac{\partial T_1}{\partial Z} + \frac{\partial \text{RAD}}{\partial Z} + \\ & + 2AX \sin^2 \psi \frac{\partial U}{\partial Z} + 2UX \sin^2 \psi \frac{\partial A}{\partial Z} \end{aligned} \quad (173)$$

$$\frac{1}{Y_L Y_T} \frac{\partial D}{\partial \psi} = - \left( \frac{1}{Y_L Y_T} \frac{\partial T_1}{\partial \psi} \right) + \left( \frac{1}{Y_L Y_T} \frac{\partial \text{RAD}}{\partial \psi} \right) + 2AUX / Y^2 \quad (174)$$

$$n^2 - 1 = \frac{-2X}{D} \left( (U-X) + UA \sin^2 \psi \right) \quad (175)$$

$$n^2 = 1 + (n^2 - 1) \quad (176)$$

$$n \frac{\partial n}{\partial X} = \frac{1}{2} (n^2 - 1) \left( \frac{1}{X} - \frac{1}{D} \frac{\partial D}{\partial X} \right) + \frac{X}{D} \quad (177)$$

$$n \frac{\partial n}{\partial Y} = - \frac{XU \sin^2 \psi}{D} \frac{\partial A}{\partial Y} - \frac{(n^2 - 1)}{2D} \frac{\partial D}{\partial Y} \quad (178)$$

$$n \frac{\partial n}{\partial Z} = - \frac{X}{D} (1+A \sin^2 \psi) \frac{\partial U}{\partial Z} - \frac{XU}{D} \sin^2 \psi \frac{\partial A}{\partial Z} - \frac{(n^2 - 1)}{2D} \frac{\partial D}{\partial Z} \quad (179)$$

$$\frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} = - \frac{XUA}{Y^2 D} - \frac{(n^2 - 1)}{2D} \left( \frac{1}{Y_L Y_T} \frac{\partial D}{\partial \psi} \right) \quad (180)$$

$$Y_L Y_T \frac{\partial \psi}{\partial r} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial r} - \left( V_r \frac{\partial Y_r}{\partial r} + V_\theta \frac{\partial Y_\theta}{\partial r} + V_\varphi \frac{\partial Y_\varphi}{\partial r} \right) \frac{V \cdot Y}{V^2} \quad (181)$$

$$Y_L Y_T \frac{\partial \psi}{\partial \theta} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \theta} - \left( V_r \frac{\partial Y_r}{\partial \theta} + V_\theta \frac{\partial Y_\theta}{\partial \theta} + V_\varphi \frac{\partial Y_\varphi}{\partial \theta} \right) \frac{V \cdot Y}{V^2} \quad (182)$$

$$Y_L Y_T \frac{\partial \psi}{\partial \varphi} = \frac{Y_L^2}{Y} \frac{\partial Y}{\partial \varphi} - \left( V_r \frac{\partial Y_r}{\partial \varphi} + V_\theta \frac{\partial Y_\theta}{\partial \varphi} + V_\varphi \frac{\partial Y_\varphi}{\partial \varphi} \right) \frac{V \cdot Y}{V^2} \quad (183)$$

$$n \frac{\partial n}{\partial r} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial r} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial r} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial r} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial r} \right) \quad (184)$$

$$n \frac{\partial n}{\partial \theta} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \theta} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \theta} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \theta} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial \theta} \right) \quad (185)$$

$$n \frac{\partial n}{\partial \varphi} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial \varphi} + n \frac{\partial n}{\partial Y} \frac{\partial Y}{\partial \varphi} + n \frac{\partial n}{\partial Z} \frac{\partial Z}{\partial \varphi} + \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( Y_L Y_T \frac{\partial \psi}{\partial \varphi} \right) \quad (186)$$



$$n \frac{\partial n}{\partial V_r} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_r Y_L^2 - (V \cdot Y) Y_r \right) / V^2 \quad (187)$$

$$n \frac{\partial n}{\partial V_\theta} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_\theta Y_L^2 - (V \cdot Y) Y_\theta \right) / V^2 \quad (188)$$

$$n \frac{\partial n}{\partial V_\varphi} = \left( \frac{n}{Y_L Y_T} \frac{\partial n}{\partial \psi} \right) \left( V_\varphi Y_L^2 - (V \cdot Y) Y_\varphi \right) / V^2 \quad (189)$$

$$nn' = n^2 - \left( 2X \left( n \frac{\partial n}{\partial X} \right) + Y \left( n \frac{\partial n}{\partial Y} \right) + Z \left( n \frac{\partial n}{\partial Z} \right) \right) \quad (190)$$

$$\text{polarization} = \rho = \frac{i(T_1 - \text{RAD}) Y \sqrt{V^2}}{U(U-X)(C-B)(V \cdot Y)} \quad (191a)$$

$$\text{longitudinal polarization} = \frac{X(.5i(C-B)\rho + A \cos \psi) \sqrt{\sin^2 \psi}}{\rho((U-X)(1+.5i(C-B)\rho \cos \psi) + A(U-X \cos^2 \psi))} \quad (191b)$$

where

$$\cos \psi = \frac{V \cdot Y}{Y \sqrt{V^2}} \quad (192)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (193)$$

### 5.8 Sen-Wyller Formula no Field

This subroutine uses the function G which is related to the function F (defined in the previous section) as follows:

$$G(w) = wF(w) \quad (194)$$

The following equations parallel the formulas used in this version of RINDEX.

$$F_1 = F \left( \frac{1}{Z} \right) \quad (195)$$

$$F_1' = F' \left( \frac{1}{Z} \right) \quad (196)$$

$$G_1 = G \left( \frac{1}{Z} \right) \quad (197)$$

$$G_1' = G' \left( \frac{1}{Z} \right) \quad (198)$$

$$n^2 = 1 - X G_1 \quad (199)$$

$$n \frac{\partial n}{\partial X} = -\frac{1}{2} G_1 \quad (200)$$

$$n \frac{\partial n}{\partial Z} = \frac{X G_1'}{2 Z^2} \quad (201)$$

$$n \frac{\partial n}{\partial r} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial r} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial r} \quad (202)$$

$$n \frac{\partial n}{\partial \theta} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \theta} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial \theta} \quad (203)$$

$$n \frac{\partial n}{\partial \varphi} = \left( n \frac{\partial n}{\partial X} \right) \frac{\partial X}{\partial \varphi} + \left( n \frac{\partial n}{\partial Z} \right) \frac{\partial Z}{\partial \varphi} \quad (204)$$

$$nn' = n^2 - \left( 2 X \left( n \frac{\partial n}{\partial X} \right) + Z \left( n \frac{\partial n}{\partial Z} \right) \right) \quad (205)$$

$$\text{polarization} = i \quad (206a)$$

$$\text{longitudinal polarization} = 0 \quad (206b)$$

$$n \frac{\partial n}{\partial t} = n \frac{\partial n}{\partial X} \frac{\partial X}{\partial t} \quad (207)$$

## 6. IONOSPHERIC MODELS

When using the program, one must specify ionospheric models which define electron density, collision frequency (if the effect of collisions is being considered), and the earth's magnetic field (if its effects are being taken into account) as a function of position in space. Each of these three characteristics of the ionosphere is defined by a separate subroutine.

Appendices 3, 4, 5, and 6 contain descriptions, input parameter forms and listings of ionospheric models that now exist. These ionospheric models are not likely to cover the needs of everyone who wants to use the program. Anticipating this when we wrote the program, we made it possible to add models easily. The user may make up his own ionospheric models by simply writing subroutines to define electron density, collision frequency, and the earth's magnetic field (and their gradients) as a function of position in space in spherical polar coordinates, following the form of the subroutines in appendices 3, 4, 5, and 6.

Appendix 3 contains electron density models; appendix 4 contains models of irregularities which may be applied as perturbations to any of the electron density models; appendix 5 contains models of the earth's magnetic field; and appendix 6 contains collision frequency models.

Having several versions of the subroutines for refractive index, electron density, collision frequency, and the earth's magnetic field gives the user not only a wide choice among ionospheric models, but also a variety of compromises between cost and an accurate description of the ionosphere, while still keeping the program simple.

## 7. FINDING THE RAY PATHS THAT CONNECT A TRANSMITTER AND RECEIVER

The reason for using a ray tracing program is to find all important ray paths that connect a given transmitter and receiver (either or both of which may be on a satellite) on a particular frequency and such properties of these ray paths as group time delay, phase time delay, and absorption of the wave. "All important" ray paths include those that reflect from the various ionospheric layers (including multiple reflections) and that propagate off the great circle path.

Since basically all that a ray tracing program can do is to calculate the path of a ray when given the transmitter location, frequency, and direction of transmission, it cannot directly calculate those ray paths that arrive at a specified receiver. The problem is to know, before tracing the ray, in which directions to transmit the ray so that it will arrive at the receiver. Since there are no general solutions to this problem, the user of a ray tracing program must rely on some sort of trial and error technique to find those ray paths that connect the transmitter and receiver. This involves varying the direction of transmission until a ray is found that reaches the receiver. If a ray tracing program does this automatically, we say that it has a homing feature. This program does not have such a feature. To find all the paths connecting the transmitter with the receiver requires a very elaborate homing routine because "homing in" on a receiver takes more judgment and common sense than speed in performing massive calculations. Therefore, the person using the program is more fitted to this task than is the computer program itself.

As an aid, however, the program allows the user to specify the receiver height, the number of hops, and a range of azimuth and elevation angles-of-transmission that he thinks will include those rays that

will arrive at the receiver. The program then calculates a ray path for each of the azimuth and elevation angles-of-transmission specified in the range. Usually only in the case of ionospheres with large horizontal gradients will the azimuth angle-of-transmission have to be varied. The program will calculate each ray path far enough to intersect or make a closest approach to the receiver height for the requested number of hops. The user can then interpolate between those rays which surround the receiver.

We define the point of "closest approach" as the point on the ray path where the wave normal direction is horizontal. It approximates an apogee if the receiver is above the apogee height and it approximates a perigee if the receiver is below the perigee height. The approximation is good for oblique propagation. When the earth's magnetic field is neglected, a point of "closest approach" is exactly an apogee or perigee.

We count one hop every time the ray crosses the receiver height. If the receiver is on the ground, a ground reflection counts as one hop for the downcoming ray before the ground reflection, and another hop for the upgoing ray after the ground reflection. We count two hops every time the ray passes through a point of "closest approach" to the receiver height. This procedure helps make rays that have the same hop number have a ground range that is a continuous function of the direction of transmission.

## 8. OUTPUT

### 8.1 Printout

Periodically and at selected points during a ray trace, the program will print information giving the position of the current ray path point, the direction of the wave normal, and the cumulative values of quantities being integrated along the ray path such as group path, phase path, absorption, and Doppler shift. Appendix 8c contains a sample of the printout.

## 8.2 Punched Cards

The program will punch a card at the beginning of each ray, a card at each ground reflection, a card at each crossing of the receiver height, two cards for each closest approach to the receiver height, and a card at the end of each ray to summarize the main results of the ray path calculations. These cards are explained in figures 1 and 2.

These cards are very useful as input data to other computer programs and for plotting the results of the ray tracing. In fact, these cards represent the most useful form of output for production ray path calculations. This method, called the rayset information-storage technique, was developed by Dr. T. A. Croft (Croft and Gregory, 1963) of Stanford University.

## 8.3 Plots of the Ray Path

A plot of the actual ray path, especially for very irregular ionospheres, can be helpful in understanding what sometimes seems like strange results in light of the input data. Thus, the program has an option for plotting, providing, of course, that the user has a plotter and plotting subroutines such as those described in appendix 7. The program can plot the projection of the ray path on any vertical plane or on the ground. The input parameter forms for plots of the ray path (appendix 1i) give more details. Appendix 8e contains sample plots of the raypath.

## 9. DECK SET UP

The versatility gained by having several versions of some of the subroutines is somewhat offset because the user must learn the deck set up in order to make necessary substitutions. Figure 3 shows the deck setup, including the subroutines that make up the main deck and those which are frequently exchanged with alternate versions. The order of the subroutines is unimportant.



1 2 3 4	5 6 7 8 9	10 11 12 13	14 15 16 17 18 19	20 21 22 23 24	25 26 27 28 29 30 31	32 33 34 35 36 37 38 39 40	41 42 43 44 45 46 47 48 49 50	51 52 53 54 55 56 57 58 59 60	61 62 63 64 65 66 67 68 69 70	71 72 73 74 75 76 77 78 79 80			
1581469	14911561	-29	0	1514389	4513	-1610	11	0	-0	1721M			
Col. 1-9	Col. 10-18	Col. 19-24	Col. 25-30	Col. 31-36	Col. 37-44	Col. 45-50	Col. 51-56	Col. 57-62	Col. 63-68	Col. 69-73	Col. 74-78	Col. 79	Col. 80
Height of ray, km	Ground range between transmitter and receiver, km	At transmitter	At ray point	Elevation angle of arrival, deg	Straight line distance between transmitter and ray point, km	Group path minus straight line distance between transmitter and ray point, km	Phase path minus straight line distance between transmitter and ray point, km	Absorption, db	Doppler shift, Hz	Real part of wave polarization	Imaginary part of wave polarization	Hop number	Alphabetic character indicating type of rayset*
00000 1234 1111 2222 1234 3333 4444 1234 5555 6666 1234 7777 8888 1234 9999	00000 7891011 1111 2222 7891011 33333 4444 7891011 55555 6666 7891011 77777 88888 7891011 99999	00100 32425 1111 22 124252 333 444 11192 555 666 11192 777 888 11192 999	100 10031 11 22 20313 33 44 30313 55 66 30313 77 88 30313 99	100 526273 111 222 526273 133 444 526273 155 666 526273 177 188 526273 199	01 6575 111 222 15751 133 144 6575 155 166 1661 77 88 157 99999999	0 4 1 2 0 - 4 - 5 - 6 - 7 - 8 - 9	0 76861 111 222 76869 333 144 76869 555 666 162638 777 888 162638 999	01 6575 111 222 15751 133 144 6575 155 166 1661 77 88 157 99999999	000 162631 1 222 162638 333 144 162631 555 666 162638 777 888 162638 999	0 1374 11 22 374 33 14 374 55 66 374 77 88 374 99	0 1374 11 22 374 33 14 374 55 66 374 77 88 374 99	888 1878 999	M

▲ indicates implied decimal

- \* G ray ground reflected. The height punched out is the apogee height since the last ground reflect.
- M ray made a closest approach to the receiver height
- P ray penetrated
- R ray at the receiver height. The height punched out is the height of the ray farthest from receiver height since last crossing of receiver height
- S program reached maximum number of steps

Figure 2. Sample minimum distance rayset.



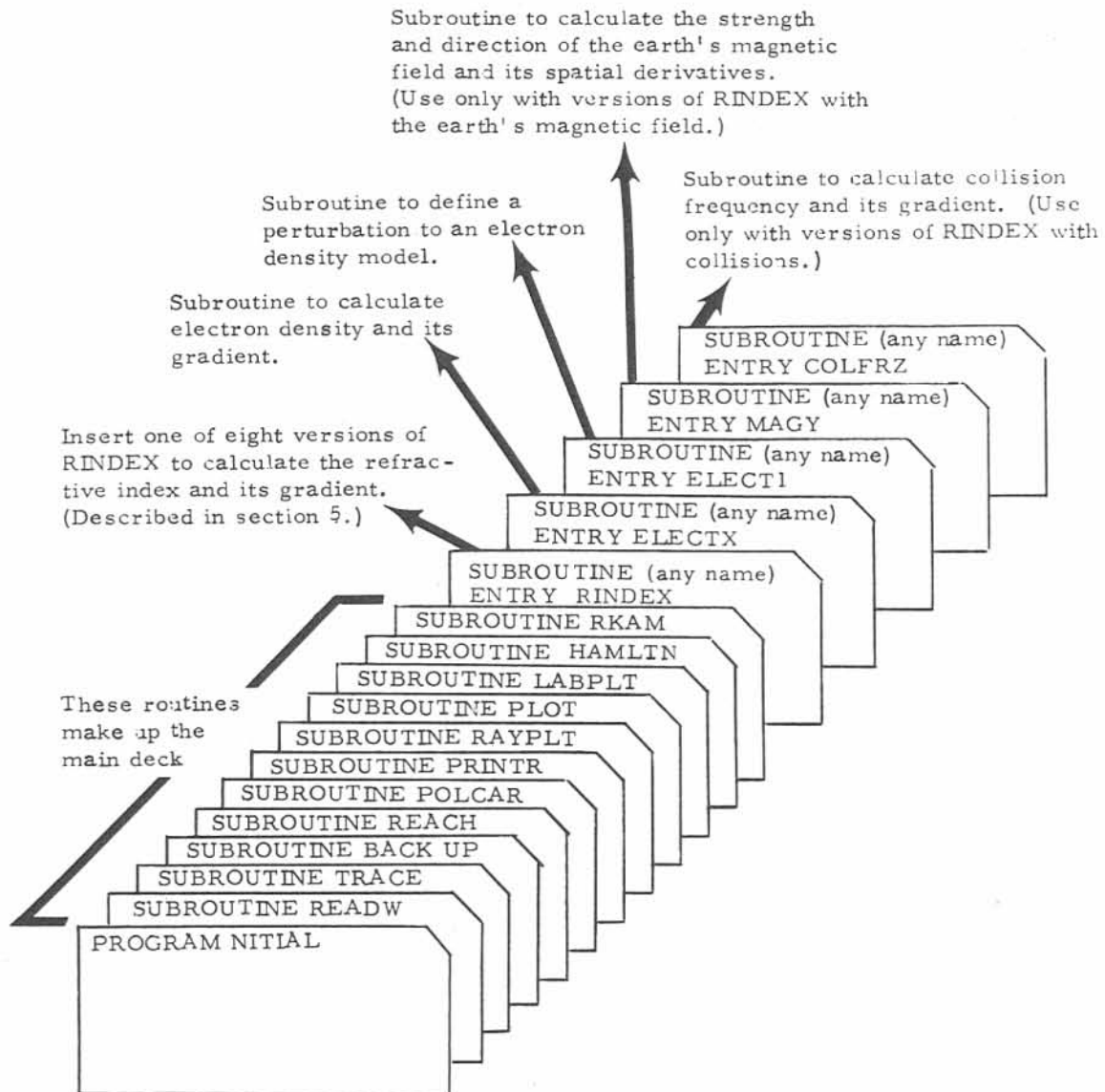


Figure 3. Program deck set-up

## 10. INPUT

The input data for a ray tracing program divide themselves naturally into two groups:

First, data that control the type of ray trace requested, such as the transmitter location and frequency, plus parameters describing analytic models of the ionosphere. Since there are few of these, efficiency in packing such data can be exchanged for versatility and ease of data handling. Therefore, by putting only one piece of data on each card, we gain the conveniences of reading in these data in any order and of having the program read in only those data that are different from those of the previous case. A number in the first three columns of each card identifies the data being read in. Table 2 defines the identifying numbers that are subscripts for a linear array,  $W$ . The last 56 columns of the card are available for comments.

We have also provided a method for conversion of units for input. The computer program needs angles in radians, whereas people usually like to use angles in degrees. The program is set up for angles in radians, but putting a "1" in column 18 allows the user to enter the angle in degrees and have the program make the conversion. A "1" in column 19 allows the user to enter central earth angles as the great circle distance along the ground in kilometers. (The program will calculate the latitude of a transmitter which is 500 km north of the equator, for instance.) The program expects distances in kilometers. A "1" in column 20 indicates a distance in nautical miles, and a "1" in column 21 indicates a distance in feet.

Appendix 8b contains a sample of how the cards are to be punched. If two or more cards have the same identifying number, the last one dominates. A card with the first three columns blank indicates the end of this type of data cards.

Table 2. Description of the Input Data for the W Array

W1	= 1. for ordinary ray = -1. for extraordinary ray
W2*	Radius of the earth in km
W3	Height of transmitter above the earth in km
W4	North geographic latitude of the transmitter
W5	East geographic longitude of the transmitter
W7	Initial frequency in MHz
W8	Final frequency in MHz
W9	Step in frequency in MHz (zero for a fixed frequency)
W11	Initial azimuth angle of transmission
W12	Final azimuth angle of transmission
W13	Step in azimuth angle of transmission (zero for a fixed azimuth)
W15	Initial elevation angle of transmission
W16	Final elevation angle of transmission
W17	Step in elevation angle of transmission (zero for a fixed elevation)
W20	Receiver height above the earth in km
W21	Nonzero to skip to the next frequency after the ray has penetrated the ionosphere
W22*	Maximum number of hops
W23*	Maximum number of steps per hop
W24*	North geographic latitude of the north geomagnetic pole
W25*	East geographic longitude of north geomagnetic pole
W41*	=1. for Runge-Kutta integration =2. for Adams-Moulton integration without error checking =3. for Adams-Moulton integration with relative error check =4. for Adams-Moulton integration with absolute error check
W42*	Maximum allowable single step error
W43*	Ratio of maximum single step error to minimum single step error
W44*	Initial integration step size in km (step in group path)
W45*	Maximum step length in km
W46*	Minimum step length in km
W47*	Factor by which to increase or decrease step length
W57	=1. to integrate, =2. to integrate and print phase path
W58	=1. to integrate, =2. to integrate and print absorption
W59	=1. to integrate, =2. to integrate and print doppler shift
W60	=1. to integrate, =2. to integrate and print path length
W71	Number of steps between periodic printout
W72	Nonzero to punch raysets on cards
W81	=0. to not plot ray path =1. to plot projection of ray path on a vertical plane =2. to plot projection of ray path on the ground
W82-88	Parameters used when plotting
W100-149	Parameters for analytic electron density models
W150-199	Parameters for perturbations to electron density models
W200-249	Parameters for analytic magnetic field models
W250-299	Parameters for analytic collision frequency models

\*These values have been initialized in the main program but may be reset by reading them in. See Appendix 1b for the initial values.

A second group of input data cards are necessary if nonanalytic ionospheric models such as the electron density profile defined by subroutine TABLEX or the collision frequency profile defined by subroutine TABLEZ are used. Each subroutine defining a nonanalytic ionospheric model reads in data cards according to a format defined in that subroutine. An element in the W array controls the reading of these cards. (See table 2.) Figure 4 shows the order in which these data cards should be arranged.

## 11. ACCURACY

The numerical integration subroutine has a built-in mechanism to check errors and adjust the integration step length accordingly. If the errors get larger than a maximum specified by the user, the routine will decrease the step length in order to maintain the accuracy. On the other hand, if the accuracy is greater than that required by the user, the routine will increase the step length in order to reduce the computing cost. The user specifies the desired accuracy in W42 (see table 2). W42 is the maximum allowable relative error in any single step for any of the equations being integrated. To get a very accurate (but expensive) ray trace, one can use a small W42 (about  $10^{-5}$  or  $10^{-6}$ ). For a cheap, approximate ray trace, one should use a large W42 ( $10^{-3}$  or even  $10^{-2}$ ). For cases in which all of the variables being integrated increase monotonically, the total relative error can be guaranteed to be less than W42. Otherwise, the total relative error cannot be easily estimated.

The far left column of the printout from the ray path calculation gives an indication of the integration error in the magnitude of the vector which points in the wave normal direction. Although the calculation of this error is made independently of the error calculation in the numerical integration routine, we have found that except near reflection for vertical or near vertical incidence this error is usually of the same order

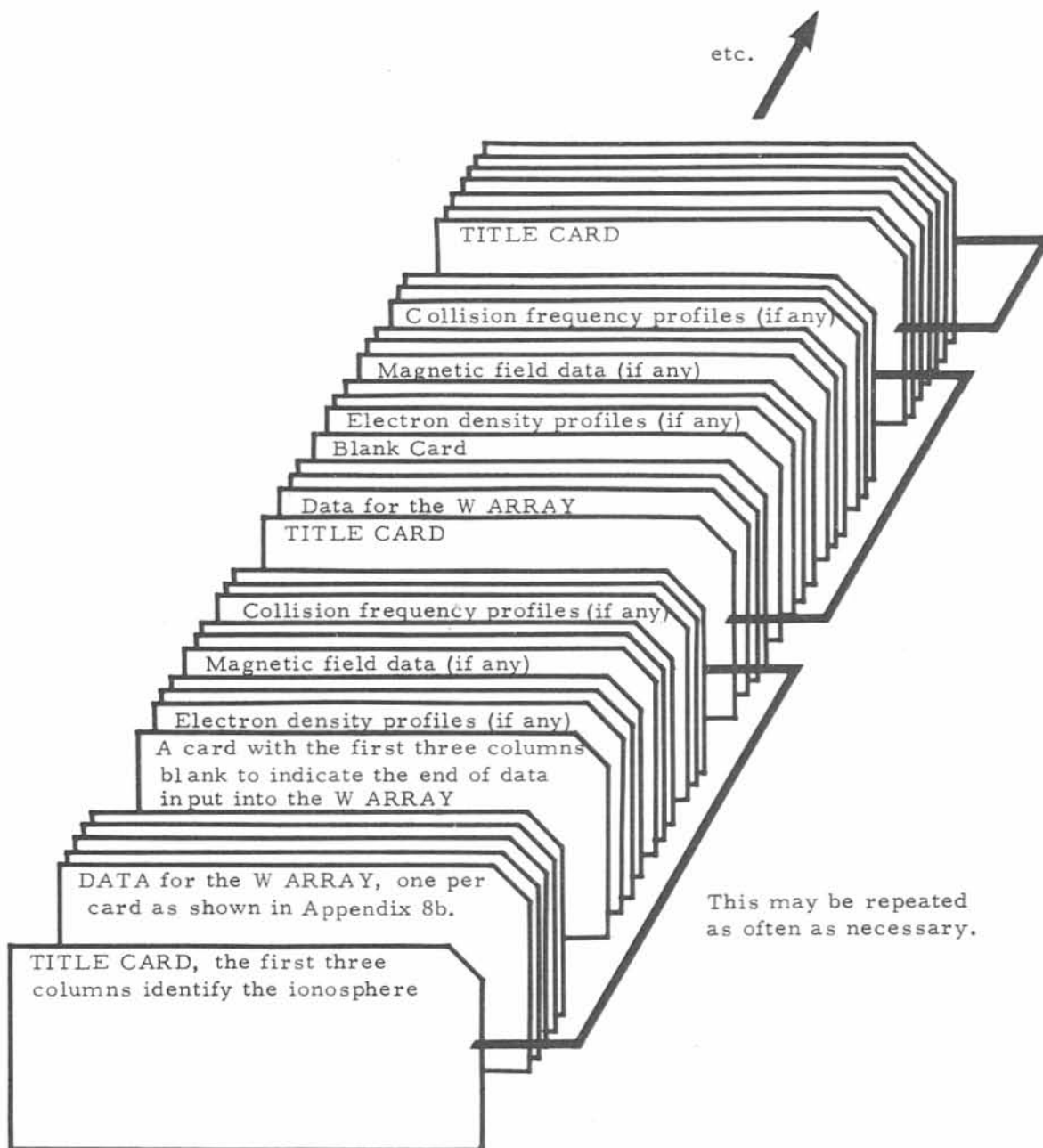


Figure 4. Data deck set-up.

of magnitude as that specified in W42. We have found that whenever this error has exceeded W42 by several orders of magnitude, the electron density subroutine we had written was calculating a gradient of electron density inconsistent with the spatial variation of electron density being calculated. See the general description of electron density models in Appendix 3a for more information.

## 12. COORDINATE SYSTEMS

The program uses two different spherical polar coordinate systems, namely, a geographic and a computational coordinate system. Input data for the coordinates of the transmitter (W4 and W5) and input data for the coordinates of the north pole of the computational coordinate system (W24 and W25) are entered in geographic coordinates. (Putting W25 equal to  $0^\circ$  and W24 equal to  $90^\circ$  would superimpose the two north poles and equate the two coordinate systems.)

When the two coordinate systems do not coincide, the three types of ionospheric models calculate electron density, the earth's magnetic field, and collision frequency in terms of the computational coordinate system. In particular, the dipole model of the earth's magnetic field uses the axis of the computational coordinate system as the axis for the dipole field. Thus, when using this dipole model, the computational coordinate system is a geomagnetic coordinate system, and both electron density and collision frequency must be defined in geomagnetic coordinates. Dudziak (1961) describes the transformations between these coordinate systems.

### 13. HOW THE PROGRAM WORKS

This ray tracing program consists of various subroutines that perform specific tasks in calculating ray paths. This division of labor facilitates modifying the program to solve specific problems. Often it may be necessary to change only one or two subroutines to convert the program to a different use.

The main program (NITIAL) sets up the initial conditions (transmitter location, wave frequency, and direction of transmission) for each ray trace. In setting up the initial conditions for each ray trace, the main program (NITIAL) steps frequency, azimuth angle of transmission, and elevation angle of transmission. The details of the workings of NITIAL can be found in the flow chart in figure 5. Then subroutine TRACE calculates one ray path for the requested number of crossings of the specified receiver height. Subroutine TRACE is the heart of the ray tracing program. It is the most complicated subroutine included, but also the most important to understand. The flow chart in figure 6 should help to explain TRACE.

Subroutine RKAM integrates the differential equations numerically using an Adams-Moulton predictor-corrector method with a Runge-Kutta starter. Subroutine HAMLTN evaluates the differential equations to be integrated. Subroutine RINDEX calculates the phase refractive index and its gradients, the group refractive index, and the polarization. (Eight versions of subroutine RINDEX are included.) Subroutines ELECTX, ELECT1, MAGY, and COLFRZ calculate the ionospheric electron density, perturbations to the electron density (irregularities), the earth's magnetic field, and the electron collision frequency, respectively. Several versions of these four subroutines are included and it is easy to add more. Subroutine REACH calculates a straight-line segment of a ray path in free space between the earth and the

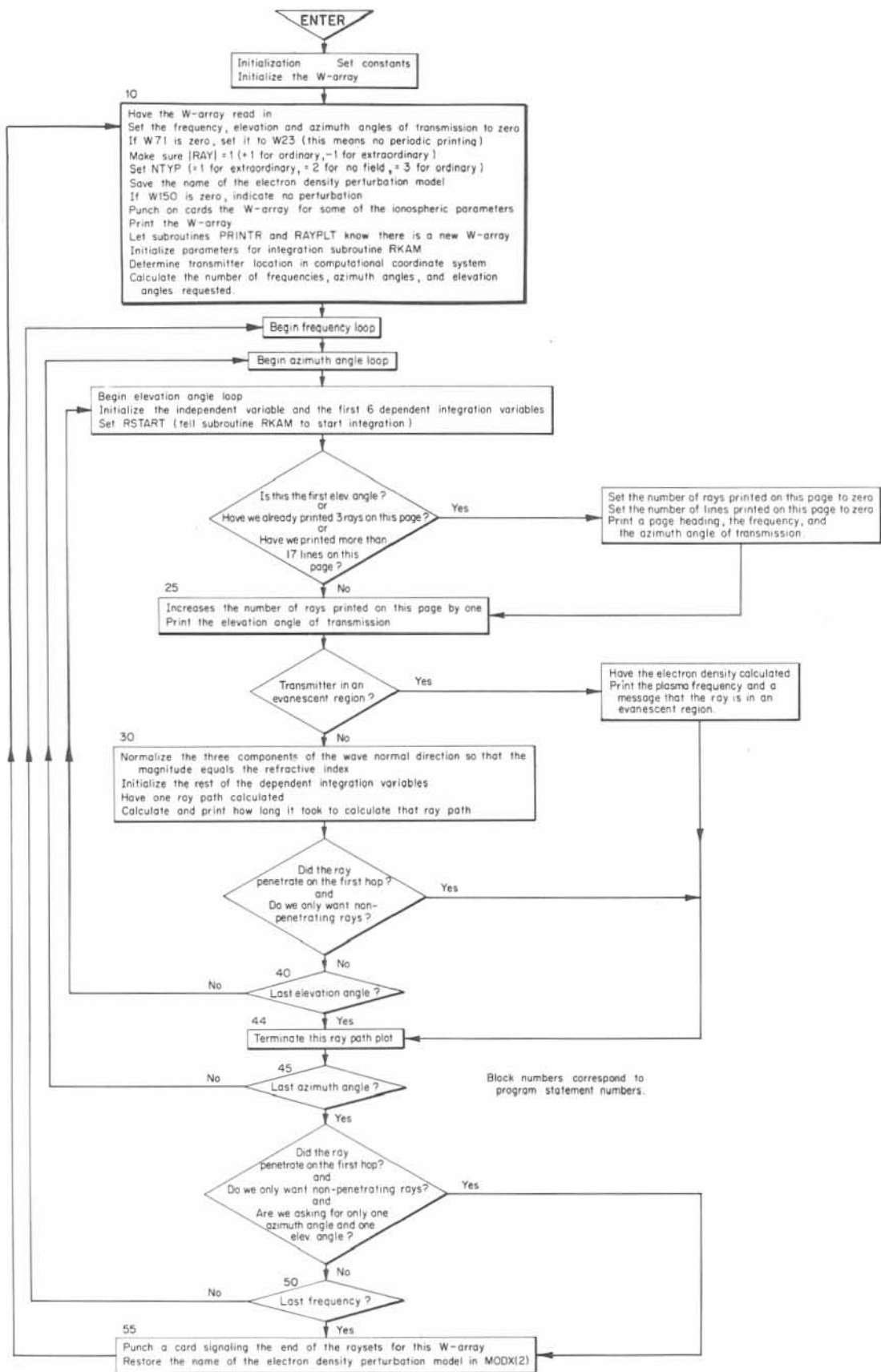


Figure 5. Flow chart for program NITIAL.



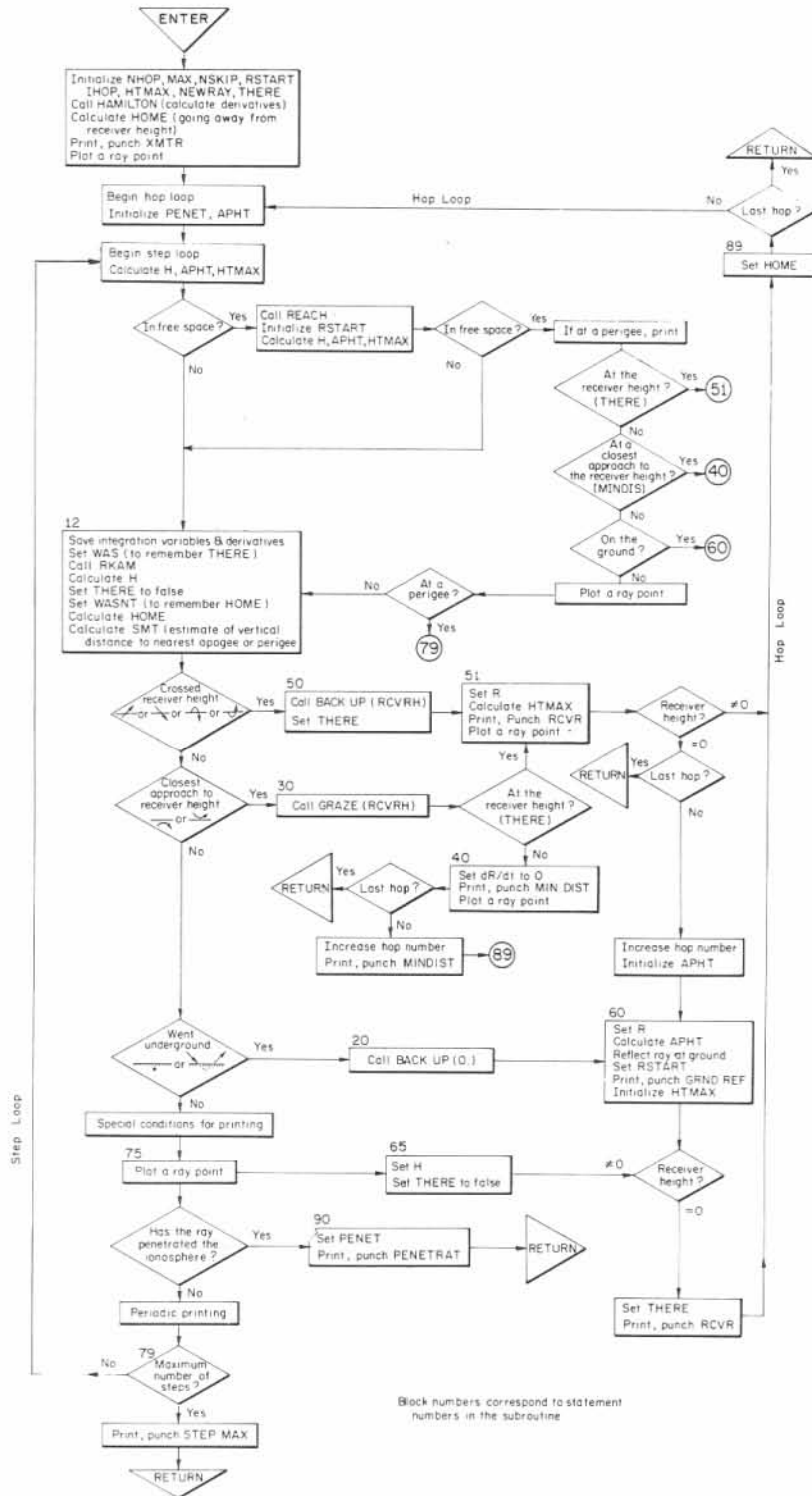


Figure 6. Flow chart for subroutine TRACE.

The ray graphics illustrate the path of a ray during a single step.

ionosphere or between ionospheric layers. Subroutine BACK UP finds an intersection of the ray with the receiver height or with the ground. Subroutine PRINTR prints information describing the ray path and punches the results on cards (raysets). Subroutine RAYPLT plots the ray path. The block diagram in figure 7 shows the relationship among these (and other) subroutines.

The listings of most of the subroutines have comments that should help in understanding how they work. In addition, Tables 3 through 14 define the variables in the common blocks.

#### 14. ACKNOWLEDGMENTS

Part of the organization of this program into subroutines follows that of the program of Dudziak (1961), in particular for subroutines RKAM, HAMLTN , RINDEX, ELECTX, MAGY, and COLFRZ. Also, the coordinate transformation in subroutine PRINTR and the method for data input via the W array are taken from the program of Dudziak (1961). The term "rayset," the idea of punching results of each hop for each ray trace onto cards, and the idea of automatically plotting ray paths come from the program of Croft and Gregory (1963). The quasi-parabolic layer electron density model QPARAB is taken from the paper by Croft and Hoogasian (1968). Notice that the quasi-parabolic layer that is now in the program is slightly different from the one in the program of Jones (1966). Subroutine RKAM is a modification of subroutine RKAMSUB, which was written by G. J. Lastman and is available through the CDC CO-OP library (the CO-OP identification is D2 UTEX RKAMSUB). Subroutine GAUSEL was written by L. David Lewis, Space Environment Laboratory, National Oceanic and Atmospheric Administration. Subroutine FSW was written in conjunction with Helmut Kopka of the Max-Planck-Institut für Aeronomie, Lindau/Harz, Germany.

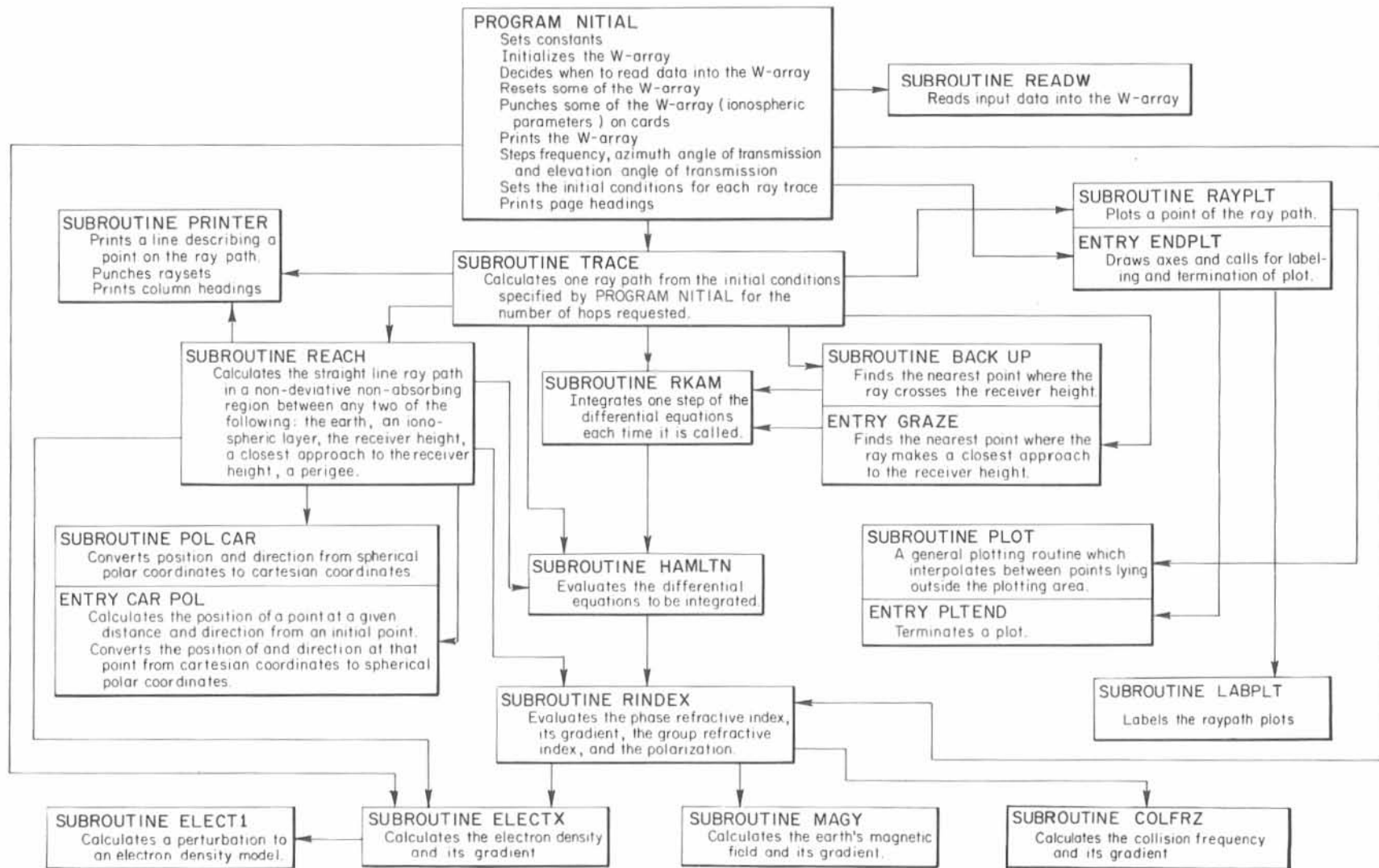


Figure 7. Block diagram for the ray tracing program.

Table 3. Definitions of the Parameters in Blank Common

Position in Common	Variable Name	Definition
1-20	R	The dependent variables in the differential equations being integrated-- the definitions of the first six are fixed, but the others may be varied by the program user.
1	R(1)	$r$
2	R(2)	$\theta$
3	R(3)	$\varphi$
4	R(4)	$k_r$
5	R(5)	$k_\theta$
6	R(6)	$k_\varphi$
7-12	R(7)-R(12)	Those variables the user has chosen to integrate, taken in the following order: P -phase path in kilometers A -absorption in decibels $\Delta f$ -Doppler shift in hertz s -geometrical path length in kilometers
13-20	R(13)-R(20)	Reserved for future expansion.
21	T	Group path in kilometers (the independent variable in the differential equations).
22	STP	Step length in group path.
23-42	DRDT	The derivatives of the dependent variables with respect to the independent variable T.

R and T are initialized in program NITIAL and changed in subroutines RKAM, REACH, and BACK UP.

STP is calculated in subroutine RKAM.

DRDT is calculated in subroutine HAMLTN and used in subroutine RKAM.

Table 4. Definitions of the Parameters in Common Block /CONST/

Position in Common	Variable Name	Definition
1	PI	$\pi$
2	PIT2	$2\pi$
3	PID2	$\pi/2$
4	DEGS	$180.0/\pi$
5	RAD	$\pi/180.0$
6	K	Ratio of the square of the plasma frequency to the electron density in $\text{MHz}^2 \text{cm}^3 = r_0^2 c^2 / \pi = e^2 / (4\pi^2 \epsilon_0 m)$ , where $r_0$ is the classical electron radius, $c$ is the free space speed of light, $e$ is the charge on the electron, $m$ is the mass of the electron, and $\epsilon_0$ is the capacitivity of a vacuum.
7	C	Free space speed of light in km/sec.
8	LOGTEN	$\log_e 10$

These parameters are set in program NITIAL.

Table 5. Definitions of the Parameters in Common Block /RK/

Position in Common	Variable Name	Definition
1	N	The number of equations being integrated.
2	STEP	The initial step in group path in kilometers.
3	MODE	Defines type of integration used (same as W41), see Table 2.
4	E1MAX	Maximum allowable single step error (same as W42).
5	E1MIN	Minimum allowable single step error (= W42/W43).
6	E2MAX	Maximum step length (same as W45).
7	E2MIN	Minimum step length (same as W46).
8	FACT	Factor by which to increase or decrease step length (same as W47).
9	RSTART	Nonzero to initialize numerical integration, zero to continue integration.

These parameters are calculated in program NITIAL (some are temporarily reset in subroutine BACK UP) and are used in subroutine RKAM.

Table 6. Definition of the Parameters in Common Block /RIN/

Postion in Common	Variable Name	Definition
1, 2, 3	MODRIN	Description of version of RINDEX in BCD.
4	COLL	= 1 if this version of RINDEX includes collisions, = 0 otherwise.
5	FIELD	= 1 if this version of RINDEX includes the earth's magnetic field, = 0 otherwise.
6	SPACE	TRUE, if the ray is in a nondeviative, nonabsorbing medium.
7, 8	KAY2	$k^2$ , square of the complex phase refractive index times $\omega^2/c^2$ .
9, 10	H	Hamiltonian (complex)
11, 12	PHPT	$\partial H/\partial t$ (complex)
13, 14	PHPR	$\partial H/\partial r$ (complex)
15, 16	PHPTH	$\partial H/\partial \theta$ (complex)
17, 18	PHPPH	$\partial H/\partial \varphi$ (complex)
19, 20	PHPOM	$\partial H/\partial \omega$ (complex)
21, 22	PHPKR	$\partial H/\partial k_r$ (complex)
23, 24	PHPKTH	$\partial H/\partial k_\theta$ (complex)
25, 26	PHPKPH	$\partial H/\partial k_\varphi$ (complex)
27, 28	KPHPK	$\vec{k} \cdot \partial H/\partial \vec{k}$ (complex)  $= k_r \partial H/\partial k_r + k_\theta \partial H/\partial k_\theta + k_\varphi \partial H/\partial k_\varphi$
29, 30	POLAR	Characteristic polarization of the wave; equal to the ratio of the component of the electric field perpendicular with the earth's magnetic field to the transverse component of the electric field parallel with the earth's magnetic field (complex) (Budden, 1961, p. 49, eq. (5.13)).

Table 6. (Continued)

Position in Common	Variable Name	Definition
31, 32	LPOLAR	Characteristic longitudinal polarization of the wave; equal to the ratio of the longitudinal component of the electric field to the component of the electric field perpendicular with the earth's magnetic field. (complex) Budden, 1961, p. 54, eq. (5.38)).
33	SGN	= +1 or -1; used for ray tracing in complex space.

These parameters are calculated in subroutine RINDEX and used in subroutine HAMLTN.

Note: In some subroutines, the real and imaginary parts of the complex variables have separate names.



Table 7. Definitions of the Parameters in Common Block /FLG/  
(See Subroutine TRACE)

Position in Common	Variable Name	Definition
1	NTYP	= 1 for extraordinary, = 2 for no field, = 3 for ordinary
2	NEWWR	Set equal to .TRUE. to tell subroutine RAYPLT there is a new W array.
3	NEWWP	Set equal to .TRUE. to tell subroutine PRINTR there is a new W array.
4	PENET	Set equal to .TRUE. if the ray just penetrated.
5	LINES	Number of lines printed on the current page.
6	IHOP	Hop number (at the beginning of each ray, subroutine TRACE sets this parameter to zero so that subroutine RAYPLT will begin a new line in plotting the ray path and subroutine PRINTR will print column headings and punch a transmitter rayset).
7	HPUNCH	The height to be punched on the raysets.

Table 8. Definitions of the Parameters in Common Block /XX/

Position in Common	Variable Name	Definition
1	MODX(1)	BCD name of the electron density subroutine.
2	MODX(2)	BCD name of the subroutine defining a perturbation to the electron density model.
3	X	X in Appleton-Hartree formula, square of the ratio of the plasma frequency to the wave frequency.
4	PXPR	$\frac{\partial X}{\partial r}$
5	PXPTH	$\frac{\partial X}{\partial \theta}$
6	PXPPH	$\frac{\partial X}{\partial \varphi}$
7	PXPT	$\frac{\partial X}{\partial t}$ , where t is time; used for calculating Doppler shifts.
8	HMAX	Height of maximum electron density.

These parameters are calculated in subroutine ELECTX, possibly modified in subroutine ELECT1, and are mainly used in subroutine RINDEX.

Table 9. Definitions of the Parameters in Common Block /YY/

Position in Common	Variable Name	Definition
1	MODY	BCD name of the subroutine defining the earth's magnetic field.
2	Y	Y in the Appleton-Hartree formula, ratio of the electron gyrofrequency to the wave frequency.
3	PYPR	$\frac{\partial Y}{\partial r}$
4	PYPTH	$\frac{\partial Y}{\partial \theta}$
5	PYPPH	$\frac{\partial Y}{\partial \varphi}$
6	YR	$Y_r$ , proportional to the component of the earth's magnetic field in the r direction.
7	PYRPR	$\frac{\partial Y_r}{\partial r}$
8	PYRPT	$\frac{\partial Y_r}{\partial \theta}$
9	PYRPP	$\frac{\partial Y_r}{\partial \varphi}$
10	YTH	$Y_\theta$
11	PYTPR	$\frac{\partial Y_\theta}{\partial r}$
12	PYTPT	$\frac{\partial Y_\theta}{\partial \theta}$

Table 9. (Continued)

Position in Common	Variable Name	Definition
13	PYTPP	$\frac{\partial Y}{\partial \varphi}$
14	YPH	$Y_{\varphi}$
15	PYPPR	$\frac{\partial Y}{\partial r}$
16	PYPPT	$\frac{\partial Y}{\partial \theta}$
17	PYPPP	$\frac{\partial Y}{\partial \varphi}$

These parameters are calculated in subroutine MAGY and are mainly used in subroutine RINDEX.

Table 10. Definitions of the Parameters in Common Block /ZZ/

Position in Common	Variable Name	Definition
1	MODZ	BCD name of the collision frequency subroutine.
2	Z	Z in the Appleton-Hartree formula, ratio of the electron-neutral collision frequency to the angular wave frequency.
3	PZPR	$\frac{\partial Z}{\partial r}$
4	PZPTH	$\frac{\partial Z}{\partial \theta}$
5	PZPPH	$\frac{\partial Z}{\partial \varphi}$

These parameters are calculated in subroutine COLFRZ and are mainly used in subroutine RINDEX.

Table 11. Definitions of the Parameters in Common Block /TRAC/

Position in Common	Variable Name	Definition
1	GROUND	.TRUE. if the ray is on the surface of the earth.
2	PERIGE	.TRUE. if the ray has just made a perigee.
3	THERE	.TRUE. if the ray is at the receiver height.
4	MINDIS	.TRUE. if the ray has just made a closest approach to the receiver height.
5	NEWRAY	Set equal to .TRUE. to tell subroutine REACH that this is a new ray.
6	SMT	An estimation of the vertical distance to an apogee or perigee of the ray.

These parameters are used for communication between subroutine TRACE and subroutines REACH and BACK UP.

Table 12. Definition of the Parameter in Common Block /COORD/

Position in Common	Variable Name	Definition
1	S	The straight line distance along the ray from the position of the ray where REACH was called to the present position.

This parameter is used for communication between subroutine REACH and subroutine POL CAR.

Table 13. Definitions of the Parameters in Common Block /PLT/

Position in Common	Variable Name	Definition
1	XMIN0, XL	The x coordinate of the left side of the plotting area in kilometers.
2	XMAX0, XR	The x coordinate of the right side of the plotting area in kilometers.
3	XMIN0, YB	The y coordinate of the bottom of the plotting area in kilometers.
4	YMAX0, YT	The y coordinate of the top of the plotting area in kilometers.
5	RESET	Set equal to one whenever the plotting area is changed.

These parameters are used for communication between subroutine RAYPLT and subroutine PLOT.

Table 14. Definitions of the Parameters in Common Block /DD/

Position in Common	Variable Name	Definition
1	IN	Intensity. IN = 0 specifies normal intensity. IN = 1 specifies high intensity.
2	IOR	Orientation. IOR = 0 specifies upright orientation. IOR = 1 specifies rotated orientation (90° counterclockwise).
3	IT	Italics (Font). IT = 0 specifies non-Italic (Roman) symbols. IT = 1 specifies Italic symbols.
4	IS	Symbol size. IS = 0 specifies miniature size. IS = 1 specifies small size. IS = 2 specifies medium size. IS = 3 specifies large size.
5	IC	Symbol case. IC = 0 specifies upper case. IC = 1 specifies lower case.
6	ICC	Character code, 0-63 (R1 format). ICC and IC together specify the symbol plotted.
7	IX	X-coordinate, 0-1023.
8	IY	Y-coordinate, 0-1023.



We also want to thank those who have used our program and have pointed out errors or made suggestions. In particular, we are grateful to Dr. T. M. Georges of the Wave Propagation Laboratory, National Oceanic and Atmospheric Administration, for his suggestions resulting from extensive use of the program, for development of some of the ionospheric models (DCHAPT, DTORUS, WAVE, WAVE2), and for financing part of the development of ray tracing through a spitze.

Examples of use of the ray tracing program are shown in the reports by Stephenson and Georges (1969) and Georges (1971).

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## APPENDIX 1. LISTINGS OF THE MAIN PROGRAM AND SUBROUTINES IN THE MAIN DECK

This appendix contains listings of the main program and those subroutines that have only one version (with the exception of subroutine RAYPLT, which has a do-nothing version for users lacking a plotter to plot ray paths). Thus, the routines which form the contents of this appendix will be used in all ray path calculations.

Additionally, this appendix contains the main input parameter form for ray tracing and the input parameter forms for plotting. These forms are very useful when using the program because they indicate the input parameters needed for ray path calculations

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PROGRAM NITIAL
C      SETS THE INITIAL CONDITIONS FOR EACH RAY AND CALLS TRACE
DIMENSION MFLD(2)
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,C,LOGTEN
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR,
1      LPOLAR,SGN
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPH,PXPT,HMAX
COMMON /YY/ MODY,Y(16) /ZZ/ MODZ,Z(4)
COMMON R(20),T,STP,ORDT(20) /WW/ IJ(10),W0,W(400)
EQUIVALENCE (RAY,W(1)),(EARTH,W(2)),(XMTRH,W(3)),(TLAT,W(4)),
1 (TLON,W(5)),(F,W(6)),(FBEG,W(7)),(FEND,W(8)),(FSTEP,W(9)),
2 (AZ1,W(10)),(AZBEG,W(11)),(AZEND,W(12)),(AZSTEP,W(13)),
3 (BETA,W(14)),(ELBEG,W(15)),(ELEND,W(16)),(ELSTEP,W(17)),
4 (ONLY,W(21)),(HOP,W(22)),(MAXSTP,W(23)),(PLAT,W(24)),(PLON,W(25)),
5 (INTYP,W(41)),(MAXERR,W(42)),(ERATIO,W(43)),(STEP1,W(44)),
6 (STPMAX,W(45)),(STPMIN,W(46)),(FACTR,W(47)),(SKIP,W(71)),
7 (RAYSET,W(72)),(PLT,W(81)),(PERT,W(150))
LOGICAL SPACE,NEWWR,NEWWP,PENET
REAL N2,N2I,LOGTEN,K,MAXSTP,INTYP,MAXERR,MU
COMPLEX PNP,POLAR,LPOLAR
NDATE=IDATE(0)
SECOND=KLOCK(0)*.001
COLL=4H NO
IF (COLL.NE.0.) KOLL=4HWITH
C***** CONSTANTS
PI=3.1415926536
PIT2=2.*PI
PID2=PI/2.
DEGS=180./PI
RAD=PI/180.
C=2.997925E5
K=2.81785E-15*C**2/PI
LOGTEN=ALOG(10.)
C***** INITIALIZE SOME VARIABLES IN THE W ARRAY
DO 5 NW=1,400
5 W(NW)=0.
PLON=0.
PLAT=PID2
EARTH=6370.
INTYP=3.
MAXERR=1.E-4
ERATIO=50.
STEP1=1.
STPMAX=100.
STPMIN=1.E-8
FACTR=0.5
MAXSTP=1000.
HOP=1.
C***** READ W ARRAY AND PRINT NON-ZERO VALUES
10 CALL READ W
F=BETA-AZ1=0.
IF (SKIP.EQ.0.) SKIP=MAXSTP
12 RAY=SIGN(1.,RAY)
NTYP=2.+FIELD*RAY
GO TO (13,14,15),NTYP
13 MFLD(1)=8HEXTRORD
MFLD(2)=5HINARY
GO TO 16
14 MFLD(1)=8HNO FIELD
MFLD(2)=1H
GO TO 16
NIT I001
NIT I002
NIT I003
NIT I004
NIT I005
NIT I006
NIT I007
NIT I008
NIT I009
NIT I010
NIT I011
NIT I012
NIT I013
NIT I014
NIT I015
NIT I016
NIT I017
NIT I018
NIT I019
NIT I020
NIT I021
NIT I022
NIT I025
NIT I026
NIT I027
NIT I028
NIT I029
NIT I030
NIT I031
NIT I032
NIT I033
NIT I034
NIT I035
NIT I036
NIT I037
NIT I038
NIT I039
NIT I040
NIT I041
NIT I042
NIT I043
NIT I044
NIT I045
NIT I046
NIT I047
NIT I048
NIT I049
NIT I050
NIT I051
NIT I052
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NIT I055
NIT I056
NIT I057
NIT I058
NIT I059
NIT I060
NIT I061
NIT I062
NIT I063
NIT I064
NIT I065

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15 MFLD(1)=8HORDINARY                                NIT1066
MFLD(2)=8H                                            NIT1067
16 MODSAV=MODX(2)                                    NIT1068
IF (PERT.EQ.0.) MODX(2)=6H                            NIT1069
IF (RAYSET.NE.0.) PUNCH 2000, ID,MODX(1),(W(NW),NW=101,107), NIT1070
1 MODX(2),(W(NW),NW=151,157),MODY,(W(NW),NW=201,207), NIT1071
2 MODZ,(W(NW),NW=251,257)                             NIT1072
2000 FORMAT (10A8,4(/A8,2X,7E10.3))                  NIT1073
PRINT 1000, ID,NUATE,MODX,MODY,MODZ,MODRIN,MFLD,KOLL NIT1074
1000 FORMAT (1H1,10A8,25X,A8/4(1X,A8),24X,5A8,1X,A0,A2,1X,A4, NIT1075
1 11H COLLISIONS/)                                    NIT1076
PRINT 1050                                             NIT1077
1050 FORMAT (85H INITIAL VALUES FOR THE W ARRAY -- ALL ANGLES IN RADIAN NIT1078
1S, ONLY NONZERO VALUES PRINTED/)                  NIT1079
DO 17 NW=1,400                                        NIT1080
IF (W(NW).NE.0.) PRINT 1700, NW,W(NW)               NIT1081
1700 FORMAT (I4,E19.11)                               NIT1082
17 CONTINUE                                           NIT1083
C***** LET SUBROUTINES PRINTR AND RAYPLT KNOW THERE IS A NEW W ARRAY NIT1084
NEWWP=.TRUE.                                          NIT1085
NEWWR=.TRUE.                                          NIT1086
C***** INITIALIZE PARAMETERS FOR INTEGRATION SUBROUTINE RKAM NIT1087
N=6                                                    NIT1088
DO 20 NR=7,20                                         NIT1089
IF (W(50+NR).NE.0.) N=N+1                            NIT1090
20 CONTINUE                                           NIT1091
MODE=INTYP                                           NIT1092
STEP=STEP1                                           NIT1093
E1MAX=MAXERR                                         NIT1094
E1MIN=MAXERR/ERATIO                                  NIT1095
E2MAX=STPMAX                                         NIT1096
E2MIN=STPMIN                                         NIT1097
FACT=FACTR                                           NIT1098
C***** DETERMINE TRANSMITTER LOCATION IN COMPUTATIONAL COORDINATE NIT1099
C***** SYSTEM (GEOMAGNETIC COORDINATES IF DIPOLE FIELD IS USED) NIT1100
R0=EARTH+XMTRH                                       NIT1101
SP=SIN (PLAT)                                         NIT1102
CP=SIN (PID2-PLAT)                                    NIT1103
SDPH=SIN (TLON-PLON)                                  NIT1104
CDPH=SIN (PID2-(TLON-PLON))                          NIT1105
SL=SIN (TLAT)                                         NIT1106
CL=SIN (PID2-TLAT)                                    NIT1107
ALPHA=ATAN2(-SDPH*CP,-CDPH*CP*SL+SP*CL)              NIT1108
TH0=ACOS (CDPH*CP*CL+SP*SL)                          NIT1109
PH0=ATAN2(SDPH*CL,CDPH*SP*CL-CP*SL)                  NIT1110
C***** LOOP ON FREQUENCY, AZIMUTH ANGLE, AND ELEVATION ANGLE NIT1111
NFREQ=1                                               NIT1112
IF (FSTEP.NE.0.) NFREQ=(FEND-FBEG)/FSTEP+1.5        NIT1113
NAZ=1                                                 NIT1114
IF (AZSTEP.NE.0.) NAZ=(AZEND-AZBEG)/AZSTEP+1.5      NIT1115
NBETA=1                                               NIT1116
IF (ELSTEP.NE.0.) NBETA=(ELEND-ELBEG)/ELSTEP+1.5    NIT1117
DO 50 NF=1,NFREQ                                     NIT1118
F=FBEG+(NF-1)*FSTEP                                  NIT1119
DO 45 J=1,NAZ                                        NIT1120
AZ1=AZBEG+(J-1)*AZSTEP                               NIT1121
AZA=AZ1*DEGS                                         NIT1122
GAMMA=PI-AZ1+ALPHA                                    NIT1123
SGAMMA=SIN (GAMMA)                                   NIT1124
CGAMMA=SIN (PID2-GAMMA)                              NIT1125
DO 40 I=1,NBETA                                      NIT1126
BETA=ELBEG+(I-1)*ELSTEP                             NIT1127

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      FL=BETA*DFGS
      CBETA=SIN (PID2-BETA)
      R(1)=R0
      R(2)=TH0
      R(3)=PH0
      R(4)=SIN (BETA)
      R(5)=CBETA*CGAMMA
      R(6)=CBETA*SGAMMA
      T=0.
      RSTART=1.
C      SGN=1. (NEED FOR RAY TRACING IN COMPLEX SPACE.)
C***** ALLOW IONOSPHERIC MODEL SUBROUTINES TO READ AND PRINT DATA
      CALL RINDEX
      IF (I.NE.1.AND.NPAGE.LT.3.AND.LINES.LE.17) GO TO 25
      NPAGE=LINFS=0
      PRINT 1000, ID,NDATE,MODX,MODY,MODZ,MODKIN,MFLD,KULL
      PRINT 2400, F,AZA
2400  FORMAT (18X,11HFREQUENCY =,F12.6,37H MHZ, AZIMUTH ANGLE OF TRANSMISSION =,F12.6,4H DEG)
      25 NPAGE=NPAGE+1
      PRINT 2500, EL
2500  FORMAT (/31X,33HELEVATION ANGLE OF TRANSMISSION =,F12.6,4H DEG/)
      IF (N2.GT.0.) GO TO 30
      CALL ELECTX
      FN=SIGN (SQRT (ABS (X))*F,X)
      PRINT 2900, FN
2900  FORMAT (58H0TRANSMITTER IN EVANESCENT REGION, TRANSMISSION IMPOSSIBLE/20H0PLASMA FREQUENCY = ,E17.10)
      GO TO 44
      30 MU=SQRT (N2/(R(4)**2+R(5)**2+R(6)**2))
      DO 34 NN=4,6
      34 R(NN)=R(NN)*MU
      DO 35 NN=7,N
      35 R(NN)=0.
      CALL TRACE
      OSEC=SECOND
      SECOND=KLOCK(0)*.001
      DIFF=SECOND-OSEC
      PRINT 3500, DIFF
3500  FORMAT (36X,26HTHIS RAY CALCULATION TOOK ,F8.3,4H SEC)
      IF (PENET.AND.ONLY.NE.0..AND.IHOP.EQ.1) GO TO 44
      40 CONTINUE
      44 IF (PLT.NE.0.) CALL ENDPLT
      45 CONTINUE
      IF (PENET.AND.ONLY.NE.0..AND.IHOP.EQ.1.AND.NAZ.EQ.1.AND.NBETA.EQ.1)
      1 GO TO 55
      50 CONTINUE
      55 IF (RAYSET.NE.0.) PUNCH 5000
5000  FORMAT (78X,1H-)
      MODX(2)=MODSAV
      GO TO 10
      END

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NIT1128
NIT1129
NIT1130
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NIT1176
NIT1177
NIT1178
NIT1179-

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SUBROUTINE READ W                                READ001
C                                                READ002
C          READS W ARRAY                          READ003
C          A 1 IN THE FOLLOWING COLUMNS WILL MAKE THE DESCRIBED CONVERSIONS
COL 18 DEGREES TO RADIANS                        READ004
COL 19 GREAT CIRCLE DISTANCE IN KM TO RADIANS   READ005
COL 20 NAUTICAL MILES TO KM                     READ006
COL 21 FEET TO KM                               READ007
C                                                READ008
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,DUM(3)     READ009
COMMON /WW/ ID(10),W0,W(400)                    READ010
EQUIVALENCE (EARTH,W(2))                        READ011
INTEGER DEG,FEET                                READ012
READ 1000, ID                                    READ013
1000 FORMAT (10A8)                               READ014
      IF (EOF,60) 3,4                             READ015
      3 CALL EXIT                                 READ016
      4 READ 1100, NW,W(NW),DEG,KM,NM,FEET       READ017
1100 FORMAT (I3,E14.7,5I1)                       READ018
      IF (NW.EQ.0) GO TO 10                       READ019
      IF (NW.GT.0.AND.NW.LE.400) GO TO 5         READ020
      PRINT 4000, NW                              READ021
4000 FORMAT(15H1THE SUBSCRIPT ,I3,77H ON THE W-ARRAY INPUT IS OUT OF BOREAD022
      UNDS. ALLOWABLE VALUES ARE 1 THROUGH 400. ) READ023
      CALL EXIT                                  READ024
      5 IF (DEG.NE.0.) W(NW)=W(NW)*RAD           READ025
      IF (KM.NE.0) W(NW)=W(NW)/EARTH            READ026
      IF (NM.NE.0) W(NW)=W(NW)*1.852           READ027
      IF (FEET.NE.0) W(NW)=W(NW)*3.048006096E-4 READ028
      GO TO 4                                    READ029
10 RETURN                                        READ030
      END                                        READ031-

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SUBROUTINE TRACE                                TRAC001
C          CALCULATES THE RAY PATH               TRAC002
C          DIMENSION ROLD(20),OROLD(20)         TRAC003
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART TRAC004
COMMON /FLG/ NTYP,NEWHR,NEWWP,PENET,LINES,IHOP,HPUNCH      TRAC005
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT        TRAC006
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,PNP(10),POLAR,LPOLAR TRAC007
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX           TRAC008
COMMON R(20),T,STP,ORDT(20) /WW/ ID(10),W0,W(400)        TRAC009
LOGICAL SPACE,HOME,WASNT,UNDRGD,GROUND,PERIGE,THERE,MINDIS,NEWWR, TRAC010
1 NEWWP,PENET,NEWRAY,WAS                                TRAC011
REAL MAXSTP                                             TRAC012
COMPLEX N2,PNP,POLAR,LPOLAR                             TRAC013
EQUIVALENCE (EARTH,W(2)),(RCVRH,W(20)),(HOP,W(22)),(MAXSTP,W(23)) TRAC014
1,(SKIP,W(71)),(RAYSET,W(72)),(PLT,W(81))              TRAC015
NHOP=HOP                                               TRAC016
MAX=MAXSTP                                             TRAC017
NSKIP=SKIP                                             TRAC018
RSTART=1.                                             TRAC019
CALL HAMLTN                                           TRAC020
HOME=ORDT(1)*(R(1)-EARTH-RCVRH).GE.0.                 TRAC021
C***** IHOP=0 TELLS PRINTR TO PRINT HEADING AND PUNCH A TRANSMITTER TRAC022
C***** RAYSET AND TELLS RAYPLT TO START A NEW RAY    TRAC023
IHOP=0                                                TRAC024
CALL PRINTR (8HXMTR ,0.)                               TRAC025
IF (PLT.NE.0.) CALL RAYPLT                             TRAC026
HTMAX=0.                                              TRAC027
NEWRAY=.TRUE.                                         TRAC028
THERE=R(1)-EARTH.EQ.RCVRH                             TRAC029

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C***** LOOP ON NUMBER OF HOPS	TRAC030
10 IHOP=IHOP+1	TRAC031
IF (IHOP.GT.NHOP) RETURN	TRAC032
PENET=.FALSE.	TRAC033
APHT=RCVRH	TRAC034
C***** LOOP ON MAXIMUM NUMBER OF STEPS PER HOP	TRAC035
DO 79 J=1,MAX	TRAC036
H=R(1)-EARTH	TRAC037
IF (ABS(H-RCVRH).GT.ABS(APHT-RCVRH)) APHT=H	TRAC038
HTMAX=AMAX1(H,HTMAX)	TRAC039
IF (.NOT.SPACE) GO TO 12	TRAC040
CALL REACH	TRAC041
RSTART=1.	TRAC042
H=R(1)-EARTH	TRAC043
IF (ABS(H-RCVRH).GT.ABS(APHT-RCVRH)) APHT=H	TRAC044
HTMAX=AMAX1(H,HTMAX)	TRAC045
IF (.NOT.SPACE) GO TO 12	TRAC046
IF (PERIGE) CALL PRINTR (8HPERIGEE ,0.)	TRAC047
IF (THERE) GO TO 51	TRAC048
IF (MINDIS) GO TO 40	TRAC049
IF (GROUND) GO TO 60	TRAC050
IF (PLT.NE.0.) CALL RAYPLT	TRAC051
IF (PERIGE) GO TO 79	TRAC052
12 DO 13 L=1,N	TRAC053
ROLD(L)=R(L)	TRAC054
13 DROLD(L)=DRDT(L)	TRAC055
TOLD=T	TRAC056
WAS=THERE	TRAC057
CALL RKAM	TRAC058
H=R(1)-EARTH	TRAC059
THERE=.FALSE.	TRAC060
WASNT=.NOT.HOME	TRAC061
HOME=DRDT(1)*(H-RCVRH).GE.0.	TRAC062
TMP=(DRDT(1)-DROLD(1))*(T-TOLD)	TRAC063
SMT=0.	TRAC064
IF (TMP.NE.0.) SMT=0.5*(R(1)-ROLD(1)+0.5*TMP)**2/ABS(TMP)	TRAC065
IF (((H-RCVRH)*(ROLD(1)-EARTH-RCVRH).LT.0..AND..NOT.WAS).OR.	TRAC066
1 (WAS.AND.DRDT(1)*DROLD(1).LT.0..AND.HOME)) GO TO 50	TRAC067
IF (HOME.AND.WASNT) GO TO 30	TRAC068
IF (H.LT.0..OR.DRDT(1).GT.0..AND.DROLD(1).LT.0..AND.SMT.GT.H)	TRAC069
1 GO TO 20	TRAC070
IF (DROLD(1).LT.0..AND.DRDT(1).GT.0.) CALL PRINTR(8HPERIGEE ,0.)	TRAC071
IF (DROLD(1).GT.0..AND.DRDT(1).LT.0.) CALL PRINTR(8HAPOGEE ,0.)	TRAC072
IF (DROLD(2)*DRDT(2).LT.0.) CALL PRINTR(8HMAX LAT ,0.)	TRAC073
IF (DROLD(3)*DRDT(3).LT.0.) CALL PRINTR(8HMAX LONG,0.)	TRAC074
DO 14 I=4,6	TRAC075
IF (ROLD(I)*R(I).LT.0.) CALL PRINTR(8HWAVE REV,0.)	TRAC076
14 CONTINUE	TRAC077
GO TO 75	TRAC078
C***** RAY WENT UNDERGROUND	TRAC079
20 CALL BACK UP(0.)	TRAC080
GO TO 60	TRAC081
C***** RAY MAY HAVE MADE A CLOSEST APPROACH	TRAC082
30 CALL GRAZE(RCVRH)	TRAC083
IF (THERE) GO TO 51	TRAC084
40 DRDT(1)=0.	TRAC085
HPUNCH=R(1)-EARTH	TRAC086
CALL PRINTR(8HMIN DIST,RAYSET)	TRAC087
IF (PLT.NE.0.) CALL RAYPLT	TRAC088
IF (IHO>.GE.NHOP) RETURN	TRAC089
IHOP=IHOP+1	TRAC090
CALL PRINTR (8HMIN DIST,RAYSET)	TRAC091
GO TO 89	TRAC092

C***** RAY CROSSED RECEIVER HEIGHT	TRAC093
50 CALL BACK UP(RCVRH)	TRAC094
THERE=.TRUE.	TRAC095
51 R(1)=EARTH+RCVRH	TRAC096
HTMAX=AMAX1(RCVRH,HTMAX)	TRAC097
HPUNCH=APHT	TRAC098
CALL PRINTR(8HRCVR ,RAYSET)	TRAC099
IF (PLT.NE.0.) CALL RAYPLT	TRAC100
IF (RCVRH.NE.0.) GO TO 89	TRAC101
IF (IHOP.GE.NHOP) RETURN	TRAC102
IHOP=IHOP+1	TRAC103
APHT=RCVRH	TRAC104
C***** GROUND REFLECT	TRAC105
60 R(1)=EARTH	TRAC106
IF (ABS(RCVRH).GT.ABS(APHT-RCVRH)) APHT=0.	TRAC107
R(4)=ABS (R(4))	TRAC108
DRDT(1)=ABS (DRDT(1))	TRAC109
RSTART=1.	TRAC110
HPUNCH=HTMAX	TRAC111
CALL PRINTR(8HGRND REF,RAYSET)	TRAC112
HTMAX=0.	TRAC113
IF (RCVRH.NE.0.) GO TO 65	TRAC114
THERE=.TRUE.	TRAC115
HPUNCH=APHT	TRAC116
CALL PRINTR (8HRCVR ,RAYSET)	TRAC117
GO TO 89	TRAC118
55 H=0.	TRAC119
THERE=.FALSE.	TRAC120
C*****	TRAC121
75 IF (PLT.NE.0.) CALL RAYPLT	TRAC122
IF (H.GT.HMAX.AND.H.GT.RCVRH.AND.DRDT(1).GT.0.) GO TO 90	TRAC123
IF (MOD(J,NSKIP).EQ.0) CALL PRINTR(8H ,0.)	TRAC124
79 CONTINUE	TRAC125
C***** EXCEEDED MAXIMUM NUMBER OF STEPS	TRAC126
HPUNCH=H	TRAC127
CALL PRINTR(8HSTEP MAX,RAYSET)	TRAC128
RETURN	TRAC129
C*****	TRAC130
89 HOME=.TRUE.	TRAC131
GO TO 10	TRAC132
C***** RAY PENETRATED	TRAC133
90 PENET=.TRUE.	TRAC134
HPUNCH=H	TRAC135
CALL PRINTR(8HPENETRAT,RAYSET)	TRAC136
RETURN	TRAC137
END	TRAC138

SUBROUTINE BACK UP(HS)	BACK001
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART	BACK002
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT	BACK003
COMMON R(20),T,STP,DRDT(20) /WW/ ID(10),W0,W(400)	BACK004
EQUIVALENCE (EARTH,W(2)),(INTYP,W(41)),(STEP1,W(44))	BACK005
REAL INTYP	BACK006
LOGICAL GROUND,PERIGE,THERE,MINDIS,NEWRAY,HOME	BACK007
C***** DIAGNOSTIC PRINTOUT	BACK008
C CALL PRINTR (8HBACK UP0,0.)	BACK009
C***** GOING AWAY FROM THE HEIGHT HS	BACK010
HOME=DRDT(1)*(R(1)-EARTH-HS).GE.0.	BACK011
IF (HS.GT.0..AND..NOT.HOME.OR.HS.EQ.0..AND.DRDT(1).GT.0.) GO TO 30	BACK012

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C***** FIND NEAREST INTERSECTION OF RAY WITH THE HEIGHT HS          BACK013
DO 10 I=1,10                                                            BACK014
STEP=-(R(1)-EARTH-HS)/DRDT(1)                                          BACK015
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)                              BACK016
IF (ABS(R(1)-EARTH-HS).LT..5E-4.AND.ABS(STEP).LT.1.) GO TO 60        BACK017
C***** DIAGNOSTIC PRINTOUT                                            BACK018
C   CALL PRINTR(8HBACK UP1,0.)                                         BACK019
MODE=1                                                                    BACK020
RSTART=1.                                                                BACK021
CALL RKAM                                                                BACK022
10 RSTART=1.                                                            BACK023
C   BACK024
C***** FIND NEAREST CLOSEST APPROACH OF RAY TO THE HEIGHT HS        BACK025
ENTRY GRAZE                                                                BACK026
THERE=.FALSE.                                                            BACK027
C***** DIAGNOSTIC PRINTOUT                                            BACK028
C   CALL PRINTR (8HGRAZE 0 ,0.)                                         BACK029
IF (SMT.GT.ABS(R(1)-EARTH-HS)) GO TO 30                                  BACK030
DO 20 I=1,10                                                            BACK031
STEP=-R(4)/DRDT(4)                                                      BACK032
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)                              BACK033
IF (ABS(R(4)).LE.1.E-6.AND.ABS(STEP).LT.1.) GO TO 60                 BACK034
C***** DIAGNOSTIC PRINTOUT                                            BACK035
C   CALL PRINTR (8HGRAZE 1 ,0.)                                         BACK036
MODE=1                                                                    BACK037
RSTART=1.                                                                BACK038
CALL RKAM                                                                BACK039
RSTART=1.                                                                BACK040
IF (DRDT(4)*(R(1)-EARTH-HS).LT.0.) GO TO 30                            BACK041
IF (R(5).EQ.0..AND.R(6).EQ.0.) GO TO 60                                BACK042
20 CONTINUE                                                              BACK043
C***** IF A CLOSEST APPROACH COULD NOT BE FOUND IN 10 STEPS, IT      BACK044
C***** PROBABLY MEANS THAT THE RAY INTERSECTS THE HEIGHT HS          BACK045
30 CONTINUE                                                                BACK046
C***** DIAGNOSTIC PRINTOUT                                            BACK047
C   CALL PRINTR (8HBACK UP2,0.)                                         BACK048
MODE=1                                                                    BACK049
C***** ESTIMATE DISTANCE TO NEAREST INTERSECTION OF RAY WITH HEIGHT  BACK050
C***** HS BEHIND THE PRESENT RAY POINT                                BACK051
STEP=(-R(4)-SQRT(R(4)**2-2.*(R(1)-EARTH-HS)*DRDT(4)))/DRDT(4)        BACK052
RSTART=1.                                                                BACK053
CALL RKAM                                                                BACK054
RSTART=1.                                                                BACK055
C***** FIND NEAREST INTERSECTION OF RAY WITH HEIGHT HS              BACK056
DO 40 I=1,10                                                            BACK057
STEP=-(R(1)-EARTH-HS)/DRDT(1)                                          BACK058
STEP=SIGN(AMIN1(ABS(STP),ABS(STEP)),STEP)                              BACK059
IF (ABS(R(1)-EARTH-HS).LT..5E-4.AND.ABS(STEP).LT.1.) GO TO 60        BACK060
C***** DIAGNOSTIC PRINTOUT                                            BACK061
C   CALL PRINTR (8HBACK UP3,0.)                                         BACK062
MODE=1                                                                    BACK063
RSTART=1.                                                                BACK064
CALL RKAM                                                                BACK065
40 RSTART=1.                                                            BACK066
50 THERE=.TRUE.                                                          BACK067
C***** RESET STANDARD MODE AND INTEGRATION TYPE                      BACK068
60 MODE=INTYP                                                            BACK069
STEP=STEP1                                                                BACK070
RETURN                                                                    BACK071
END                                                                        BACK072-

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SUBROUTINE REACH
C          CALCULATES THE STRAIGHT LINE RAY PATH BETWEEN THE EARTH
C          AND THE IONOSPHERE OR BETWEEN IONOSPHERIC LAYERS
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON /TRAC/ GROUND,PERIGE,THERE,MINDIS,NEWRAY,SMT
COMMON /COORD/ S
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR,
1          LPOLAR
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON P(20),T,STP,DRDT(20) /HW/ ID(10),WD,W(400)
EQUIVALENCE (EARTH,R(2)),(XMTRH,W(3)),(RCVRH,W(20))
LOGICAL CROSS,CROSSG,CROSSR,SPACE,GROUND,PERIGE,THERE,MINDIS,
1          NEWRAY,RSPACE
REAL N2,N2I
COMPLEX PNP,POLAR,LPOLAR
DATA (NSTEP=500)
CALL HAPLTN
H=R(1)-EARTH
IF (.NOT.NEWRAY.AND..NOT.RSPACE) CALL PRINTR(8HEXIT ION,0.)
NEWRAY=.FALSE.
V=SQRT(R(4)**2+R(5)**2+R(6)**2)
C***** NORMALIZE THE WAVE NORMAL DIRECTION TO ONE
R(4)=R(4)/V
R(5)=R(5)/V
R(6)=R(6)/V
C***** NEGATIVE OF DISTANCE ALONG RAY TO CLOSEST APPROACH TO CENTER
C***** OF EARTH
UP=R(1)*R(4)
RADG=EARTH**2-R(1)**2*(R(5)**2+R(6)**2)
DISTG=SQRT (AMAX1(0.,RADG))
C***** DISTANCE ALONG RAY TO FIRST INTERSECTION WITH OR CLOSEST
C***** APPROACH TO THE EARTH
SG=-UP-DISTG
C***** CROSSG IS TRUE IF THE RAY WILL INTERSECT OR TOUCH THE EARTH
CROSSG=UP.LT.0..AND.RADG.GE.0.
RADR=(EARTH+RCVRH)**2-R(1)**2*(R(5)**2+R(6)**2)
DISTR=SQRT (AMAX1(0.,RADR))
C***** DISTANCE ALONG RAY TO THE FIRST INTERSECTION WITH OR CLOSEST
C***** APPROACH TO THE RECEIVER HEIGHT
SR=DISTR-UP
IF (UP.LT.0..AND.DISTR.LT.-UP.AND.R(1).NE.EARTH+RCVRH) SR=-DISTR
1 -UP
C***** CROSSR IS TRUE IF THE RAY WILL INTERSECT WITH OR MAKE A
C***** CLOSEST APPROACH TO THE RECEIVER HEIGHT
CROSSR=R(4).LT.0..OR.R(1).LT.(EARTH+RCVRH)
CROSS=CROSSG.OR.CROSSR
C***** MAXIMUM DISTANCE IN WHICH TO LOOK FOR THE IONOSPHERE
S1=AMIN1(SR,SG)
IF(.NOT.CROSSG) S1=SR
IF (UP.GE.0.) GO TO 15
CROSS=.TRUE.
C***** IF RAY IS GOING DOWN, S1 IS AT MOST THE DISTANCE TO A PERIGEE
S1=AMIN1(S1,-UP)
C***** CONVERT THE POSITION AND DIRECTION OF THE RAY TO CARTESIAN
C***** COORDINATES
15 CALL POL CAR
SSTEP=100.
S=SSTEP
DO 20 I=1,NSTEP
IF ((S-SSTEP).GT.S1.AND.CROSS) GO TO 25
C***** CONVERT POSITION AND DIRECTION TO SPHERICAL POLAR COORDINATES
C***** AT A DISTANCE S ALONG THE RAY
CALL CAR POL
CALL ELECTX
C***** FREE SPACE

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REAC001
REAC002
REAC003
REAC004
REAC005
REAC006
REAC007
REAC008
REAC009
REAC010
REAC011
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REAC 14
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REAC 55
REAC 56
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REAC 58
REAC 59
REAC 60
REAC 61
REAC 62
REAC 63
REAC 64
REAC 65

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IF (X.EQ.0.) GO TO 20 REAC 66
CALL RINDEX REAC 67
C***** EFFECTIVELY FREE SPACE REAC 68
IF (SPACE) GO TO 20 REAC 69
IF (SSTEP.LT.0.5E-4) GO TO 25 REAC 70
C***** RAY IN THE IONOSPHERE. STEP BACK OUT REAC 71
S=S-SSTEP REAC 72
C***** DECREASE STEP SIZE REAC 73
SSTEP=SSTEP/10. REAC 74
20 S=S+SSTEP REAC 75
PRINT 2000, NSTEP REAC 76
2000 FORMAT (9H EXCEEDED,I5,26H STEPS IN SUBROUTINE REACH) REAC 77
CALL EXIT REAC 78
25 IF(CROSS) S=AMIN1(S,S1) REAC 79
C***** CONVERT POSITION AND DIRECTION TO SPHERICAL POLAR COORDINATES REAC 80
C***** AT A DISTANCE S ALONG THE RAY REAC 81
CALL CAP POL REAC 82
C***** AVOID THE RAY BEING SLIGHTLY UNDERGROUND REAC 83
R(1)=AMAX1(R(1),EARTH) REAC 84
C***** ONE STEP INTEGRATION REAC 85
IF (N.LT.7) GO TO 31 REAC 86
DO 30 NN=7,N REAC 87
30 R(NN)=P(NN)+S*DROT(NN) REAC 88
31 T=T+S REAC 89
CALL RINDEX REAC 90
C***** AT A PERIGEE REAC 91
PERIGE=S.EQ.(-UP) REAC 92
C***** CORRECT MINOR ERRORS REAC 93
IF (PERIGE) R(4)=0. REAC 94
C***** KEEP CONSISTENCY AFTER CORRECTING MINOR ERRORS REAC 95
DROT(1)=R(4) REAC 96
C***** ON THE GROUND REAC 97
GROUND=S.EQ.SG.AND.CROSSG REAC 98
C***** AT THE RECEIVER HEIGHT REAC 99
THERE=S.EQ.SR.AND.CROSSR.AND..NOT.PERIGE REAC100
C***** AT A CLOSEST APPROACH TO THE RECEIVER HEIGHT REAC101
MINDIS=PERIGE.AND.S.EQ.SR.AND.CROSSR REAC102
RSPACE=SPACE REAC103
V=SQRT(N2/(R(4)**2+R(5)**2+R(6)**2)) REAC104
C***** RENORMALIZE THE WAVE NORMAL DIRECTION TO = SQRT(REAL(N**2)) REAC105
R(4)=R(4)*V REAC106
R(5)=R(5)*V REAC107
R(6)=R(6)*V REAC108
RSTART=1. REAC109
IF (.NOT.SPACE) CALL PRINTR (8HENTR ION,0.) REAC110
RETURN REAC111
END REAC112-

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SUBROUTINE POL CAR POLC001
DIMENSION X0(6),X(6),R0(4) POLC002
COMMON R(6) /COORD/ S POLC003
COMMON /CONST/ PI,PIT2,PID2,DUM(5) POLC004
C POLC005
C CONVERTS SPHERICAL COORDINATES TO CARTESIAN POLC006
IF (R(5).EQ.0..AND.R(6).EQ.0.) GO TO 1 POLC007
VERT=0. POLC008
SINA=SIN(R(2)) POLC009
COSA=SIN(PID2-R(2)) POLC010
SINP=SIN(R(3)) POLC011
COSP=SIN(PID2-R(3)) POLC012
X0(1)=R(1)*SINA*COSP POLC013
X0(2)=R(1)*SINA*SINP POLC014
X0(3)=R(1)*COSA POLC015

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X(4)=R(4)*SINA*COSP+R(5)*COSA*COSP-R(6)*SINP
X(5)=R(4)*SINA*SINP+R(5)*COSA*SINP+R(6)*COSP
X(6)=R(4)*COSA-R(5)*SINA
RETURN
C      VERTICAL INCIDENCE
1 VERT=1.
  RO(1)=R(1)
  RO(2)=R(2)
  RO(3)=R(3)
  RO(4)=SIGN (1.,R(4))
  RETURN
C
C      STEPS THE RAY A DISTANCE S, AND THEN
C      CONVERTS CARTESIAN COORDINATES TO SPHERICAL COORDINATES
ENTRY CAR POL
IF (VERT.NE.0.) GO TO 2
X(1)=X0(1)+S*X(4)
X(2)=X0(2)+S*X(5)
X(3)=X0(3)+S*X(6)
TEMP=SQRT(X(1)**2+X(2)**2)
R(1)=SQRT(X(1)**2+X(2)**2+X(3)**2)
R(2)=ATAN2(TEMP,X(3))
R(3)=ATAN2(X(2),X(1))
R(4)=(X(1)*X(4)+X(2)*X(5)+X(3)*X(6))/R(1)
R(5)=(X(3)*(X(1)*X(4)+X(2)*X(5))-(X(1)**2+X(2)**2)*X(6))/
1 (R(1)*TEMP)
R(6)=(X(1)*X(5)-X(2)*X(4))/TEMP
RETURN
C      VERTICAL INCIDENCE
2 R(1)=RO(1)+RO(4)*S
  R(2)=RO(2)
  R(3)=RO(3)
  R(4)=RO(4)
  R(5)=0.
  R(6)=0.
  RETURN
  END
SUBROUTINE PRINTR(NWHY,CARD)
C      PRINTS OUTPUT AND PUNCHES RAYSETS WHEN REQUESTED
DIMENSION G(3,3),G1(3,3),TYPE(3),HEADR1(20),HEADR2(20),UNITS(20),
1 HEAD1(20),HEAD2(20),UNIT(20),RPRINT(20),NPR(20)
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,DUM(3)
COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,N2,N2I,PNP(10),POLAR(2),
1 LPOLAR(2)
COMMON R(20),T /WW/ ID(10),W0,W(400)
EQUIVALENCE (THETA,R(2)),(PHI,R(3))
EQUIVALENCE (EARTH,R(4)),(XMTRH,W(3)),(TLAT,W(4)),(TLON,W(5)),
1 (F,W(6)),(AZ1,W(10)),(BETA,W(14)),(RCVRH,W(20)),(HOP,W(22)),
2 (PLAT,W(24)),(PLON,W(25)),(RAYSET,W(72))
LOGICAL SPACE,NEWWR,NEWWP,PENET
REAL N2,N2I,LPOLAR
COMPLEX PNP
DATA (TYPE=1HX,1HN,1HO)
2,(HEADR1(7)=6H PHAS),(HEADR2(7)=6H PATH),(UNITS(7)=6H KM),
3 (HEADR1(8)=6H ABSO),(HEADR2(8)=6H RPTION),(UNITS(8)=6H DB),
4 (HEADR1(9)=6H DOP),(HEADR2(9)=6H PLER),(UNITS(9)=6H C/S),
5 (HEADR1(10)=6H PATH),(HEADR2(10)=6H LENGTH),(UNITS(10)=6H KM)
CALL RINDEX
IF (.NOT.NEWWP) GO TO 10

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POLC016
POLC017
POLC018
POLC019
POLC020
POLC021
POLC022
POLC023
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POLC042
POLC043
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POLC048
POLC049
POLC050
POLC051
POLC 52-

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PRIN001
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PRIN015
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PRIN020
PRIN023
PRIN024
PRIN025

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C***** NEW M ARRAY -- REINITIALIZE
NEWWP=.FALSE.
SPL=SIN (PLON-TLON)
CPL=SIN (PID2-(PLON-TLON))
SP=SIN (PLAT)
CP=SIN (PID2-PLAT)
SL=SIN (TLAT)
CL=SIN (PID2-TLAT)
C***** MATRIX TO ROTATE COORDINATES
G(1,1)=CPL*SP*CL-CP*SL
G(1,2)=SPL*SP
G(1,3)=-SL*SP*CL-CL*CP
G(2,1)=-SPL*CL
G(2,2)=CPL
G(2,3)=SL*SP
G(3,1)=CL*CP*CPL+SP*SL
G(3,2)=CP*SP
G(3,3)=-SL*CP*CL+SP*CL
DENM=G(1,1)*G(2,2)*G(3,3)+G(1,2)*G(3,1)*G(2,3)+G(2,1)*G(3,2)*G(1,3)
1)-G(2,2)*G(3,1)*G(1,3)-G(1,2)*G(2,1)*G(3,3)-G(1,1)*G(3,2)*G(2,3)
C***** THE MATRIX G1 IS THE INVERSE OF THE MATRIX G
G1(1,1)=(G(2,2)*G(3,3)-G(3,2)*G(2,3))/DENM
G1(1,2)=(G(3,2)*G(1,3)-G(1,2)*G(3,3))/DENM
G1(1,3)=(G(1,2)*G(2,3)-G(2,2)*G(1,3))/DENM
G1(2,1)=(G(3,1)*G(2,3)-G(2,1)*G(3,3))/DENM
G1(2,2)=(G(1,1)*G(3,3)-G(3,1)*G(1,3))/DENM
G1(2,3)=(G(2,1)*G(1,3)-G(1,1)*G(2,3))/DENM
G1(3,1)=(G(2,1)*G(3,2)-G(3,1)*G(2,2))/DENM
G1(3,2)=(G(3,1)*G(1,2)-G(1,1)*G(3,2))/DENM
G1(3,3)=(G(1,1)*G(2,2)-G(2,1)*G(1,2))/DENM
R0=EARTH+XMTRH
C***** CARTESIAN COORDINATES OF TRANSMITTER
XR=R0*G(1,1)
YR=R0*G(2,1)
ZR=R0*G(3,1)
CTHR=G(3,1)
STHR=SIN (ACOS (GTHR))
PHIR=ATAN2 (YR,XR)
ALPH=ATAN2 (G(3,2),G(3,3))
C*****
NR=6
NP=0
DO 7 NN=7,20
IF (W(NN+50).EQ.0.) GO TO 7
C***** DEPENDENT VARIABLE NUMBER NN IS BEING INTEGRATED
C***** NR IS THE NUMBER OF DEPENDENT VARIABLES BEING INTEGRATED
NR=NR+1
IF (W(NN+50).NE.2.) GO TO 7
C***** DEPENDENT VARIABLE NUMBER NN IS BEING INTEGRATED AND PRINTED.
C***** NP IS THE NUMBER OF DEPENDENT VARIABLES BEING INTEGRATED AND
C***** PRINTED
NP=NP+1
C***** SAVE THE INDEX OF THE DEPENDENT VARIABLE TO PRINT
NPR(NP)=NR
HEAD1(NP)=HEADR1(NN)
HEAD2(NP)=HEADR2(NN)
UNIT(NP)=UNITS(NN)
7 CONTINUE
VP1=MIN0(NP,3)
PDEV=ABSORB=DOPP=0.

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PRIN026
PRIN027
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PRIN083
PRIN084
PRIN085

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C***** PRINT COLUMN HEADINGS AT THE BEGINNING OF EACH RAY
10 IF (IHOP.NE.0) GO TO 12
   PRINT 1100, (HEAD1(NN),HEAD2(NN),NN=1,NP1)
1100 FORMAT (44X,7HAZIMUTH/43X,9HDEVIATION,8X,9HELEVATION/
1 19X,16HHEIGHT RANGE,1X,2(5X,12HXMTR LOCAL),5X,26HPOLARIZATI
20N GROUP PATH,5A6,A5)
   PRINT 1150, (JNIT(NN),NN=1,NP1)
1150 FORMAT (13X,2(8X,2HKM),2X,2(6X,3HDEG,5X,3HDEG),6X,12HREAL IMAG,
1 7X,2HKM,4X,3(4X,A6,2X))
   IF (RAYSET.EQ.0.) GO TO 12
C***** PUNCH A TRANSMITTER RAYSET
   TLOND=TLON*DEGS
   IF (TLOND.LT.0.) TLOND=TLOND+360.
   TLATD=TLAT*DEGS
   IF (TLATD.LT.0.) TLATD=TLATD+360.
   AZ=AZ1*DEGS
   EL=BETA*DEGS
   NHOP=HOP
   PUNCH 1200, ID(1),TYPE(NTYP),XMTRH,TLATD,TLOND,RCVRH,F,AZ,EL,POLAR
1,NHOP,1HT
1200 FORMAT (A3,A1,4PF9.0,3P2F6.0,4P2F9.0,5P2F10.0,5X,2P2F5.0,I1,A1)
C*****
12 V=0.
   IF (N2.NE.0.) V=(R(4)**2+R(5)**2+R(6)**2)/N2-1.
   H=R(1)-EARTH
   STH=SIN (THETA)
   CTH=SIN (PID2-THETA)
C***** CARTESIAN COORDINATES OF RAY POINT, ORIGIN AT TRANSMITTER
   XP=R(1)*STH*SIN (PID2-PHI)-XR
   YP=R(1)*STH*SIN (PHI)-YR
   ZP=R(1)*CTH-ZR
C***** CARTESIAN COORDINATES OF RAY POINT, ORIGIN AT TRANSMITTER AND
C***** ROTATED
   EPS=XP*G1(1,1)+YP*G1(1,2)+ZP*G1(1,3)
   ETA=XP*G1(2,1)+YP*G1(2,2)+ZP*G1(2,3)
   ZETA=XP*G1(3,1)+YP*G1(3,2)+ZP*G1(3,3)
   RCE2=ETA**2+ZETA**2
   RCE=SQRT (RCE2)
C***** GROUND RANGE
   RANGE=EARTH*ATAN2(RCE,EARTH+EPS)
C***** ANGLE OF WAVE NORMAL WITH LOCAL HORIZONTAL
   ELL=ATAN2(R(4),SQRT (R(5)**2+R(6)**2))*DEGS
C***** STRAIGHT LINE DISTANCE FROM TRANSMITTER TO RAY POINT
   SR=SQRT (RCE2+EPS**2)
   IF (NP.LT.1) GO TO 16
   DO 15 I=1,NP
   NN=NPR(I)
15 RPRINT(I)=R(NN)
16 IF (SR.GE.1.E-6) GO TO 20
C***** TOO CLOSE TO TRANSMITTER TO CALCULATE DIRECTION FROM
C***** TRANSMITTER
   PRINT 1500, V,NWHY,H,RANGE,ELL,POLAR,T,(RPRINT(NN),NN=1,NP1)
1500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,26X,F8.3,F9.3,F8.3,4F12.4)
   GO TO 40
C***** ELEVATION ANGLE OF RAY POINT FROM TRANSMITTER
20 EL=ATAN2(EPS,RCE)*DEGS
   IF (RCE.GE.1.E-6) GO TO 30
C***** NEARLY DIRECTLY ABOVE OR BELOW TRANSMITTER. CAN NOT CALCULATE
C***** AZIMUTH DIRECTION FROM TRANSMITTER ACCURATELY
   PRINT 2500, V,NWHY,H,RANGE,EL,ELL,POLAR,T,(RPRINT(NN),NN=1,NP1)
2500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,17X,F9.3,F8.3,F9.3,F8.3,
1 4F12.4)
   GO TO 40

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C***** AZIMUTH ANGLE OF RAY POINT FROM TRANSMITTER
30 ANGA=ATAN2(ETA,ZETA)
AZDEV=180.-AMOD(540.-(AZI-ANGA)*DEGS,360.)
IF (R(5).NE.0..OR.R(6).NE.0.) GO TO 34
C***** WAVE NORMAL IS VERTICAL, SO AZIMUTH DIRECTION CANNOT BE
C***** CALCULATED
PRINT 3000, V,NWHY,H,RANGE,AZDEV,EL,ELL,POLAR,T,(RPRINT(NN),NN=1,
1 NP1)
3000 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,F9.3,8X,F9.3,F8.3,F9.3,F8.3,
1 4F12.4)
GO TO 40
34 ANA=ANGA-ALPH
SANA=SIN(ANA)
SPHI=SANA*STHR/STH
CPHI=-SIN(PID2-ANA)*SIN(PID2-(PHI-PHIR))+SANA*SIN(PHI-PHIR)
1 *CTHR
AZA=180.-AMOD(540.-(ATAN2(SPHI,CPHI)-ATAN2(R(6),R(5)))*DEGS,360.)
PRINT 3500, V,NWHY,H,RANGE,AZDEV,AZA,EL,ELL,POLAR,T,(RPRINT(NN),NN=
1 =1,NP1)
3500 FORMAT (1X,E6.0,1X,A8,F10.4,F11.4,2(F9.3,F8.3),F9.3,F8.3,
1 4F12.4)
C*****
40 LINES=LINES+1
IF (NP.LE.3) GO TO 45
C***** ADDITIONAL LINE TO PRINT REMAINING DEPENDENT INTEGRATION
C***** VARIABLES
PRINT 4000, (RPRINT(NN),NN=4,NP)
4000 FORMAT (99X,3F12.4)
LINES=LINES+1
45 IF (CARD.EQ.0.) RETURN
C
C***** PUNCH A RAYSET
IF (AZDEV.LT.-90.) AZDEV=AZDEV+360.
IF (AZA.LT.-90.) AZA=AZA+360.
TDEV=T-SR
NR=6
IF (W(57).EQ.0.) GO TO 47
C***** PHASE PATH
NR=NR+1
PDEV=R(NR)-SR
47 IF (W(58).EQ.0.) GO TO 48
C***** ABSORPTION
NR=NR+1
ABSORB=R(NR)
C***** DOPPLER SHIFT
48 IF (W(59).NE.0.) DOPP=R(NR+1)
PUNCH 4500, HPUNCH,RANGE,AZDEV,AZA,ELL,SR,TDEV,PDEV,ABSORB,DOPP,
1 POLAR,IHOP,NWHY
4500 FORMAT (4P2F9.0,3P3F6.0,3PF8.0,3P4F6.0,2P2F5.0,I1,A1)
RETURN
END

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PRIN149
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PRIN199-

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INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON A VERTICAL PLANE

Coordinates of the left edge of the graph:

Latitude = \_\_\_\_\_ rad  
deg north (W83)  
km

Longitude = \_\_\_\_\_ rad  
deg east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = \_\_\_\_\_ rad  
deg north (W85)  
km

Longitude = \_\_\_\_\_ rad  
deg east (W86)  
km

Height above the ground of the bottom of the graph = \_\_\_\_\_ km (W88)

Distance between tic marks = \_\_\_\_\_ rad  
deg (W87)  
km

(W81 = 1.)

INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON THE GROUND

Coordinates of the left edge of the graph:

Latitude = \_\_\_\_\_ rad  
deg north (W83)  
km

Longitude = \_\_\_\_\_ rad  
deg east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = \_\_\_\_\_ rad  
deg north (W85)  
km

Longitude = \_\_\_\_\_ rad  
deg east (W86)  
km

Factor to expand lateral deviation scale by = \_\_\_\_\_ (W82)

Distance between tic marks on range scale = \_\_\_\_\_ rad  
deg (W87)  
km

(W81 = 2.)

	SUBROUTINE RAYPLT	YPLT001
C	REPLACES SUBROUTINES RAYPLT,PLOT, AND LABPLT IF PLOTS ARE	YPLT002
C	NOT WANTED OR IF A PLOTTER IS NOT AVAILABLE	YPLT003
	COMMON /WW/ ID(10),W0,W(400)	YPLT004
	EQUIVALENCE (PLT,W(81))	YPLT005
	PLT=0.	YPLT006
	ENTRY ENDPLT	YPLT007
	RETURN	YPLT008
	END	YPLT 9-
	SUBROUTINE RAYPLT	RAYP001
C	W(81)=1. PLOTS PROJECTION OF RAYPATH ON VERTICAL PLANE	RAYP002
C	=2. PLOTS PROJECTION OF RAYPATH ON GROUND	RAYP003
	COMMON /PLT/ XL,XR,YB,YT,RESET	RAYP004
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	RAYP005
	COMMON /FLG/ NTYP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH	RAYP006
	COMMON R(6) /WW/ ID(10),W0,W(400)	RAYP007
	EQUIVALENCE (TH,R(2)),(PH,R(3))	RAYP008
	EQUIVALENCE (EARTH,R(2)),(PLAT,W(24)),(PLON,W(25)),(PLT,W(81)),	RAYP009
	1 (FACTR,W(82)),(LLAT,W(83)),(LLON,W(84)),(RLAT,W(85)),(RLON,W(86)),	RAYP010
	2,(TIC,W(87)),(HB,W(88))	RAYP011
	REAL LLAT,LLON,LTIC	RAYP012
	LOGICAL NEWWR,NEWWP,PENET	RAYP013
	IF (.NOT.NEWWR) GO TO 5	RAYP014
C		RAYP015
C	NEW W ARRAY -- REINITIALIZE	RAYP016
	NEWWR=.FALSE.	RAYP017
	RESET=1.	RAYP018
	CONVERT COORDINATES OF VERTICAL PLANE FROM GEOGRAPHIC TO GEOMAGNETIC	RAYP019
	SW=SIN (PLAT)	RAYP020
	CW=SIN (PID2-PLAT)	RAYP021
	SLM=SIN (LLAT)	RAYP022
	CLM=SIN (PID2-LLAT)	RAYP023
	SRM=SIN (RLAT)	RAYP024
	CRM=SIN (PID2-RLAT)	RAYP025
	CDPHI=SIN (PID2-(LLON-PLON))	RAYP026
	PHL=ATAN2(SIN (LLON-PLON)*CLM,CDPHI*SW*CLM-CW*SLM)	RAYP027
	CTHL=CDPHI*CW*CLM+SW*SLM	RAYP028
	STHL=SIN (ACOS (CTHL))	RAYP029
	CDPHI=SIN (PID2-(RLON-PLON))	RAYP030
	PHR=ATAN2(SIN (RLON-PLON)*CRM,CDPHI*SW*CRM-CW*SRM)	RAYP031
	CTHR=CDPHI*CW*CRM+SW*SRM	RAYP032
	STHR=SIN (ACOS (CTHR))	RAYP033
	CLR=CTHL*CTHR+STHL*STHR*SIN (PID2-(PHL-PHR))	RAYP034
	SLR=SQRT (1.-CLR**2)	RAYP035
	IF (PLT.EQ.2.) GO TO 3	RAYP036
	FACTR=1.	RAYP037
	R0=EARTH+HB	RAYP038
	ALPHA=.5*ACOS (CLR)	RAYP039
	XR=R0*SIN (ALPHA)	RAYP040
	XL=-XR	RAYP041
	YB=R0*SIN (PID2-ALPHA)	RAYP042
	YT=YB+2.*XR	RAYP043
	GO TO 5	RAYP044
	3 IF (FACTR.EQ.0.) FACTR=1.	RAYP045
	ALPH1=ATAN2(STHR*SIN (PHR-PHL),(CTHR-CTHL*CLR)/STHL)	RAYP046
	XL=0.	RAYP047
	XR=EARTH*ACOS (CLR)	RAYP048
	YT=0.5*XR/FACTR	RAYP049
	YB=-YT	RAYP050

C	5	STH=SIN (TH)	RAYP051
		CTH=SIN (PID2-TH)	RAYP052
		CR=CTHR*CTH+STHR*STH*SIN (PID2-(PHR-PH))	RAYP053
		CL=CTHL*CTH+STHL*STH*SIN (PID2-(PHL-PH))	RAYP054
		CEA=ATAN2(CR-CL*CLR,CL*SLR)	RAYP055
		NEW=1	RAYP056
		IF (IHOP.NE.0) NEW=0	RAYP057
		IF (PLT.EQ.2.) GO TO 10	RAYP058
		CALL PLOT (R(1)*SIN(CEA-ALPHA),R(1)*SIN(PID2-(CEA-ALPHA)),NEW)	RAYP059
		RETURN	RAYP060
	10	SL=SQRT (1.-CL**2)	RAYP061
		TMP1=STH*SIN (PH-PHL)	RAYP062
		TMP2=(CTH-CTHL*CL)/STHL	RAYP063
		ALPH2=0.	RAYP064
		IF (TMP1.NE.0..OR.TMP2.NE.0.) ALPH2=ATAN2(TMP1,TMP2)	RAYP065
		CALL PLOT (EARTH*CEA,EARTH*ASIN(SL*SIN (ALPH1-ALPH2)),NEW)	RAYP066
		RETURN	RAYP067
C			RAYP068
C		DRAW AXES AND CALL FOR LABELING AND TERMINATION OF THIS PLOT	RAYP069
		ENTRY ENDPLT	RAYP070
		TICKX=0.01*(YT-YB)	RAYP071
		IF (PLT.EQ.2.) GO TO 25	RAYP072
		R1=EARTH-TICKX	RAYP073
		X=XL	RAYP074
		Y=YB	RAYP075
		CALL PLOT (X,Y,1)	RAYP076
		NTIC=2	RAYP077
		IF (TIC.NE.0.) NTIC=NTIC+2.*ALPHA/TIC	RAYP078
		NLINE=MAX0 (1,100/NTIC)	RAYP079
		DO 20 I=1,NTIC	RAYP080
		ANG=-ALPHA+(I-1)*TIC	RAYP081
		CALL PLOT (R1*SIN (ANG),R1*SIN (PID2-ANG),0)	RAYP082
		CALL PLOT (X,Y,0)	RAYP083
		DO 20 J=1,NLINE	RAYP084
		ANG=ANG+TIC/NLINE	RAYP085
		X=EARTH*SIN (ANG)	RAYP086
		Y=EARTH*SIN (PID2-ANG)	RAYP087
	20	CALL PLOT (X,Y,0)	RAYP088
		CALL PLOT (XR,YB,0)	RAYP089
		GO TO 50	RAYP090
	25	DTIC=TIC*EARTH	RAYP091
		LTIC=DTIC/FACTR	RAYP092
		TICY=XL+0.01*(XR-XL)	RAYP093
		NTIC=YT/LTIC	RAYP094
		TIC1=-LTIC*NTIC	RAYP095
		CALL PLOT (XL,YB,1)	RAYP096
		NTIC=2*NTIC+1	RAYP097
		DO 30 I=1,NTIC	RAYP098
		Y=TIC1+(I-1)*LTIC	RAYP099
		CALL PLOT (XL,Y,0)	RAYP100
		CALL PLOT (TICY,Y,0)	RAYP101
	30	CALL PLOT (XL,Y,0)	RAYP102
		CALL PLOT (XL,YT,0)	RAYP103
		CALL PLOT (XL,0.,1)	RAYP104
		NTIC=(XR-XL)/DTIC	RAYP105
		DO 40 I=1,NTIC	RAYP106
		X=I*DTIC	RAYP107
		CALL PLOT (X,0.,0)	RAYP108
		CALL PLOT (X,TICKX,0)	RAYP109
	40	CALL PLOT (X,0.,0)	RAYP110
		CALL PLOT (XR,0.,0)	RAYP111
	50	CALL LABPLT	RAYP112
		CALL PLTEND	RAYP113
		RETURN	RAYP114
		END	RAYP115
			RAYP116-

	SUBROUTINE PLOT (X,Y,NEW)	PLOT001
	COMMON /PLT/ XMIN0,XMAX0,YMIN0,YMAX0,RESET	PLOT002
	COMMON /DD/ INT,IOR,IT,IS,IC,ICC,IX,IY	PLOT003
	DATA (INITAL=1),(MINX=0),(MINY=0),(MAXX=1023),(MAXY=1023),	PLOT004
	1 (MINX0=23),(MINY0=23),(MAXX0=1023),(MAXY0=1023)	PLOT005
C		PLOT006
C	INITIALIZE LIBRARY PLOTTING ROUTINES	PLOT007
	IF (INITAL.EQ.0) GO TO 1	PLOT008
	INITAL=0	PLOT009
	CALL DDINIT (1,1H )	PLOT010
C		PLOT011
C	COMPUTE SCALE FACTORS	PLOT012
	1 IF (RESET.EQ.0.) GO TO 5	PLOT013
	RESET=0.	PLOT014
	XSCALE=(MAXX0-MINX0)/(XMAX0-XMIN0)	PLOT015
	YSCALE=(MAXY0-MINY0)/(YMAX0-YMIN0)	PLOT016
	XMIN=XMIN0-(MINX0-MINX)/XSCALE	PLOT017
	YMIN=YMIN0-(MINY0-MINY)/YSCALE	PLOT018
	XMAX=XMAX0+(MAXX-MAXX0)/XSCALE	PLOT019
	YMAX=YMAX0+(MAXY-MAXY0)/YSCALE	PLOT020
C		PLOT021
C	START A NEW LINE	PLOT022
	5 IF (NEW.EQ.0) GO TO 10	PLOT023
	IX=MINX0+(X-XMIN0)*XSCALE	PLOT024
	IY=MINY0+(Y-YMIN0)*YSCALE	PLOT025
	IF (IX.GE.MINX.AND.IX.LE.MAXX.AND.IY.GE.MINY.AND.IY.LE.MAXY)	PLOT026
	1 CALL DDBP	PLOT027
	GO TO 50	PLOT028
C		PLOT029
C	HORIZONTAL DISPLACEMENT	PLOT030
	10 XS=X-XOLD	PLOT031
	YS=Y-YOLD	PLOT032
	IF (XS) 11,12,16	PLOT033
C	NEGATIVE	PLOT034
	11 X1=XMAX	PLOT035
	X2=XMIN	PLOT036
	GO TO 20	PLOT037
C	ZERO	PLOT038
	12 IF (YS) 13,50,14	PLOT039
	13 S1=(YMAX-YOLD)/YS	PLOT040
	S2=(YMIN-YOLD)/YS	PLOT041
	GO TO 40	PLOT042
	14 S1=(YMIN-YOLD)/YS	PLOT043
	S2=(YMAX-YOLD)/YS	PLOT044
	GO TO 40	PLOT045
C	POSITIVE	PLOT046
	16 X1=XMIN	PLOT047
	X2=XMAX	PLOT048
C		PLOT049
C	VERTICAL DISPLACEMENT	PLOT050
	20 IF (YS) 21,22,26	PLOT051
C	NEGATIVE	PLOT052
	21 Y1=YMAX	PLOT053
	Y2=YMIN	PLOT054
	GO TO 30	PLOT055
C	ZERO	PLOT056
	22 S1=(X1-XOLD)/XS	PLOT057
	S2=(X2-XOLD)/XS	PLOT058
	GO TO 40	PLOT059
C	POSITIVE	PLOT060
	26 Y1=YMIN	PLOT061
	Y2=YMAX	PLOT062
C		PLOT063
	30 S1=AMAX1((X1-XOLD)/XS,(Y1-YOLD)/YS)	PLOT064
	S2=AMIN1((X2-XOLD)/XS,(Y2-YOLD)/YS)	PLOT065

C		PLOT LINE -- CHECKING FOR BORDER CROSSINGS	PLOT066
C			PLOT067
	40	S=SQRT(XS**2+YS**2)	PLOT068
		IF (S2.LT.0..OR.S*S1-S.GT.0.) GO TO 50	PLOT069
		IF (S1.LT.0.) GO TO 42	PLOT070
C		PREVIOUS POINT OFF GRAPH	PLOT071
		IX=MINX0+(XOLD+XS*S1-XMIN0)*XSCALE+0.5	PLOT072
		IY=MINY0+(YOLD+YS*S1-YMIN0)*YSCALE+0.5	PLOT073
		CALL DDBP	PLOT074
	42	IF (S*S2-S.GT.0.) GO TO 44	PLOT075
C		CURRENT POINT OFF GRAPH	PLOT076
		IX=MINX0+(XOLD+XS*S2-XMIN0)*XSCALE+0.5	PLOT077
		IY=MINY0+(YOLD+YS*S2-YMIN0)*YSCALE+0.5	PLOT078
		CALL DDVC	PLOT079
		GO TO 50	PLOT080
C		CURRENT POINT ON GRAPH	PLOT081
	44	IX=MINX0+(X-XMIN0)*XSCALE+0.5	PLOT082
		IY=MINY0+(Y-YMIN0)*YSCALE+0.5	PLOT083
		CALL DDVC	PLOT084
C			PLOT085
C		EXIT ROUTINE	PLOT086
	50	XOLD=X	PLOT087
		YOLD=Y	PLOT088
		RETURN	PLOT089
C			PLOT090
C		TERMINATE THE CURRENT PLOT	PLOT091
		ENTRY PLTEND	PLOT092
		CALL DDFR	PLOT093
		RETURN	PLOT094
		END	PLOT 95-

		SUBROUTINE LABPLT	LABP001
C		LABEL THE CURRENT PLOT	LABP002
		DIMENSION LABEL(9),TYPE(3)	LABP003
		COMMON /OD/ INT,IOR,IT,IS,IC,ICC,IX,IY	LABP004
		COMMON /CONST/ PI,PIT2,PID2,DEGS,DUM(4)	LABP005
		COMMON /FLG/ NTYP,NEWWP,NEWWP,PENET,LINES,IHOP,HPUNCH	LABP006
		COMMON /WW/ ID(10),W0,W(400)	LABP007
		EQUIVALENCE (EARTH,W(2)),(F,W(6)),(AZ1,W(10)),(PLT,W(81)),	LABP008
	1	(FACTR,W(82)),(TIC,W(87))	LABP009
		LOGICAL NEWWR,NEWWP,PENET	LABP010
		REAL LTIC	LABP011
		DATA (TYPE=8HEXTRAORD,8HNO FIELD,8HORDINARY)	LABP012
		IOR=IT=0	LABP013
		IS=2	LABP014
		IX=0 \$ IY=1023 \$ CALL ODTAB \$ CALL DDTXT (7,ID)	LABP015
		NDATE=IDATE(0)	LABP016
		CALL DDTXT (1,NDATE)	LABP017
		AZA=AZ1*DEGS	LABP018
		DTIC=TIC*EARTH	LABP019
		ENCODE (72,1000,LABEL) F,AZA,TYPE(NTYP),DTIC	LABP020
	1000	FORMAT (3HF =,F7.3,6H, AZ =,F7.2,2H, ,A8,2H, ,F7.2,24H KM BETWEEN	LABP021
		1TICK MARKS#.)	LABP022
		IX=0 \$ IY=991 \$ CALL ODTAB \$ CALL DDTXT (9,LABEL)	LABP023
		IF (PLT.EQ.1.) RETURN	LABP024
		LTIC=DTIC/FACTR	LABP025
		ENCODE (32,2000,LABEL) LTIC	LABP026
	2000	FORMAT (F7.2,24H KM BETWEEN TICK MARKS#.)	LABP027
		IOR=1	LABP028
		IX=0 \$ IY=0 \$ CALL ODTAB \$ CALL DDTXT (4,LABEL)	LABP029
		IOR=0	LABP030
		RETURN	LABP031
		END	LABP 32-



<pre> SUBROUTINE RKAM C   NUMERICAL INTEGRATION OF DIFFERENTIAL EQUATIONS COMMON /RK/ NN,SPACE,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART COMMON Y(20),T,STEP,DYDT(20) DIMENSION DELY(4,20),BET(4),XV(5),FV(4,20),YU(5,20) TYPE DOJBLE YU IF (RSTART.EQ.0.) GO TO 1000 LL=MM=1 IF (MODE.EQ.1) MM=4 ALPHA=T EPM=0.0 BET(1)=BET(2)=0.5 BET(3)=1.0 BET(4)=0.0 STEP=SPACE R=19.0/270.0 XV(MM)=T IF (E1MIN.LE.0.) E1MIN=E1MAX/55. IF (FACT.LE.0.) FACT=0.5 CALL HAMLTN DO 320 I=1,NN FV(MM,I)=DYDT(I) 320 YU(MM,I)=Y(I) RSTART=0. GO TO 1001 1000 IF (MODE.NE.1) GO TO 2000 C C   RUNGE-KUTTA 1031 DO 1034 K=1,4 DO 1350 I=1,NN DELY(K,I)=STEP*FV(MM,I) Z=YU(MM,I) 1350 Y(I)=Z+BET(K)*DELY(K,I) T=BET(K)*STEP+XV(MM) CALL HAMLTN DO 1034 I=1,NN 1034 FV(MM,I)=DYDT(I) DO 1039 I=1,NN DEL=(DELY(1,I)+2.0*DELY(2,I)+2.0*DELY(3,I)+DELY(4,I))/6.0 1039 YU(MM+1,I)=YU(MM,I)+DEL MM=MM+1 XV(MM)=XV(MM-1)+STEP DO 1400 I=1,NN 1400 Y(I)=YU(MM,I) T=XV(MM) CALL HAMLTN IF (MODE.EQ.1) GO TO 42 DO 150 I=1,NN 150 FV(MM,I)=DYDT(I) IF (MM.LE.3) GO TO 1001 C C   ADAMS-MOULTON 2000 DO 2048 I=1,NN DEL=STEP*(55.*FV(4,I)-59.*FV(3,I)+37.*FV(2,I)-9.*FV(1,I))/24. Y(I)=YU(4,I)+DEL 2048 DELY(1,I)=Y(I) T=XV(4)+STEP CALL HAMLTN XV(5)=T DO 2051 I=1,NN DEL=STEP*(9.*DYDT(I)+19.*FV(4,I)-5.*FV(3,I)+FV(2,I))/24. YU(5,I)=YU(4,I)+DEL 2051 Y(I)=YU(5,I) CALL HAMLTN IF (MODE.LE.2) GO TO 42 </pre>	<pre> RKAM001 RKAM002 RKAM003 RKAM004 RKAM005 RKAM006 RKAM007 RKAM008 RKAM009 RKAM010 RKAM011 RKAM012 RKAM013 RKAM014 RKAM015 RKAM016 RKAM017 RKAM018 RKAM019 RKAM020 RKAM021 RKAM022 RKAM023 RKAM024 RKAM025 RKAM026 RKAM027 RKAM028 RKAM029 RKAM030 RKAM031 RKAM032 RKAM033 RKAM034 RKAM035 RKAM036 RKAM037 RKAM038 RKAM039 RKAM040 RKAM041 RKAM042 RKAM043 RKAM044 RKAM045 RKAM046 RKAM047 RKAM048 RKAM049 RKAM050 RKAM051 RKAM052 RKAM053 RKAM054 RKAM055 RKAM056 RKAM057 RKAM058 RKAM059 RKAM060 RKAM061 RKAM062 RKAM063 RKAM064 RKAM065 </pre>
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C		RKAM066
C	ERROR ANALYSIS	RKAM067
	SSE=0.0	RKAM068
	DO 3033 I=1,NN	RKAM069
	EPSIL=R*ABS(Y(I)-DELY(1,I))	RKAM070
	IF (MODE.EQ.3.AND.Y(I).NE.0.) EPSIL=EPSIL/ABS(Y(I))	RKAM071
	IF (SSE.LT.EPSIL) SSE=EPSIL	RKAM072
3033	CONTINUE	RKAM073
	IF (E1MAX.GT.SSE) GO TO 3035	RKAM074
	IF (ABS(STEP).LE.E2MIN) GO TO 42	RKAM075
	LL=MM=1	RKAM076
	STEP=STEP*FACT	RKAM077
	GO TO 1001	RKAM078
3035	IF (LL.LE.1.OR.SSE.GE.E1MIN.OR.E2MAX.LE.ABS(STEP)) GO TO 42	RKAM079
	LL=2	RKAM080
	MM=3	RKAM081
	XV(2)=XV(3)	RKAM082
	XV(3)=XV(5)	RKAM083
	DO 5363 I=1,NN	RKAM084
	FV(2,I)=FV(3,I)	RKAM085
	FV(3,I)=DYDT(I)	RKAM086
	YU(2,I)=YJ(3,I)	RKAM087
5363	YU(3,I)=YJ(5,I)	RKAM088
	STEP=2.0*STEP	RKAM089
	GO TO 1001	RKAM090
C		RKAM091
C	EXIT ROUTINE	RKAM092
42	LL=2	RKAM093
	MM=4	RKAM094
	DO 12 K=1,3	RKAM095
	XV(K)=XV(K+1)	RKAM096
	DO 12 I=1,NN	RKAM097
	FV(K,I)=FV(K+1,I)	RKAM098
12	YU(K,I)=YU(K+1,I)	RKAM099
	XV(4)=XV(5)	RKAM100
	DO 52 I=1,NN	RKAM101
	FV(4,I)=DYDT(I)	RKAM102
52	YU(4,I)=YU(5,I)	RKAM103
	IF (MODE.LE.2) RETURN	RKAM104
	E=ABS(XV(4)-ALPHA)	RKAM105
	IF (E.LE.EPM) GO TO 2000	RKAM106
	EPM=E	RKAM107
	RETURN	RKAM108
	END	RKAM109-

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SUBROUTINE HAMLTN
C***** CALCULATES HAMILTONS EQUATIONS FOR RAY TRACING
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,
1 H,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPTHI,PHPPH,PHPPHI
2 ,PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI, PHPKPH,PHPKPI
3 ,KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI
COMMON R(20),T,STP,DRDT(20) /WW/ IO(10),W0,W(400)
EQUIVALENCE (TH,R(2)),(PH,R(3)),(KR,R(4)),(KTH,R(5)),(KPH,R(6)),
1 (DTHDT,DRDT(2)),(DPHDT,DRDT(3)),(DKRDT,DRDT(4)),(DKTHDT,DRDT(5)),
2 (DKPHDT,DRDT(6)),(F,W(6))
REAL KR,KTH,KPH,KPHPK,KPHPKI,LPOLAR,LPOLRI,LOGTEN,K,KAY2,KAY2I
OM=PIT2*1.E6*F
STH=SIN(TH)
CTH=SIN(PID2-TH)
RSTH=R(1)*STH
RCTH=R(1)*CTH
CALL RINDEX
DRDT=-PHPKR/(PHPOM*C)
DTHDT=-PH2KTH/(PHPOM*R(1)*C)
DPHDT=-PHPKPH/(PHPOM*RSTH*C)
DKRDT=PHPR/(PHPOM*C)+KTH*DTHDT+KPH*STH*DPHDT
DKTHDT=(PHPTH/(PHPOM*C)-KTH*DRDT+KPH*RCTH*DPHDT)/R(1)
DKPHDT=(PHPPH/(PHPOM*C)-KPH*STH*DRDT-KPH*RCTH*DTHDT)/RSTH
NR=6
C***** PHASE PATH
IF (W(57).EQ.0.) GO TO 10
NR=NR+1
DRDT(NR)=- KPHPK/PHPOM/OM
C***** ABSORPTION
10 IF (W(58).EQ.0.) GO TO 15
NR=NR+1
DRDT(NR)= 10./LOGTEN*KPHPK*KAY2I/(KR*KR+KTH*KTH+KPH*KPH)/PHPOM/C
C***** DOPPLER SHIFT
15 IF (W(59).EQ.0.) GO TO 20
NR=NR+1
DRDT(NR)=-PHPT/PHPOM/C/PIT2
C***** GEOMETRICAL PATH LENGTH
20 IF (W(60).EQ.0.) GO TO 25
NR=NR+1
DRDT(NR)=-SQRT(PHPKR**2+PHPKTH**2+PHPKPH**2)/PHPOM /C
C***** OTHER CALCULATIONS
25 CONTINUE
RETURN
END
HAML001
HAML002
HAML003
HAML004
HAML005
HAML006
HAML007
HAML008
HAML009
HAML010
HAML011
HAML012
HAML013
HAML014
HAML015
HAML016
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HAML041
HAML042
HAML043
HAML044
HAML045-

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## APPENDIX 2. VERSIONS OF THE REFRACTIVE INDEX SUBROUTINE (RINDEX)

This ray tracing program gains versatility without sacrificing speed by having several versions of some of the subroutines. For example, the 8 versions of the refractive index subroutine allow the user to decide for each ray path calculation whether to include or ignore various aspects of the propagation medium such as the earth's magnetic field or collisions between electrons and neutral air molecules.

If collisions are included, the user has the option of using the Appleton-Hartree formula (which assumes a constant collision frequency) or the Sen-Wyller formula (which assumes a Maxwell distribution of electron energies and a collision frequency proportional to energy). The Sen-Wyller formula is generally assumed to be more accurate, especially in the lower ionosphere, but the Appleton-Hartree formula can often be used with an effective collision frequency profile to save computer time.

When the effect of the earth's magnetic field is included and ray paths are calculated near vertical incidence, a spitze (Davies, 1965, p. 202) often occurs in the ray path. (At a spitze, the usual formulas for refractive index become indeterminate because the wave normal is parallel with the earth's magnetic field and the wave frequency equals the local plasma frequency.) Two versions of the refractive index subroutine have been developed to calculate ray paths through a spitze. These two versions will also work in the absence of a spitze, but the standard versions are much faster.

The input to the refractive index subroutines is through blank common and common blocks /XX/, /YY/, and /ZZ/. Output is through common block /RIN/. The refractive index subroutine is called through the entry RINDEX. The subroutine names are used only for user identification. The following 8 versions of the refractive index subroutine are

listed in this appendix:

a.	Subroutine AHFWFC (Appleton-Hartree formula with field, with collisions)	93
b.	Subroutine AHWFNC (Appleton-Hartree formula with field, no collisions)	94
c.	Subroutine AHNFWC (Appleton-Hartree formula no field, with collisions)	96
d.	Subroutine AHNFNC (Appleton-Hartree formula no field, no collisions)	97
e.	Subroutine BQFWFC (Booker Quartic with field, with collisions)	98
f.	Subroutine BQWFNC (Booker Quartic with field, no collisions)	100
g.	Subroutine SWWF (Sen-Wyller formula with field)	102
h.	Subroutine SWNF (Sen-Wyller no field)	105
	Subroutine FGSW	106
	Subroutine FSW	106
	Fresnel integral function C	108
	Fresnel integral function S	108

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SUBROUTINE A4WFC                                WFWC001
CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE WFWC002
APPLETON-HARTREE FORMULA WITH FIELD, WITH COLLISIONS WFWC003
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN WFWC004
COMMON /RIN/ MODRIN(3),CJ_L,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH, WFWC005
1 PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR WFWC006
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX WFWC007
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR WFWC008
1 ,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP WFWC009
COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH WFWC010
COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400) WFWC011
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART WFWC012
EQUIVALENCE (RAY,W(1)),(F,W(6)) WFWC013
LOGICAL SPACE WFWC014
REAL KR,KTH,KPH,K2 WFWC015
COMPLEX N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT, WFWC016
1 POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2,D2, WFWC017
2 KAY2,H,PHPT,PHPR,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH, WFWC018
3 KPHPK WFWC019
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=1.), WFWC020
1 (FIELD=1.), WFWC021
2 (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.), WFWC022
3 (Y=0.),(PYPR=0.),(PYPTH=0.),(PYPPH=0.),(YR=0.),(PYRPR=0.), WFWC023
4 (PYRPT=0.),(PYRPP=0.),(YTH=0.),(PYTPR=0.),(PYTPT=0.), WFWC024
5 (PYTPP=0.),(YPH=0.),(PYPPR=0.),(PYPPT=0.),(PYPPP=0.) WFWC025
6 ,(Z=0.),(PZPR=0.),(PZPTH=0.),(PZPPH=0.), WFWC026
7 (I=(0.,1.)),(ABSLIM=1.E-5) WFWC027
ENTRY RINDEX WFWC028
OM=PI*PIT2*1.E6*F WFWC029
C2=C*C WFWC030
K2=KR*KR+KTH*KTH+KPH*KPH WFWC031
OM2=OM*OM WFWC032
VR =C/OM*KR WFWC033
VTH=C/OM*KTH WFWC034
VPH=C/OM*KPH WFWC035
CALL ELECTX WFWC036
CALL MAGY WFWC037
V2=VR**2+VTH**2+VPH**2 WFWC038
VDOTY=VR*YR+VTH*YTH+VPH*YPH WFWC039
YLV=VDOTY/V2 WFWC040
YL2=VDOTY**2/V2 WFWC041
YT2=Y**2-YL2 WFWC042
YT4=YT2*YT2 WFWC043
CALL COLFRZ WFWC044
U=CMPLX(1.,-Z) WFWC045
JX=U-X WFWC046
UX2=UX*UX WFWC047
RAD=RAY*CSQRT(YT4+4.*YL2*UX2) WFWC048
D=2.*U*UX-YT2+RAD WFWC049
D2=D*D WFWC050
N2=1.-2.*X*UX/D WFWC051
PNPPS=2.*X*UX*(-1.+(YT2-2.*UX2)/RAD)/D2 WFWC052
PPSPR =YL2/Y*PYPR - (VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV WFWC053
PPSPTH=YL2/Y*PYPTH - (VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV WFWC054
PPSPPH=YL2/Y*PYPPH - (VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV WFWC055
PNPX=- (2.*U*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2 WFWC056
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y) WFWC057
PNPZ=I*X*(-2.*UX2-YT2+YT4/RAD)/D2 WFWC058
PNPR =PNPX*PXPR +PNPY*PYPR +PNPZ*PZPR +PNPPS*PPSPR WFWC059
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPTH WFWC060
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPPH WFWC061
PNPVR =PNPPS*(VR *YL2/V2-YLV*YR ) WFWC062
PNPVTH=PNPPS*(VTH*YL2/V2-YLV*YTH) WFWC063
PNPVPH=PNPPS*(VPH*YL2/V2-YLV*YPH) WFWC064
NNP=N2- (2.*X*PNPX+Y*PNPY+Z*PNPZ) WFWC065

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PNPT=PNPX*PXPT
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM
POLAR=-I*SQRT(V2)*(-YT2+RAD)/(2.*VDDTY*UX)
GAM=(-YT2+RAD)/(2.*UX)
LPOLAR=I*X*SQRT(YT2)/(UX*(U+GAM))
KAY2=OM2/C2*N2
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(REAL(KAY2)/K2)
KR =SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
C***** CALCULATES A HAMILTONIAN H
H=.5*(C2*K2/OM2-N2)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
PHPT =-PNPT
PHPR =-PNPR
PHPTH=-PNPTH
PHPPH=-PNPPH
PHPOM=-NNP/OM
PHPKR =C2/OM2*KR -C/OM*PNPVR
PHPKTH=C2/OM2*KTH-C/OM*PNPVTH
PHPKPH=C2/OM2*KPH-C/OM*PNPVPH
KPHPK=N2
RETURN
END
WFNC066
WFNC067
WFNC068
WFNC069
WFNC070
WFNC071
WFNC072
WFNC073
WFNC074
WFNC075
WFNC076
WFNC077
WFNC078
WFNC079
WFNC080
WFNC081
WFNC082
WFNC083
WFNC084
WFNC085
WFNC086
WFNC087
WFNC088
WFNC089
WFNC090
WFNC091
WFNC092

SUBROUTINE AHWFNC
C CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C APPLETON-HARTREE FORMULA WITH FIELD, NO COLLISIONS
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,
1 H,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPTI,PHPPH,PHPPHI
2 ,PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPI
3 ,KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI,SGN
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR
1 ,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP
COMMON /ZZ/ MODZ,Z(4)
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),WQ,W(400)
EQUIVALENCE (RAY,W(1)),(F,W(6))
LOGICAL SPACE
REAL KR,KTH,KPH,K2,KPHPK,KHPKI,KAY2,KAY2I,N2,NNP,LPOLAR,LPOLRI
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=0.),
1 (FIELD=1.),(KAY2I=0.),(HI=0.),(PHPTI=0.),(PHPRI=0.),
2 (PHPTI=0.),(PHPPHI=0.),(PHPOMI=0.),(PHPKRI=0.),(PHPKTI=0.),
3 (PHPKPI=0.),(KHPKI=0.),(POLAR=0.),(LPOLAR=0.),
4 (X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.),
5 (Y=0.), (PYPR=0.), (PYPTH=0.), (PYPPH=0.), (YR=0.), (PYRPR=0.),
6 (PYRPT=0.), (PYRPP=0.), (YTH=0.), (PYTPR=0.), (PYTPT=0.),
7 (PYTPP=0.), (YPH=0.), (PYPPR=0.), (PYPPT=0.), (PYPPP=0.),
8 (MODZ=1H),(U=1.)
ENTRY RINDEX
OM=PIT2*1.E6*F
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
VR =C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
WFNC001
WFNC002
WFNC003
WFNC004
WFNC005
WFNC006
WFNC007
WFNC008
WFNC009
WFNC010
WFNC011
WFNC012
WFNC013
WFNC014
WFNC015
WFNC016
WFNC017
WFNC018
WFNC019
WFNC020
WFNC021
WFNC022
WFNC023
WFNC024
WFNC025
WFNC026
WFNC027
WFNC028
WFNC029
WFNC030
WFNC031
WFNC032
WFNC033
WFNC034

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CALL ELECTX	WFNC035
CALL MAGY	WFNC036
V2=VR**2+VTH**2+VPH**2	WFNC037
VDOTY=VR*YR+VTH*YTH+VPH*YPH	WFNC038
YLV=VDOTY/V2	WFNC039
YL2=VDOTY**2/V2	WFNC040
YT2=Y**2-YL2	WFNC041
YT4=YT2*YT2	WFNC042
UX=U-X	WFNC043
UX2=UX*UX	WFNC044
RAD=RAY*SQRT(YT4+4.*YL2*UX2)	WFNC045
D=2.*UX-YT2+RAD	WFNC046
D2=0*D	WFNC047
N2=1.-2.*X*UX/D	WFNC048
PNPPS=2.*X*UX*(-1.+(YT2-2.*UX2)/RAD)/D2	WFNC049
PSPR=YL2/Y**2*YPR-(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV	WFNC050
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV	WFNC051
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV	WFNC052
PNPX=-(2.*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2	WFNC053
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)	WFNC054
NNP=N2-(2.*X**2*PNPX+Y*PNPY)	WFNC055
PNPR=PNPX*PXPR+PNPY*PYPR+PNPPS*PPSPR	WFNC056
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPPS*PPSPTH	WFNC057
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPPS*PPSPPH	WFNC058
PNPVR=PNPPS*(VR*YL2-VDOTY*YR)/V2	WFNC059
PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2	WFNC060
PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2	WFNC061
PNPT=PNPX*PXPT	WFNC062
SPACE=N2.EQ.1.	WFNC063
POLARI=SQRT(V2)*(YT2-RAD)/(2.*VDOTY*UX)	WFNC064
GAM=(-YT2+RAD)/(2.*UX)	WFNC065
LPOLRI=X*SQRT(YT2)/(UX*(U+GAM))	WFNC066
KAY2=OM2/C2*N2	WFNC067
IF(RSTART.EQ.0.) GO TO 1	WFNC068
SCALE=SQRT(KAY2/K2)	WFNC069
KR=SCALE*KR	WFNC070
KTH=SCALE*KTH	WFNC071
KPH=SCALE*KPH	WFNC072
1 CONTINUE	WFNC073
C***** CALCULATES A HAMILTONIAN H	WFNC074
H=.5*(C2*K2/OM2-N2)	WFNC075
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO	WFNC076
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.	WFNC077
PHPT=-PNPT	WFNC078
PHPR=-PNPR	WFNC079
PHPTH=-PNPTH	WFNC080
PHPPH=-PNPPH	WFNC081
PHPOH=-NNP/OM	WFNC082
PHPKR=C2/OM2*KR-C/OM*PNPVR	WFNC083
PHPKTH=C2/OM2*KTH-C/OM*PNPVTH	WFNC084
PHPKPH=C2/OM2*KPH-C/OM*PNPVPH	WFNC085
KHPK=N2	WFNC086
RETURN	WFNC087
END	WFNC088-



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SUBROUTINE A4NFWC
C      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C      APPLETON-HARTREE FORMULA -- NO FIELD, WITH COLLISIONS
COMMON /CONST/ PI,PIT2,PIJ2,DEGS,RADIAN,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH,
1      PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,
2      SGN
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON /YY/ MODY,Y(16)
COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH
COMMON /R</ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON R,TH,PH,KR,KTH,KPH /HW/ ID(10),W0,W(400)
EQUIVALENCE (RAY,W(1)),(F,W(6))
LOGICAL SPACE
REAL KR,KTH,KPH,K2
COMPLEX KAY2,H,PHPT,PHPR,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,
1      KPHPK,POLAR,LPOLAR,U,I,PNPX,PNPZ,
2      N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=1.),
1      (FIELD=0.), (POLAR=(0.,1.)), (LPOLAR=(0.,0.)),
2      (X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.),
3      (MODY=1H),
4      (Z=0.), (PZPR=0.), (PZPTH=0.), (PZPPH=0.),
5      (I=(0.,1.)), (ABSLIM=1.E-5), (PNPVR=0.), (PNPVTH=0.), (PNPVPH=0.)
ENTRY RINDEX
OM=PIT2*1.E6*F
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
VR =C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
CALL ELECTX
CALL COLFRZ
J=1.-I*Z
N2=1.-X/U
PNPX=-1./(2.*U)
PNPZ=-I*X/(2.*U**2)
NNP=N2-(2.*X*PNPX+Z*PNPZ)
PNPR =PNPX*PXPR +PNPZ*PZPR
PNPTH=PNPX*PXPTH+PNPZ*PZPTH
PNPPH=PNPX*PXPPH+PNPZ*PZPPH
PNPT=PNPX*PXPT
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM
KAY2=OM2/C2*N2
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(REAL(KAY2)/K2)
KR =SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
C***** CALCULATES A HAMILTONIAN H
H=.5*(C2*K2/OM2-N2)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
PHPT =-PNPT
PHPR =-PNPR
PHPTH=-PNPTH
PHPPH=-PNPPH
PHPOM=-NNP/OM
PHPKR =C2/OM2*KR
PHPKTH=C2/OM2*KTH
PHPKPH=C2/OM2*KPH
KPHPK=N2
RETURN
END

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SUBROUTINE AHNFNC
C      CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C      APPLETON-HARTREE FORMULA -- NO FIELD, NO COLLISIONS
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,
1      H,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPTI,PHPPH,PHPPHI,
2      PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPI
3      ,KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI,SGN
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON /YY/ MODY,Y(16) /ZZ/ MODZ,Z(4)
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400)
EQUIVALENCE (RAY,W(1)),(F,W(6))
LOGICAL SPACE
REAL N2,NNP,KR,KTH,KPH,K2,KPHPK,KHPKI,KAY2,KAY2I,LPOLAR,LPOLRI
DATA (MODRIN=8HAPPLETON,8H-HARTREE,8H FORMULA),(COLL=0.),
1      (FIELD=0.),(KAY2I=0.),(HI=0.),(PHPTI=0.),(PHPRI=0.),(
2      (PHPTI=0.),(PHPPHI=0.),(PHPOMI=0.),(PHPKRI=0.),(PHPKTI=0.),(
3      (PHPKPI=0.),(KPHPKI=0.),(POLAR=0.),(POLARI=1.),(LPOLAR=0.),(
4      (LPOLRI=1.),(
4      (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),(
5      (MODY=1H),(MODZ=1H),(
5      (NNP=1.),(PNPX=-0.5),(PNPVR=0.),(PNPVTH=0.),(PNPVPH=0.)
ENTRY RINDEX
OM=PIT2*1.E6*F
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
VR=C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
CALL ELECTX
PNPR=PNPX*PXPR
PNPTH=PNPX*PXPTH
PNPPH=PNPX*PXPPH
PNPT=PNPX*PXPT
N2=1.-X
SPACE=N2.EQ.1.
KAY2=OM2/C2*N2
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(KAY2/K2)
KR=SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
C***** CALCULATES A HAMILTONIAN H
H=.5*(C2*K2/OM2-N2)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
PHPT=-PNPT
PHPR=-PNPR
PHPTH=-PNPTH
PHPPH=-PNPPH
PHPOM=-NNP/OM
PHPKR=C2/OM2*KR
PHPKTH=C2/OM2*KTH
PHPKPH=C2/OM2*KPH
KPHPK=N2
RETURN
END
NFNC001
NFNC002
NFNC003
NFNC004
NFNC005
NFNC006
NFNC007
NFNC008
NFNC009
NFNC010
NFNC011
NFNC012
NFNC013
NFNC014
NFNC015
NFNC016
NFNC017
NFNC018
NFNC019
NFNC020
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NFNC041
NFNC042
NFNC043
NFNC044
NFNC045
NFNC046
NFNC047
NFNC048
NFNC049
NFNC050
NFNC051
NFNC052
NFNC053
NFNC054
NFNC055
NFNC056
NFNC057
NFNC058
NFNC059
NFNC060

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SUBROUTINE BQWFWC
C***** CALCULATES A HAMILTONIAN H
C***** (= BOOKER QUARTIC FOR VERTICAL INCIDENCE, S=0, C=1)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
C***** WITH FIELD, WITH COLLISIONS
COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),CJLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH,
1 PHPPH,PHPOH,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,
2 SGN
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR
1 ,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP
COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH
COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400)
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON /FLG/ NTP,NEWWR,NEWWP,PENET,LINES,IHOP,HPUNCH
EQUIVALENCE (RAY,W(1)),(F,W(6))
LOGICAL SPACE
REAL KR,KTH,KPH,K2,KDOTY,K4,KDOTY2
COMPLEX KAY2,I,PHPT,PHPR,PHPTH,PHPPH,PHPOH,PHPKR,PHPKTH,PHPKPH,
1 POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2,D2,
2 KPHPK,U2,A,B,ALPHA,BETA,GAMMA,PHPX,PHPY2,PHPK2,PHPU,PHPZ,
3 N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVP,PNPVP,PNPVP,NNP,PNPT
DATA (MODRIN=8HBOOKER Q,8HUARTIC, ,8HS=0, C=1),(COLL=1.),
1 (FIELD=1.),
2 (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),(
3 (Y=0.),(PYPR=0.),(PYPTH=0.),(PYPPH=0.),(YR=0.),(PYRPR=0.),(
4 (PYRPT=0.),(PYRPP=0.),(YTH=0.),(PYTPR=0.),(PYTPT=0.),(
5 (PYTPP=0.),(YPH=0.),(PYPPR=0.),(PYPPT=0.),(PYPPP=0.),(
6 (Z=0.),(PZPR=0.),(PZPTH=0.),(PZPPH=0.),(
7 (I=(0.,1.)),(ABSLIM=1.E-5),(SGN=1.)
ENTRY RINDEX
OM=PIT2*1.E6*F
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
CALL ELECTX
IF(X.LT..1) GO TO 2
K4=K2*K2
OM4=OM2*OM2
C4=C2*C2
CALL MAGY
Y2=Y*Y
KDOTY=KR*YR+KTH*YTH+KPH*YPH
KDOTY2=KDOTY*KDOTY
CALL COLFRZ
U=CHPLX(1.,-Z)
U2=U*U
UX=U-X
UX2=UX*UX
A=UX*U2-U*Y2
B=-2.*U*UX2+Y2*(2.*U-X)
ALPHA=A*C4*K4+X*KDOTY2*C4*K2
BETA=B*C2*OM2*K2-X*KDOTY2*C2*OM2
GAMMA=(UX2-Y2)*UX*OM4
H=ALPHA+BETA+GAMMA
PHPX=-U2*C4*K4+KDOTY2*C4*K2+(4.*U*UX-Y2)*C2*OM2*K2-KDOTY2*C2*OM2+
1 (-3.*UX2+Y2)*OM4
PHPY2=-U2*C4*K4+(2.*U-X)*C2*OM2*K2-UX*OM4
PHPKY2 =X*C2*(C2*K2-OM2)
PHPU=(2.*U*UX+U2-Y2)*C4*K4+2.*(Y2-UX2-2.*U*UX)*C2*K2*OM2+(3.*UX2
1 -Y2)*OM4
PHPZ=-I*PHPU
PHPK2=2.*A*C4*K2+X*KDOTY2*C4+B*C2*OM2

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BQWC001
BQWC002
BQWC003
BQWC004
BQWC005
BQWC006
BQWC007
BQWC008
BQWC009
BQWC010
BQWC011
BQWC012
BQWC013
BQWC014
BQWC015
BQWC016
BQWC017
BQWC018
BQWC019
BQWC020
BQWC021
BQWC022
BQWC023
BQWC024
BQWC025
BQWC026
BQWC027
BQWC028
BQWC029
BQWC030
BQWC031
BQWC032
BQWC033
BQWC034
BQWC035
BQWC036
BQWC037
BQWC038
BQWC039
BQWC040
BQWC041
BQWC042
BQWC043
BQWC044
BQWC045
BQWC046
BQWC047
BQWC048
BQWC049
BQWC050
BQWC051
BQWC052
BQWC053
BQWC054
BQWC055
BQWC056
BQWC057
BQWC058
BQWC059
BQWC060
BQWC061
BQWC062
BQWC063
BQWC064
BQWC065

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PHPT=PHPX*PXPT
PHPR =PHPX*PXPR +PHPY2*2.*Y*PYPR +PHPKY2 *2.*KDOTY*
1 (KR*PYRPR+KTH*PYTPR+KPH*PYPPR) +PHPZ*PZPR
PHPTH=PHPX*PXPTH+PHPY2*2.*Y*PYPTH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPT+KTH*PYTPT+KPH*PYPPT) +PHPZ*PZPTH
PHPPH=PHPX*PXPPH+PHPY2*2.*Y*PYPPH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPP+KTH*PYTPP+KPH*PYPPP) +PHPZ*PZPPH
PHPOM=(2.*BETA+4.*GAMMA)/OM
1 -2.*PHPX*X/OM-2.*PHPY2*Y2/OM-2.*PHPKY2 *KDOTY2/OM -PHPZ*Z/OM
PHPKR= 2.*PHPK2*KR +2.*KDOTY*PHPKY2 *YR
PHPKTH=2.*PHPK2*KTH+2.*KDOTY*PHPKY2 *YTH
PHPKPH=2.*PHPK2*KPH+2.*KDOTY*PHPKY2 *YPH
<AY2=K2*(-BETA+SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/(2.*ALPHA)
C
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT((-REAL(BETA)+SGN*RAY*SQRT(REAL(BETA)**2
1 -4.*REAL(ALPHA)*REAL(GAMMA)))/(2.*REAL(ALPHA)))
<R =SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1
CONTINUE
C***** THE FOLLOWING 3 CARDS USED FOR RAY TRACING IN COMPLEX SPACE
C IF(CABS((-BETA-SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/ALPHA-2.))
C 1LT.CABS((-BETA+SGN*RAY*CSQRT(BETA**2-4.*ALPHA*GAMMA))/ALPHA-2.)
C 2 .AND.RSTART.EQ.0.) SGN=-SGN
KPHPK=4.*ALPHA+2.*BETA
SPACE=CABS(C2*KAY2/OM2-1.) .LT. ABSLIM
POLAR =SQRT(K2)*(U+X*OM2/(C2*KAY2-OM2))/KDOTY*I
LPOLR = SQRT(Y2-KDOTY2/K2)/UX*(1.-C2*KAY2/OM2)*I
RETURN
C CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C APPLETON-HARTREE FORMULA WITH FIELDS, WITH COLLISIONS
2
CONTINUE
VR =C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
CALL MAGY
V2=VR**2+VTH**2+VPH**2
VDOTY=VR*YR+VTH*YTH+VPH*YPH
YLV=VDOTY/V2
YL2=VDOTY**2/V2
YT2=Y**2-YL2
YT4=YT2*YT2
CALL COLFRZ
U=CMPLX(1.,-Z)
UX=U-X
UX2=UX*UX
RAD=SGN*RAY*CSQRT(YT4+4.*YL2*UX2)
D=2.*U*UX-YT2+RAD
D2=D*D
N2=1.-2.*X*UX/D
PNPPS=2.*X*UX*(-1.+(YT2-2.*UX2)/RAD)/D2
PPSPR= YL2/Y*PYPR -(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV
PNPX=- (2.*U*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2
PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)
PNPZ=I*X*(-2.*UX2-YT2+YT4/RAD)/D2
NPN= N2-(2.*X*PNPX+Y*PNPY+Z*PNPZ)
PNPR =PNPX*PXPR +PNPY*PYPR +PNPZ*PZPR +PNPPS*PPSPR
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPTH
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPPH
PNPVR =PNPPS*(VR *YL2-VDOTY*YR )/V2
PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2
PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2
BQWC066
BQWC067
BQWC068
BQWC069
BQWC070
BQWC071
BQWC072
BQWC073
BQWC074
BQWC075
BQWC076
BQWC077
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BQWC120
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BQWC122
BQWC123
BQWC124
BQWC125
BQWC126
BQWC127
BQWC128
BQWC129
BQWC130

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	PNPT=PNPX*PXPT	BQWC131
	SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM	BQWC132
	POLAR=-I*SQRT(V2)*(-YT2+RAD)/(2.*VDOTY*UX)	BQWC133
	GAM=(-YT2+RAD)/(2.*UX)	BQWC134
	LPOLAR=I*X*SQRT(YT2)/(UX*(U+GAM))	BQWC135
	KAY2=OM2/C2*N2	BQWC136
	IF(RSTART.EQ.0.) GO TO 3	BQWC137
	SCALE=SQRT(REAL(KAY2)/K2)	BQWC138
	KR =SCALE*KR	BQWC139
	KTH=SCALE*KTH	BQWC140
	KPH=SCALE*KPH	BQWC141
3	CONTINUE	BQWC142
	H=.5*(C2*K2/OM2-N2)	BQWC143
	PHPT =-PNPT	BQWC144
	PHPR =-PNPR	BQWC145
	PHPTH=-PNPTH	BQWC146
	PHPPH=-PNPPH	BQWC147
	PHPOM=-NNP/OM	BQWC148
	PHPKR =C2/OM2*KR -C/OM*PNPVR	BQWC149
	PHPKTH=C2/OM2*KTH-C/OM*PNPVTH	BQWC150
	PHPKPH=C2/OM2*KPH-C/OM*PNPVPH	BQWC151
	KPHPK=N2	BQWC152
	RETURN	BQWC153
	END	BQWC154

	SUBROUTINE BQWFNC	BQNC001
	C***** CALCULATES A HAMILTONIAN H	BQNC002
	C***** (= BOOKER QUARTIC FOR VERTICAL INCIDENCE, S=0, C=1)	BQNC003
	C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO	BQNC004
	C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.	BQNC005
	C***** WITH FIELD, NO COLLISIONS	BQNC006
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,C,LOGTEN	BQNC007
	COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,KAY2I,	BQNC008
1	H,HI,PHPT,PHPTI,PHPR,PHPRI,PHPTH,PHPTI,PHPPH,PHPHI,	BQNC009
2	PHPOM,PHPOMI,PHPKR,PHPKRI,PHPKTH,PHPKTI,PHPKPH,PHPKPI	BQNC010
3	KPHPK,KPHPKI,POLAR,POLARI,LPOLAR,LPOLRI,SGN	BQNC011
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	BQNC012
	COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR	BQNC013
1	PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP	BQNC014
	COMMON /ZZ/ MODZ,Z(4)	BQNC015
	COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART	BQNC016
	COMMON R,TH,PH,KR,KTH,KPH /WW/ ID(10),W0,W(400)	BQNC017
	EQUIVALENCE (RAY,W(1)),(F,W(6))	BQNC018
	LOGICAL SPACE	BQNC019
	REAL N2,NNP,LPOLAR,LPOLARI,KR,KTH,KPH,K2,KDOTY,K4,KDOTY2,	BQNC020
1	KPHPK,KPHPKI,KAY2,KAY2I	BQNC021
	DATA (MODRIN=8HBOOKER Q,8HUARTIC, ,8HS=0, C=1),(COLL=0.),	BQNC022
1	(FIELD=1.),(KAY2I=0.),(HI=0.),(PHPTI=0.),(PHPRI=0.),	BQNC023
2	(PHPTI=0.),(PHPHI=0.),(PHPOMI=0.),(PHPKRI=0.),(PHPKTI=0.),	BQNC024
3	(PHPKPI=0.),(KPHPKI=0.),(POLAR=0.),(LPOLAR=0.),	BQNC025
4	(X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),	BQNC026
5	(Y=0.),(PYPR=0.),(PYPTH=0.),(PYPPH=0.),(YR=0.),(PYRPR=0.),	BQNC027
6	(PYRPT=0.),(PYRPP=0.),(YTH=0.),(PYTPR=0.),(PYTPT=0.),	BQNC028
7	(PYTPP=0.),(YPH=0.),(PYPPR=0.),(PYPPT=0.),(PYPPP=0.),	BQNC029
8	(MODZ=1H),(U=1.),(U2=1.)	BQNC030
	ENTRY RINDEX	BQNC031
	OM=PIT2*1.E6*	BQNC032
	C2=C*C	BQNC033
	K2=KR*KR+KTH*KTH+KPH*KPH	BQNC034
	OM2=OM*OM	BQNC035

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CALL ELECTX
IF(X.LT..1) GO TO 2
K4=K2*K2
OM4=OM2*OM2
C4=C2*C2
CALL MAGY
Y2=Y*Y
KDOTY=KR*YR+KTH*YTH+KPH*YPH
KDOTY2=KDOTY*KDOTY
UX=U-X
UX2=UX*UX
A=UX*U2-U*Y2
B=-2.*U*UX2+Y2*(2.*U-X)
ALPHA=A*C4*K4+X*KDOTY2*C4*K2
BETA=B*C2*OM2*K2-X*KDOTY2*C2*OM2
GAMMA=(UX2-Y2)*UX*OM4
H=ALPHA+BETA+GAMMA
P4PX=-U2*C4*K4+KDOTY2*C4*K2+(4.*U*UX-Y2)*C2*OM2*K2-KDOTY2*C2*OM2+
1 (-3.*UX2+Y2)*OM4
PHPY2=-U*C4*K4+(2.*U-X)*C2*OM2*K2-UX*OM4
PHPKY2 =X*C2*(C2*K2-OM2)
P4PK2=2.*A*C4*K2+X*KDOTY2*C4+B*C2*OM2
PHPT=PHPX*PXPT
PHPR =PHPX*PXPR +PHPY2*2.*Y*PYPR +PHPKY2 *2.*KDOTY*
1 (KR*PYRPR+KTH*PYTPR+KPH*PYPPR)
PHPTH=PHPX*PXPTH+PHPY2*2.*Y*PYPTH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPT+KTH*PYTPT+KPH*PYPPT)
PHPPH=PHPX*PXPPH+PHPY2*2.*Y*PYPPH+PHPKY2 *2.*KDOTY*
1 (KR*PYRPP+KTH*PYTPP+KPH*PYPPP)
PHPOH=(2.*BETA+4.*GAMMA)/OH
1 -2.*PHPX*X/OM-2.*PHPY2*Y2/OM-2.*PHPKY2 *KDOTY2/OM
PHPKR= 2.*PHPK2*KR +2.*KDOTY*PHPKY2 *YR
PHPKTH=2.*PHPK2*KTH+2.*KDOTY*PHPKY2 *YTH
PHPKPH=2.*PHPK2*KPH+2.*KDOTY*PHPKY2 *YPH
KAY2 = K2 *(-BETA+RAY*SQRT(BETA**2-4.*ALPHA*GAMMA))/(2.*ALPHA)
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(KAY2/K2)
KR =SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
KPHPK=4.*ALPHA+2.*BETA
SPACE=KAY2.EQ.OM2/C2
POLARI=SQRT(K2)*(U+X*OM2/(C2*KAY2-OM2))/KDOTY
LPOLRI= SQRT(Y2-KDOTY2/K2)/UX*(1.-C2*KAY2/OM2)
RETURN
C CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C APPLETON-HARTREE FORMULA WITH FIELD, NO COLLISIONS
2 CONTINUE
VR =C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
CALL MAGY
V2=VR**2+VTH**2+VPH**2
VDOTY=VR*YR+VTH*YTH+VPH*YPH
YLV=VDOTY/V2
YL2=VDOTY**2/V2
YT2=Y**2-YL2
YT4=YT2*YT2
JX=U-X
UX2=UX*UX
RAD=RAY*SQRT(YT4+4.*YL2*UX2)
D=2.*UX-YT2+RAD
D2=D*D
N2=1.-2.*X*UX/D
BQNC036
BQNC037
BQNC038
BQNC039
BQNC040
BQNC041
BQNC042
BQNC043
BQNC044
BQNC045
BQNC046
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BQNC099
BQNC100

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	PNPPS=2.*X*UX*(-1.+(YT2-2.*UX2)/RAD)/D2	BQNC101
	PPSPR=YL2/Y*PYPR-(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV	BQNC102
	PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV	BQNC103
	PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV	BQNC104
	PNPX=- (2.*UX2-YT2*(U-2.*X)+(YT4*(U-2.*X)+4.*YL2*UX*UX2)/RAD)/D2	BQNC105
	PNPY=2.*X*UX*(-YT2+(YT4+2.*YL2*UX2)/RAD)/(D2*Y)	BQNC106
	NNP=N2-(2.*X*PNPX+Y*PNPY)	BQNC107
	PNPR=PNPX*PXPR+PNPY*PYPR+PNPPS*PPSPR	BQNC108
	PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPPS*PPSPH	BQNC109
	PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPPS*PPSPPH	BQNC110
	PNPVR=PNPPS*(VR*YL2-VDOTY*YR)/V2	BQNC111
	PNPVTH=PNPPS*(VTH*YL2-VDOTY*YTH)/V2	BQNC112
	PNPVPH=PNPPS*(VPH*YL2-VDOTY*YPH)/V2	BQNC113
	PNPT=PNPX*PXPT	BQNC114
	SPACE=N2.EQ.1.	BQNC115
	POLARI=SQRT(V2)*(YT2-RAD)/(2.*VDOTY*UX)	BQNC116
	GAM=(-YT2+RAD)/(2.*UX)	BQNC117
	LPOLRI=X*SQRT(YT2)/(UX*(J+GAM))	BQNC118
	KAY2=OM2/C2*N2	BQNC119
	IF(RSTART.EQ.0.) GO TO 3	BQNC120
	SCALE=SQRT(KAY2/K2)	BQNC121
	KR=SCALE*KR	BQNC122
	KTH=SCALE*KTH	BQNC123
	KPH=SCALE*KPH	BQNC124
3	CONTINJE	BQNC125
	A=.5*(C2*K2/OM2-N2)	BQNC126
	PHPT=-PNPT	BQNC127
	PHPR=-PNPR	BQNC128
	PHPTH=-PNPTH	BQNC129
	PHPPH=-PNPPH	BQNC130
	PHPM=-NNP/OM	BQNC131
	PHPKR=C2/OM2*KR-C/OM*PNPVR	BQNC132
	PHPKTH=C2/OM2*KTH-C/OM*PNPVTH	BQNC133
	PHPKPH=C2/OM2*KPH-C/OM*PNPVPH	BQNC134
	KPHPK=N2	BQNC135
	RETURN	BQNC136
	END	BQNC137-

	SUBROUTINE SHWF	SHWF001
C	CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE	SHWF002
C	SEN-WYLLER FORMULA -- WITH FIELD	SHWF003
C	NEEDS SUBROUTINE FSW AND FUNCTIONS C AND S.	SHWF004
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RADIAN,K,SEA,LOGTEN	SHWF005
	COMMON /RIN/ MODRIN(3),CJLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH,	SHWF006
1	PHPPH,PHPM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,	SHWF007
2	SGN	SHWF008
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	SHWF009
	COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPR	SHWF010
1	PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP	SHWF011
	COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH	SHWF012
	COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART	SHWF013
	COMMON R,TH,P1,KR,KTH,KPH /HW/ ID(10),W0,W(400)	SHWF014
	EQUIVALENCE (RAY,W(1)),(F,W(6))	SHWF015
	LOGICAL SPACE	SHWF016
	REAL KR,KTH,KPH,K2	SHWF017
	COMPLEX KAY2,H,PHPT,PHPR,PHPTH,PHPPH,PHPM,PHPKR,PHPKTH,PHPKPH,	SHWF018
1	KPHPK,POLAR,LPOLAR,I,U,RAD,D,PNPPS,PNPX,PNPY,PNPZ,UX,UX2,	SHWF019
2	ALPHA,BETA,GAMMA,A,B,C,TEMP1,TEMP2,TEMP3,ALPOAL,BEPOBE,	SHWF020
3	GAPOGA,CB2,N2M1,J2,D2GA,DAL,DBET,DGAM,DADY,DAOZ,DBDY,DBDZ,	SHWF021
4	DCDY,DCDZ,DUOZ,DT1OX,DT1OY,DT1OZ,DT1OPS,DT2OX,DT2OY,DT2OZ,	SHWF022
5	DT2OPS,DRAODX,DRAODY,DRAOZ,DRDOPS,DOOX,DOOY,DOOZ,DOOPX,	SHWF023



6	UPX,N2,PNPR,PNPT1,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT	SWWF024
	DATA (MODRIN=8H SE,84N-WYLLER,8H FORMULA),(COLL=1.),	SWWF025
1	(FIELD=1.), (LPOLAR=(0.,0.)),	SWWF026
2	(X=0.), (PXPR=0.), (PXPTH=0.), (PXPPH=0.), (PXPT=0.),	SWWF027
3	(Y=0.), (PYPR=0.), (PYPTH=0.), (PYPPH=0.), (YR=0.), (PYRPR=0.),	SWWF028
4	(PYRPT=0.), (PYRPP=0.), (YTH=0.), (PYTPR=0.), (PYTPT=0.),	SWWF029
5	(PYTPP=0.), (YPH=0.), (PYPPR=0.), (PYPPT=0.), (PYPPP=0.),	SWWF030
6	(Z=0.), (PZPR=0.), (PZPTH=0.), (PZPPH=0.),	SWWF031
7	(I=(0.,1.)), (ABSLIM=1.E-5)	SWWF032
	ENTRY RINDEX	SWWF033
	OM=PIT2*1.E6*F	SWWF034
	C2=SEA*SEA	SWWF035
	K2=KR*KR+KTH*KTH+KPH*KPH	SWWF036
	OM2=OM*OM	SWWF037
	VR =SEA/OM*KR	SWWF038
	VTH=SEA/OM*KTH	SWWF039
	VPH=SEA/OM*KPH	SWWF040
	CALL ELECTX	SWWF041
	CALL MAGY	SWWF042
	OPY=1.+Y	SWWF043
	OMY=1.-Y	SWWF044
	CALL COLFRZ	SWWF045
	Z2=Z*Z	SWWF046
	CALL FSW(1./Z,ALPHA,DAL)	SWWF047
	ALPOAL=DAL/ALPHA	SWWF048
	CALL FSW(OMY/Z,BETA,DBET)	SWWF049
	BEPOBE=DBET/BETA	SWWF050
	CALL FSW(OPY/Z,GAMMA,DGAM)	SWWF051
	GAPOGA=DGAM/GAMMA	SWWF052
	U=Z/ALPHA	SWWF053
	DUDZ=(1.+ALPOAL/Z)/ALPHA	SWWF054
	U2=U*U	SWWF055
	UX=U-X	SWWF056
	JPX=U+X	SWWF057
	B=ALPHA/BETA	SWWF058
	DBDY=B*BEPOBE/Z	SWWF059
	DBDZ=-B*(ALPOAL-OMY*BEPOBE)/Z2	SWWF060
	C=ALPHA/GAMMA	SWWF061
	DCDY=-C*GAPOGA/Z	SWWF062
	DCDZ=-C*(ALPOAL-OPY*GAPOGA)/Z2	SWWF063
	A=.5*(B+C)-1.	SWWF064
	DADY=.5*(DBDY+DCDY)	SWWF065
	DAOZ=.5*(DBDZ+DCDZ)	SWWF066
	TEMP3=(1.-B*C)*U2+A*U*UPX	SWWF067
	V2= VR**2+VTH**2+VPH**2	SWWF068
	VDOY=VR*YR+VTH*YTH+VPH*YPH	SWWF069
	YL2=VDOY**2/V2	SWWF070
	YT2=Y**2-YL2	SWWF071
	Y2=Y*Y	SWWF072
	S2PSI=YT2/Y2	SWWF073
	C2PSI=YL2/Y2	SWWF074
	UX2=UX*UX	SWWF075
	CB2=(C-B)**2	SWWF076
	TEMP1=TEMP3*S2PSI	SWWF077
	DT1DX= A*U*S2PSI	SWWF078
	DT1DY=(U*JPX*DADY-U2*(B*DCDY+C*DBDY))*S2PSI	SWWF079
	DT1DZ=(2.*U*DUDZ*(1.-B*C+A)+A*X*DUDZ-U2*(B*DCDZ+C*DBDZ)+U*UPX*DAOZ	SWWF080
1	) *S2PSI	SWWF081
C	(1/YLYT) D/DPSI(TEMP1)	SWWF082
	DT1DPS=2.*TEMP1/YT2	SWWF083
	TEMP2=U2*CB2*JX2*C2PSI	SWWF084
	DT2DX=-2.*UX*U2*CB2*C2PSI	SWWF085
	DT2DY=2.*U2*UX2*C2PSI*(C-B)*(DCDY-DBDY)	SWWF086
	DT2DZ=2.*U2*UX2*C2PSI*(C-B)*(DCDZ-DBDZ)+2.*TEMP2*(1./U+1./UX)*DUDZ	SWWF087
C	(1/YLYT) D/DPSI(TEMP2)	SWWF088
	DT2DPS=-2.*TEMP2/YL2	SWWF089



```

RAD=RAY*CSQRT(TEMP1**2+TEMP2)
DRADDX=(TEMP1*DT1DX+.5*DT2DX)/RAD
DRAADDY=(TEMP1*DT1DY+.5*DT2DY)/RAD
DRAADDZ=(TEMP1*DT1DZ+.5*DT2DZ)/RAD
C (1/YLYT) D/DPSI(RAD)
DRDDPS=(TEMP1*DT1DPS+.5*DT2DPS)/RAD
D=2.*U*UX*(1.+A)-TEMP1+RAD+2.*A*U*X*S2PSI
DDDX=-2.*U-DT1DX*DRADDX+2.*A*U*S2PSI
DDDY=2.*U*UX*DADY-DT1DY*DRAADDY+2.*U*S2PSI*DADY
DDOZ=2.*(1.+A)*DUDZ*(U+UX)+2.*U*UX*DAOZ-DT1DZ*DRAADDZ+2.*X*S2PSI*
1 (A*DUDZ+J*JADZ)
C (1/YLYT) D/DPSI(D)
DDDPS=-DT1DPS+DRDDPS+2.*A*U*X/Y2
N2M1=-2.*X*(UX+U*A*S2PSI)/D
N2=1.+N2M1
C N D/OX(N)
PNPX=-(JX+U*A*S2PSI)*(1.-X*DDDX/D)/D+X/D
C N D/OY(N)
PNPY=-X*U*S2PSI/D*DADY-.5*N2M1/D*DDDY
C N D/OZ(N)
PNPZ=-X*(1.+A*S2PSI)/D*DUDZ-X*U*S2PSI/D*DAOZ-.5*N2M1/D*DDOZ
C (N/YLYT) D/DPSI(N)
PNPPS=-X*U*A/(D*Y2) -.5*N2M1/D*DDDPS
YLV=VDOTY/V2
C (YLYT) D/DR(PSI)
PPSPR=YL2/Y*PYPR-(VR*PYRPR+VTH*PYTPR+VPH*PYPPR)*YLV
C (YLYT) J/DTHETA(PHI)
PPSPTH=YL2/Y*PYPTH-(VR*PYRPT+VTH*PYTPT+VPH*PYPPT)*YLV
C (YLYT) D/DPHI(PHI)
PPSPPH=YL2/Y*PYPPH-(VR*PYRPP+VTH*PYTPP+VPH*PYPPP)*YLV
PNPR=PNPX*PXPR+PNPY*PYPR+PNPZ*PZPR+PNPPS*PPSPR
PNPTH=PNPX*PXPTH+PNPY*PYPTH+PNPZ*PZPTH+PNPPS*PPSPH
PNPPH=PNPX*PXPPH+PNPY*PYPPH+PNPZ*PZPPH+PNPPS*PPSPPH
PNPVR=PNPPS*(VR*YL2/V2-YLV*YR)
PNPVTH=PNPPS*(VTH*YL2/V2-YLV*YTH)
PNPVPH=PNPPS*(VPH*YL2/V2-YLV*YPH)
NNP=N2-(2.*X*PNPX+Y*PNPY+Z*PNPZ)
PNPT=PNPX*PXPT
POLAR=I*(TEMP1-RAD)*Y*SQRT(V2)/(U*UX*(C-B)*VDOTY)
COSPSI=VDOTY/(Y*SQRT(V2))
LPOLAR=(.5*I*(C-B)*POLAR+A*COSPSI)*SQRT(S2PSI)/
1 (POLAR*(JX*(1.+5*I*(C-B)*COSPSI)*POLAR)+A*(U-X*C2PSI))
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM
KAY2=OM2/32*N2
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(REAL(KAY2)/K2)
KR=SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
C***** CALCULATES A HAMILTONIAN H
H=.5*(C2*K2/OM2-N2)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
PHPT=-PNPT
PHPR=-PNPR
PHPTH=-PNPTH
PHPPH=-PNPPH
PHPOH=-NNP/OM
PHPKR=C2/OM2*KR-SEA/OM*PNPVR
PHPKTH=C2/OM2*KTH-SEA/OM*PNPVTH
PHPKPH=C2/OM2*KPH-SEA/OM*PNPVPH
KHPK=N2
RETURN
END

```

```

SWWF090
SWWF091
SWWF092
SWWF093
SWWF094
SWWF095
SWWF096
SWWF097
SWWF098
SWWF099
SWWF100
SWWF101
SWWF102
SWWF103
SWWF104
SWWF105
SWWF106
SWWF107
SWWF108
SWWF109
SWWF110
SWWF111
SWWF112
SWWF113
SWWF114
SWWF115
SWWF116
SWWF117
SWWF118
SWWF119
SWWF120
SWWF121
SWWF122
SWWF123
SWWF124
SWWF125
SWWF126
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SWWF136
SWWF137
SWWF138
SWWF139
SWWF140
SWWF141
SWWF142
SWWF143
SWWF144
SWWF145
SWWF146
SWWF147
SWWF148
SWWF149
SWWF150
SWWF151
SWWF152
SWWF153
SWWF154-

```

```

SUBROUTINE SWNF
C          CALCULATES THE REFRACTIVE INDEX AND ITS GRADIENT USING THE
C          SEN-WYLLER FORMULA -- NO FIELD
C          NEEDS SUBROUTINES FGSW AND FSW AND FUNCTIONS C AND S.
COMMON /CONST/ PI,PIT2,PIJ2,DEGS,RADIAN,K,C,LOGTEN
COMMON /RIN/ MODRIN(3),COLL,FIELD,SPACE,KAY2,H,PHPT,PHPR,PHPTH,
1          PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,KPHPK,POLAR,LPOLAR,
2          SGN
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON /YY/ MODY,Y(16)
COMMON /ZZ/ MOOZ,Z,PZPR,PZPTH,PZPPH
COMMON /RK/ N,STEP,MODE,E1MAX,E1MIN,E2MAX,E2MIN,FACT,RSTART
COMMON R,TH,PH,KR,KTH,KPH /HW/ IO(10),W0,W(400)
EQUIVALENCE (RAY,W(1)),(F,W(6))
LOGICAL SPACE
REAL KR,KTH,KPH,K2
COMPLEX KAY2,I,PHPT,PHPR,PHPTH,PHPPH,PHPOM,PHPKR,PHPKTH,PHPKPH,
1          KPHPK,POLAR,LPOLAR,PNPX,PNPZ,F1,DF,G1,DG1,
2          N2,PNPR,PNPTH,PNPPH,PNPVR,PNPVTH,PNPVPH,NNP,PNPT
DATA (MODRIN=8H SE,84N-WYLLER,8H FORMULA),(COLL=1.),
1      (FIELD=0.),(POLAR=(0.,1.)),(LPOLAR=(0.,0.)),
2      (X=0.),(PXPR=0.),(PXPTH=0.),(PXPPH=0.),(PXPT=0.),(
3      (MODY=1H ),
4      (Z=0.),(PZPR=0.),(PZPTH=0.),(PZPPH=0.),(
5      (ABSLIM=1.E-5),(PNPVR=0.),(PNPVTH=0.),(PNPVPH=0.)
ENTRY RINJEX
OM=PIT2*1.E6*F
C2=C*C
K2=KR*KR+KTH*KTH+KPH*KPH
OM2=OM*OM
VR =C/OM*KR
VTH=C/OM*KTH
VPH=C/OM*KPH
CALL ELECTX
CALL COLFRZ
CALL FGSW(1./Z,F1,DF1,G1,DG1)
N2=1.-X*G1
PNPX=-.5*G1
PNPZ=.5*X*DG1/Z**2
PNPR=PNPX*PXPR+PNPZ*PZPR
PNPTH=PNPX*PXPTH+PNPZ*PZPTH
PNPPH=PNPX*PXPPH+PNPZ*PZPPH
NNP=N2-(2.*X*PNPX+Z*PNPZ)
PNPT=PNPX*PXPT
SPACE=REAL(N2).EQ.1..AND.ABS(AIMAG(N2)).LT.ABSLIM
KAY2=OM2/C2*N2
IF(RSTART.EQ.0.) GO TO 1
SCALE=SQRT(REAL(KAY2)/K2)
KR =SCALE*KR
KTH=SCALE*KTH
KPH=SCALE*KPH
1 CONTINUE
C***** CALCULATES A HAMILTONIAN H
H=.5*(C2*K2/OM2-N2)
C***** AND ITS PARTIAL DERIVATIVES WITH RESPECT TO
C***** TIME, R, THETA, PHI, OMEGA, KR, KTHETA, AND KPHI.
PHPT =-PNPT
PHPR =-PNPR
PHPTH=-PNPTH
PHPPH=-PNPPH
PHPOM=-NNP/OM
PHPKR =C2/OM2*KR
PHPKTH=C2/OM2*KTH
PHPKPH=C2/OM2*KPH
KPHPK=N2
RETURN
END
SWNF001
SWNF002
SWNF003
SWNF004
SWNF005
SWNF006
SWNF007
SWNF008
SWNF009
SWNF010
SWNF011
SWNF012
SWNF013
SWNF014
SWNF015
SWNF016
SWNF017
SWNF018
SWNF019
SWNF020
SWNF021
SWNF022
SWNF023
SWNF024
SWNF025
SWNF026
SWNF027
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SWNF046
SWNF047
SWNF048
SWNF049
SWNF050
SWNF051
SWNF052
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SWNF054
SWNF055
SWNF056
SWNF057
SWNF058
SWNF059
SWNF060
SWNF061
SWNF062
SWNF063
SWNF064
SWNF065
SWNF066
SWNF067-

```

```

SUBROUTINE FGSW (X,F,DF,G,DG)
COMPLEX F,DF,G,DG
CALL FSW (X,F,DF)
IF (ABS(X).GT.50.) GO TO 1
G=X*F
DG=F+X*DF
RETURN
1 X2=X*X
  X3=X2*X
  T2=2.*X2
  T3=3.*X2
  T4=4.*X2
  T8=8.*X2
  T12=12.*X2
  T16=16.*X2
  G=CMPLX(1.-35./T4*(1.-99./T4*(1.-195./T4*(1.-323./T4)))/T4,
12.5*(1.-63./T4*(1.-143./T4*(1.-255./T4*(1.-399./T4)))/X)
  DG=.5*CMPLX(35.*(1.-99./T2*(1.-585./T8*(1.-323./T3*(1.-2415./T16)
1 ))/X3,
2-5.*(1.-189./T4*(1.-715./T12*(1.-357./T4*(1.-513./T4)))/X2)
RETURN
END

```

FGSW001  
FGSW002  
FGSW003  
FGSW004  
FGSW005  
FGSW006  
FGSW007  
FGSW008  
FGSW009  
FGSW010  
FGSW011  
FGSW012  
FGSW013  
FGSW014  
FGSW015  
FGSW016  
FGSW017  
FGSW018  
FGSW019  
FGSW020  
FGSW021  
FGSW 22-

```

SUBROUTINE FSW (Z,F,DF)
C      F(Z) = Z*C3/2(Z) + 2.5*I*C5/2(Z) AND DF(Z) = DF/DZ
C      WHERE THE INPUT Z IS REAL AND THE OUTPUT F AND DF ARE COMPLEX.
C      NEEDS THE SUBPROGRAMS FOR THE FRESNEL INTEGRAL FUNCTIONS S AND C
DIMENSION A(10),B(10),D(10)
COMPLEX F,DF,C1,C2,C3,C8,W,TEMP,I
DATA (I=(0.,1.)),(PI=3.1415926536),(A3=1.333333333)
DATA (C2=(1.,1.)),(C3=(1.,-1.)),(C4=.79788456 ),(C6=1.333333333)
C      C4=SQRT(2./PI)
DATA (A=.36230845E-02,.29579186E+00,.23193588E+01,.91355870E+01,
1.25856287E+02,.60488560E+02,.12562218E+03,.24214980E+03,
2.44918106E+03,.84244774E+03),
3 (B=.16747479E-02,.84796280E-01,.25285001E+00,.22665857E+00,
4.83871933E-01,.13811875E-01,.98017417E-03,.26299148E-04,
5.19761006E-06,.18781476E-09),
6 (D=.10080653E-03,.46117941E-01,.38507643E+00,.68507885E+00,
7.42648105E+00,.10742102E+00,.10985920E-01,.40924533E-03,
8.41881263E-05,.54513142E-08),(G=1.5045055)
C1=2./3.*I
C8=C2*A3*SQRT(PI/2.)
X=Z
X2=X*X
X3=X2*X
IF (ABS(X).GT.50.) GO TO 500
IF (ABS(X).GT.6.) GO TO 1
IF (ABS(X).LT..05) GO TO 200
C      FRESNEL
IF (X.GT.0.) GO TO 300
100 Y=C4*SQRT(-X)
  X2=X*X
  W=(COS(X)+I*SIN(X))*(1.-C3*(C(Y)+I*S(Y)))
  F =C1+C6*(X+C3*X*X/Y*W)
  DF=A3*CMPLX(1.,X)+CMPLX(1.5,X)*A3*C3*X/Y*W
  RETURN
300 Y=C4*SQRT(X)
  X2=X*X
  W=(COS(X)+I*SIN(X))*(1.-C2*(C(Y)-I*S(Y)))
  F =C1+C6*(X-C2*X*X/Y*W)
  DF=A3*CMPLX(1.,X)-CMPLX(1.5,X)*A3*C2*X/Y*W
RETURN

```

FSW 001  
FSW 002  
FSW 003  
FSW 004  
FSW 005  
FSW 006  
FSW 007  
FSW 008  
FSW 009  
FSW 010  
FSW 011  
FSW 012  
FSW 013  
FSW 014  
FSW 015  
FSW 016  
FSW 017  
FSW 018  
FSW 019  
FSW 020  
FSW 021  
FSW 022  
FSW 023  
FSW 024  
FSW 025  
FSW 026  
FSW 027  
FSW 028  
FSW 029  
FSW 030  
FSW 031  
FSW 032  
FSW 033  
FSW 034  
FSW 035  
FSW 036  
FSW 037  
FSW 038  
FSW 039  
FSW 040

C		POWER SERIES	FSW 041
	200	X=ABS(Z)	FSW 042
		X2=X*X	FSW 043
		X3=X2*X	FSW 044
		X4=X*X3	FSW 045
		X5=X*X4	FSW 046
		TEMP=-C8* SQRT(X)*CEXP(I*X)	FSW 047
		F=CMPLX(4./3.*X-16./9.*X3+64./315.*X5,2./3.+8./3.*X2-32./45.*X4)	FSW 048
	1	+TEMP*X	FSW 049
		DF=CMPLX(4./3.-16./3.*X2+64./63.*X4,16./3.*X-128./45.*X3	FSW 050
	1	+256./945.*X5)	FSW 051
	2	+TEMP*CMPLX(1.5,X)	FSW 052
		IF(Z.GE.0.) RETURN	FSW 053
		F=-CONJG(F)	FSW 054
		JF=CONJG(DF)	FSW 055
		RETURN	FSW 056
C		HERMITE	FSW 057
	1	XQ = X**2	FSW 058
		X2=XQ	FSW 059
		FR = 0.	FSW 060
		FI = 0.	FSW 061
		DFR = 0.	FSW 062
		DFI = 0.	FSW 063
		DO 2 J = 1,10	FSW 064
		SS = A(J) + XQ	FSW 065
		SB = B(J)/SS	FSW 066
		SD = D(J)/SS	FSW 067
		FR = FR + SB	FSW 068
		FI = FI + SD	FSW 069
		DFR = DFR + SB/SS	FSW 070
	2	DFI = DFI + SD/SS	FSW 071
		F = CMPLX(X*FR,FI)*G	FSW 072
		DF = G*(FR - 2.*X*CMPLX(X*DFR,DFI))	FSW 073
		RETURN	FSW 074
C		ASYMPTOTIC	FSW 075
	500	X2=X*X	FSW 076
		X3=X2*X	FSW 077
		X4=X3*X	FSW 078
		X5=X4*X	FSW 079
		T2=2.*X2	FSW 080
		T3=3.*X2	FSW 081
		T4=4.*X2	FSW 082
		T8=8.*X2	FSW 083
		T16=16.*X2	FSW 084
		T28=28.*X2	FSW 085
		F=CMPLX((1.-35./T4*(1.-99./T4*(1.-195./T4*(1.-323./T4))))/X	FSW 086
	1,5.	(1.-63./T4*(1.-143./T4*(1.-255./T4*(1.-399./T4))))/T2	FSW 087
		DF=-CMPLX((1.-105./T4*(1.-165./T4*(1.-273./T4*(1.-2907./T28))))/X2	FSW 088
	1,5.	(1.-63./T2*(1.-429./T8*(1.-255./T3*(1.-1995./T16))))/X3	FSW 089
		RETURN	FSW 090
		END	FSW 91-

```

FUNCTION C(X)
DOUBLEPRECISION  PIH, XD, Y, V, A, QZ, QN, Q, Z
DATA (A1=0.3183099), (A2=0.10132), (B1=0.0968), (B2=0.154)
PIH = 1.570796326794897
XA = ABS(X)
IF (XA.GT.4.) GOTO 20
C
XD = X
Y = PIH*XD*XD
V = Y*Y
A = 1.D0
Z = A
M = 15.*(XA + 1.)
DO 10 I = 1, M
KZ=2*(I-1)
KV=4*(I-1)
QZ = KV + 1
QN = (KZ + 1)*(KZ + 2)*(KV + 5)
Q = QZ/QN
A = -A*Q*V
10 Z = Z + A
Z = Z*XD
C = Z
RETURN
C
20 W = PIH*X*X
XV=XA**4
C=0.5+(A1-B1/XV)*SIN(W)/XA-(A2-B2/XV)*COS(W)/XA**3
IF (X.LT.0.) C = -C
RETURN
END

```

```

FUNCTION S(X)
DOUBLEPRFCISION  PIH, XD, Y, V, A, QZ, QN, Q, Z
DATA (A1=0.3183099), (A2=0.10132), (B1=0.0968), (B2=0.154)
PIH = 1.570796326794897
C
XA = ABS(X)
IF (XA.GT.4.) GOTO 20
C
XD = X
Y = PIH*XD*XD
V = Y*Y
A = Y/3.D0
Z = A
M = 15.*(XA + 1.)
DO 10 I = 1, M
KZ=2*(I-1)
KV=4*(I-1)
QZ = KV + 3
QN = (KZ + 2)*(KZ + 3)*(KV + 7)
Q = QZ/QN
A = -A*Q*V
10 Z = Z + A
Z = Z*XD
S = Z
RETURN
C
20 W = PIH*X*X
XV=XA**4
S=0.5-(A1-B1/XV)*COS(W)/XA-(A2-B2/XV)*SIN(W)/XA**3
IF (X.LT.0.) S = -S
RETURN
END

```

### APPENDIX 3. ELECTRON DENSITY SUBROUTINES WITH INPUT PARAMETER FORMS

The following electron density models are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Tabular profiles (TABLEX)	111
b.	Subroutine GAUSEL	113
c.	Chapman layer with tilts, ripples, and gradients (CHAPX)	115
d.	Chapman layer with variable scale height (VCHAPX)	117
e.	Double, tilted $\alpha$ -Chapman layer (DCHAPT)	118
f.	Linear Layer (LINEAR)	120
g.	Plain or quasi-parabolic layer (QPARAB)	121
h.	Analytic equatorial model (BULGE)	122
i.	Exponential profile (EXPX)	124

A further source of versatility in this ray tracing program is the ease with which specific ionospheric models, suited to the users needs, may be introduced. To add electron density models not included in the program, the user must write a subroutine that calculates the normalized electron density ( $X$ ) and its gradient ( $\partial X/\partial r$ ,  $\partial X/\partial \theta$ ,  $\partial X/\partial \varphi$ ) as a function of position in spherical coordinates ( $r$ ,  $\theta$ ,  $\varphi$ ). ( $X = 80.5 \times 10^{-6} N/f^2$ , where  $N$  is the electron density in  $\text{cm}^{-3}$  and  $f$  is the wave frequency in MHz.)

Both  $X$  and its gradient must be continuous functions of position. The formulas for  $\partial X/\partial r$ ,  $\partial X/\partial \theta$ , and  $\partial X/\partial \varphi$  must be consistent with the variation of  $X$  with  $r$ ,  $\theta$ , and  $\varphi$ . Otherwise, the program will run slowly and give incorrect results.

The coordinates  $r$ ,  $\theta$ ,  $\varphi$  refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine ( $r$ ,  $\theta$ ,  $\varphi$ ) is through blank common. (See

Table 3.) The output is through common block /XX/. (See Table 8.) It is useful if the name of the subroutine suggests the model to which it corresponds. The subroutine should have an entry point ELECTX so that other subroutines in the program can call it. Any parameters needed by the subroutine should be input into W101 through W149 of the W array. (See Table 2.) If the model needs massive amounts of data, these should be read in by the subroutine following the example of TABLEX. As in the already existing electron density subroutines, provision should be made for perturbations to the electron density model (irregularities) by having the statement

```
IF(PERT.NE.0.) CALL ELECT1
```

before the RETURN statement at the end of the subroutine.





```

SUBROUTINE TABLEX
C   CALCULATES ELECTRON DENSITY AND GRADIENT FROM PROFILES HAVING
C   THE SAME FORM AS THOSE USED BY CROFTS RAY TRACING PROGRAM
C   MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS
C   NEEDS SUBROUTINE GAUSEL
DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250),
1  DFLTA(250),SLOPE(250),MAT(4,5)
COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(F,W(6)),(READFN,W(100)),(PERT,W(150))
REAL MAT,K
DATA (MODX(1)=6HTABLEX)
ENTRY ELECTX
IF (READFN.EQ.0.) GO TO 10
READFN=0.
READ 1000, NOC,(HPC(I),FN2C(I),I=1,NOC)
1000 FORMAT (I4/(F8.2,E12.4))
PRINT 1200, (HPC(I),FN2C(I),I=1,NOC)
1200 FORMAT (I1,14X,6HHEIGHT,4X,16HELECTRON DENSITY/(1X,F20.10,E20.10))
A=0.
IF (FN2C(1).NE.0.) A=ALOG (FN2C(2)/FN2C(1))/(HPC(2)-HPC(1))
FN2C(1)=K*FN2C(1)
FN2C(2)=K*FN2C(2)
SLOPE(1)=A*FN2C(1)
SLOPE(NOC)=0.
NMAX=1
DO 6 I=2,NOC
IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I
IF (I.EQ.NOC) GO TO 4
FN2C(I+1)=K*FN2C(I+1)
DO 3 J=1,3
M=I+J-2
MAT(J,1)=1.
MAT(J,2)=HPC(M)
MAT(J,3)=HPC(M)**2
3  MAT(J,4)=FN2C(M)
CALL GAUSEL (MAT,4,3,4,NRANK)
IF (NRANK.LT.3) GO TO 60
SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)
4  DO 5 J=1,2
M=I+J-2
MAT(J,1)=1.
MAT(J,2)=HPC(M)
MAT(J,3)=HPC(M)**2
MAT(J,4)=HPC(M)**3
MAT(J,5)=FN2C(M)
L=J+2
MAT(L,1)=0.
MAT(L,2)=1.
MAT(L,3)=2.*HPC(M)
MAT(L,4)=3.*HPC(M)**2
5  MAT(L,5)=SLOPE(M)
CALL GAUSEL (MAT,4,4,5,NRANK)
IF (NRANK.LT.4) GO TO 60
ALPHA(I)=MAT(1,5)
BETA(I)=MAT(2,5)
GAMMA(I)=MAT(3,5)
6  DELTA(I)=MAT(4,5)
HMAX=HPC(NMAX)
NH=2
10 H=R(1)-EARTH
F2=F*F
PXPR=0.
IF (H.GE.HPC(1)) GO TO 12
11 NH=2
X=0.

```

```

TABX001
TABX002
TABX003
TABX004
TABX005
TABX006
TABX007
TABX008
TABX009
TABX010
TABX011
TABX012
TABX013
TABX014
TABX015
TABX016
TABX017
TABX018
TABX019
TABX020
TABX021
TABX022
TABX023
TABX024
TABX025
TABX026
TABX027
TABX028
TABX029
TABX030
TABX031
TABX032
TABX033
TABX034
TABX035
TABX036
TABX037
TABX038
TABX039
TABX040
TABX041
TABX042
TABX043
TABX044
TABX045
TABX046
TABX047
TABX048
TABX049
TABX050
TABX051
TABX052
TABX053
TABX054
TABX055
TABX056
TABX057
TABX058
TABX059
TABX060
TABX061
TABX062
TABX063
TABX064
TABX065
TABX066
TABX067

```

```

IF(FN2C(1).EQ.0.) GO TO 50
X=FN2C(1)*EXP(A*(H-HPC(1)))/F2
PXPR=A*X
GO TO 50
12 IF (H.GE.HPC(NOC)) GO TO 18
NSTEP=1
IF (H.LT.HPC(NH-1)) NSTEP=-1
15 IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16
NH=NH+NSTEP
GO TO 15
16 X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH)))/F2
PXPR=(BETA(NH)+H*(2.*GAMMA(NH)+H*3.*DELTA(NH)))/F2
GO TO 50
18 X=FN2C(NOC)/F2
50 IF (PERT.NE.0.) CALL ELECT1
RETURN
60 PRINT 6000, I,HPC(I)
6000 FORMAT(4H THE,I4,55HTH POINT IN THE ELECTRON DENSITY PROFILE HAS T
1HE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)
CALL EXIT
END

```

```

TABX068
TABX069
TABX070
TABX071
TABX072
TABX073
TABX074
TABX075
TABX076
TABX077
TABX078
TABX079
TABX080
TABX081
TABX082
TABX083
TABX084
TABX085
TABX086
TABX087
TABX088

```

```

SUBROUTINE GAUSEL (C,NRD,NRR,NCC,NSF)
C***** SAME AS SUBROUTINE GAUSSEL WRITTEN BY L. DAVID LEWIS *****
DIMENSION C(NRD,NCC),L(128,2)
C BITS = 2.**-18
DATA (BITS=3.8146972656E-6)
NR=NRR
NC=NCC
IF(NC.LT.NR.OR.NR.GT.128.OR.NR.LE.0) CALL EXIT
C
C INITIALIZE.
NSF=0
NRM=NR-1
NRP=NR+1
D=1.
LSD=1
DO 1 KR=1,NR
L(KR,1)=KR
1 L(KR,2)=0
IF(NR.EQ.1) GO TO 42
C
C ELIMINATION PHASE.
DO 41 KP=1,NRM
KPP=KP+1
PM=0.
MPN=0
C
C SEARCH COLUMN KP FROM DIAGONAL DOWN FOR MAX PIVOT.
DO 2 KR=KP,NR
LKR=L(KR,1)
PT=ABS(C(LKR,KP))
IF(PT.LE.PM) GO TO 2
PM=PT
MPN=KR
LMP=LKR
2 CONTINUE
C
C IF MAX PIVOT IS ZERO, MATRIX IS SINGULAR.
IF(MPN.EQ.0) GO TO 9
NSF=NSF+1
IF(MPN.EQ.KP) GO TO 3

```

```

GAUS001
GAUS002
GAUS003
GAUS004
GAUS005
GAUS006
GAUS007
GAUS008
GAUS009
GAUS010
GAUS011
GAUS012
GAUS013
GAUS014
GAUS015
GAUS016
GAUS017
GAUS018
GAUS019
GAUS020
GAUS021
GAUS022
GAUS023
GAUS024
GAUS025
GAUS026
GAUS027
GAUS028
GAUS029
GAUS030
GAUS031
GAUS032
GAUS033
GAUS034
GAUS035
GAUS036
GAUS037
GAUS038
GAUS039
GAUS040

```

C		GAUS041
C	NEW ROW NUMBER KP HAS MAX PIVOT.	GAUS042
	LSD=-LSD	GAUS043
	L(KP,2)=L(MPN,1)=L(KP,1)	GAUS044
	L(KP,1)=LMP	GAUS045
C		GAUS046
C	ROW OPERATIONS TO ZERO COLUMN KP BELOW DIAGONAL.	GAUS047
3	MKP=L(KP,1)	GAUS048
	P=C(MKP,KP)	GAUS049
	D=D*P	GAUS050
	DO 41 KR=KPP, NR	GAUS051
	MKR=L(KR,1)	GAUS052
	Q=C(MKR,KP)/P	GAUS053
	IF(Q.EQ.0.) GO TO 41	GAUS054
C		GAUS055
C	SUBTRACT Q * PIVOT ROW FROM ROW KR.	GAUS056
	DO 4 LC=KPP, NC	GAUS057
	R=Q*C(MKP, LC)	GAUS058
	C(MKR, LC)=C(MKR, LC)-R	GAUS059
4	IF(ABS(C(MKR, LC)).LT.ABS(R)*BITS) C(MKR, LC)=0.	GAUS060
41	CONTINUE	GAUS061
C		GAUS062
C	LOWER RIGHT HAND CORNER.	GAUS063
42	LNR=L(NR,1)	GAUS064
	P=C(LNR, NR)	GAUS065
	IF(P.EQ.0.) GO TO 9	GAUS066
	NSF=NSF+1	GAUS067
	D=D*P*LSD	GAUS068
	IF(NR.EQ.NC) GO TO 8	GAUS069
C		GAUS070
C	BACK SOLUTION PHASE.	GAUS071
	DO 61 MC=NRP, NC	GAUS072
	C(LNR, MC)=C(LNR, MC)/P	GAUS073
	IF(NR.EQ.1) GO TO 61	GAUS074
	DO 6 LL=1, NRM	GAUS075
	KR=NR-LL	GAUS076
	MR=L(KR,1)	GAUS077
	KRP=KR+1	GAUS078
	DO 5 MS=KRP, NR	GAUS079
	LMS=L(MS,1)	GAUS080
	R=C(MR, MS)*C(LMS, MC)	GAUS081
	C(MR, MC)=C(MR, MC)-R	GAUS082
5	IF(ABS(C(MR, MC)).LT.ABS(R)*BITS) C(MR, MC)=0.	GAUS083
6	C(MR, MC)=C(MR, MC)/C(MR, KR)	GAUS084
61	CONTINUE	GAUS085
C		GAUS086
C	SHUFFLE SOLUTION ROWS BACK TO NATURAL ORDER.	GAUS087
	DO 71 LL=1, NRM	GAUS088
	KR=NR-LL	GAUS089
	MKR=L(KR,2)	GAUS090
	IF(MKR.EQ.0) GO TO 71	GAUS091
	MKP=L(KR,1)	GAUS092
	DO 7 LC=NRP, NC	GAUS093
	Q=C(MKR, LC)	GAUS094
	C(MKR, LC)=C(MKP, LC)	GAUS095
7	C(MKP, LC)=Q	GAUS096
71	CONTINUE	GAUS097
C		GAUS098
C	NORMAL AND SINGULAR RETURNS. GOOD SOLUTION COULD HAVE D=0.	GAUS099
8	C(1,1)=D	GAUS100
	GO TO 91	GAUS101
9	C(1,1)=0.	GAUS102
91	RETURN	GAUS103
	END	GAUS104-

## INPUT PARAMETER FORM FOR SUBROUTINE CHAPX

An ionospheric electron density model consisting of a Chapman layer with tilts, ripples, and gradients

$$f_N^2 = f_c^2 \exp\left(\alpha(1-z-e^{-z})\right)$$

$$z = \frac{h - h_{\max}}{H}$$

$$f_c^2 = f_{c0}^2 \left(1 + A \sin\left(2\pi\left(\theta - \frac{\pi}{2}\right)/B\right) + C\left(\theta - \frac{\pi}{2}\right)\right)$$

$$h_{\max} = h_{\max 0} + E\left(\theta - \frac{\pi}{2}\right) R_0$$

$f_N$  is the plasma frequency

$h$  is the height above the ground

$R_0$  is the radius of the earth in km

and  $\theta$  is the colatitude in radians.

Specify:

Critical frequency at the equator,  $f_{c0} =$  \_\_\_\_\_ MHz (W101)

Height of the maximum electron density at the equator,  $h_{\max 0} =$  \_\_\_\_\_ km (W102)

Scale height,  $H =$  \_\_\_\_\_ km (W103)

$\alpha =$  \_\_\_\_\_ (W104, 0.5 for an  $\alpha$  Chapman layer, 1.0 for a  $\beta$  Chapman layer)

Amplitude of periodic variation of  $f_c^2$  with latitude,  $A =$  \_\_\_\_\_ (W105)

Period of variation of  $f_c^2$  with latitude,  $B =$  \_\_\_\_\_  $\frac{\text{rad}}{\text{km}}$  (W106)

Coefficient of linear variation of  $f_c^2$  with latitude,  $C =$  \_\_\_\_\_  $\text{rad}^{-1}$  (W107)

Tilt of the layer,  $E =$  \_\_\_\_\_  $\frac{\text{rad}}{\text{deg}}$  (W108)

	SUBROUTINE CHAPX	CHAP001
	CHAPMAN LAYER WITH TILTS, RIPPLES, AND GRADIENTS	CHAP002
C	W(104) = 0.5 FOR AN ALPHA-CHAPMAN LAYER	CHAP003
C	= 1.0 FOR A BETA-CHAPMAN LAYER	CHAP004
C	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	CHAP005
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	CHAP006
	COMMON R(6) /WW/ ID(10),W0,W(400)	CHAP007
	EQUIVALENCE (THETA,R(2))	CHAP008
	EQUIVALENCE (EARTH,R(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	CHAP009
	1 (SH,W(103)),(ALPHA,W(104)),(A,W(105)),(B,W(106)),(C,W(107)),	CHAP010
	2 (E,W(108)),(PERT,W(150))	CHAP011
	DATA (MODX(1)=6H CHAPX)	CHAP012
	ENTRY ELECTX	CHAP013
	THETA2=THETA-PID2	CHAP014
	HMAX=HM+EARTH*E*THETA2	CHAP015
	H=R(1)-EARTH	CHAP016
	Z=(H-HMAX)/SH	CHAP017
	D=0.	CHAP018
	IF (B.NE.0.) D=PIT2/B	CHAP019
	TEMP=1.+A*SIN(D*THETA2)+C*THETA2	CHAP020
	EXZ=1.-EXP(-Z)	CHAP021
	X=(FC/F)**2*TEMP*EXP(ALPHA*(EXZ-Z))	CHAP022
	PXPR=-ALPHA*X*EXZ/SH	CHAP023
	PXPTH=X*(D*A*SIN(PID2-D*THETA2)+C)/TEMP-PXPR*EARTH*E	CHAP024
	IF (PERT.NE.0.) CALL ELECT1	CHAP025
	RETURN	CHAP026
	END	CHAP 27-

## INPUT PARAMETER FORM FOR SUBROUTINE VCHAPX

An ionospheric electron density model consisting of a Chapman layer with variable scale height

$$f_N^2 = f_C^2 \tau^{\frac{1}{2}} e^{\frac{1}{2}} (1 - \tau)$$

$$\tau = \left( \frac{h_{\max}}{h} \right)^\chi$$

h is the height above the ground.

Specify:

critical frequency,  $f_c =$  \_\_\_\_\_ MHz (W101)

height of maximum electron density,  $h_{\max} =$  \_\_\_\_\_ km (W102)

$\chi =$  \_\_\_\_\_ (W103)

SUBROUTINE VCHAPX	VCHA001
CHAPMAN LAYER WITH VARIABLE SCALE HEIGHT	VCHA002
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	VCHA003
COMMON R(6) /WW/ ID(10),W0,W(400)	VCHA004
EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	VCHA005
1 (CHI,W(103)),(PERT,W(150))	VCHA006
DATA (MODX(1)=6HVCHAPX)	VCHA007
ENTRY ELECTX	VCHA008
HMAX=HM	VCHA009
X=PXPR=0.	VCHA010
H=R(1)-EARTH	VCHA011
IF (H.LE.0.) GO TO 50	VCHA012
TAU=(HM/H)**CHI	VCHA013
X=(FC/F)**2*SQRT(TAU)*EXP(0.5*(1.-TAU))	VCHA014
PXPR=.5*X*(TAU-1.)*CHI/H	VCHA015
50 IF (PERT.NE.0.) CALL ELECT1	VCHA016
RETURN	VCHA017
END	VCHA018-

INPUT PARAMETER FORM FOR SUBROUTINE DCHAPT

An ionospheric electron density model consisting of a double, tilted  $\alpha$ -Chapman layer

$$f_N^2 = f_{c1}^2 \exp \frac{1}{2} (1 - z_1 - e^{-z_1}) + f_{c2}^2 \exp \frac{1}{2} (1 - z_2 - e^{-z_2})$$

$$z_1 = \frac{h - h_{m1}}{H_1} \quad ; \quad z_2 = \frac{h - h_{m2}}{H_2}$$

$$f_{c1}^2 = f_{c10}^2 C(\theta - \pi/2)$$

$$f_{c2}^2 = f_{c20}^2 C(\theta - \pi/2)$$

$$h_{m1} = h_{m10} + R_o E \left( \frac{\pi}{180} \right) \left( \theta - \frac{\pi}{2} \right)$$

$$h_{m2} = h_{m20} + R_o E \left( \frac{\pi}{180} \right) \left( \theta - \frac{\pi}{2} \right)$$

Specify:

$$f_{c10} = \text{_____} \text{ MHz (} f_{c1} \text{ at equator)} \quad \text{(W101)}$$

$$h_{m10} = \text{_____} \text{ Km (} h_{m1} \text{ at equator)} \quad \text{(W102)}$$

$$H_1 = \text{_____} \text{ Km} \quad \text{(W103)}$$

$$f_{c20} = \text{_____} \text{ MHz (} f_{c2} \text{ at equator)} \quad \text{(W104)}$$

$$h_{m20} = \text{_____} \text{ Km (} h_{m2} \text{ at equator)} \quad \text{(W105)}$$

$$H_2 = \text{_____} \text{ Km} \quad \text{(W106)}$$

$$C = \text{_____} \text{ rad}^{-1} \text{ (fractional change in } f_{c1}, f_{c2}, \text{ position for increases southward)} \quad \text{(W107)}$$

$$E = \text{_____} \text{ deg (positive for upward tilt to the south)} \quad \text{(W108)}$$

SUBROUTINE DCHAPT	DCHA001
TWO CHAPMAN LAYERS WITH TILTS	DCHA002
COMMON /CONST/ PI,PIT2,PID2,DUM(5)	DCHA003
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	DCHA004
COMMON R(6) /WW/ ID(10),WO,W(400)	DCHA005
EQUIVALENCE (EARTH,W(2)),(F,W(6)),(FC1,W(101)),(HM1,W(102)),	DCHA006
1 (SH1,W(103)),(FC2,W(104)),(HM2,W(105)),(SH2,W(106)),(C,W(107)),	DCHA007
2 (E,W(108)),(PERT,W(150))	DCHA008
DATA (MODX(1)=6HDCHAPT)	DCHA009
ENTRY ELECTX	DCHA010
EARTHE=EARTH*E	DCHA011
THETA2=R(2)-PID2	DCHA012
HMAX=HM1+EARTHE*THETA2	DCHA013
X=PXPR=PXPTH=0.	DCHA014
H=R(1)-EARTH	DCHA015
IF (H.LT.0.) GO TO 50	DCHA016
Z1=(H-HMAX)/SH1	DCHA017
EXPZ1=1.-EXP(-Z1)	DCHA018
TEMP=1.+C*THETA2	DCHA019
X=(FC1/F)**2*TEMP*EXP(.5*(EXPZ1-Z1))	DCHA020
PXPR=-0.5*X*EXPZ1/SH1	DCHA021
PXPTH=X*C/TEMP-PXPR*EARTHE	DCHA022
IF (FC2.EQ.0.) GO TO 50	DCHA023
Z2=(H-HM2-EARTHE*THETA2)/SH2	DCHA024
EXPZ2=1.-EXP(-Z2)	DCHA025
X2=(FC2/F)**2*TEMP*EXP(.5*(EXPZ2-Z2))	DCHA026
X=X+X2	DCHA027
PXPR2=-0.5*X2*EXPZ2/SH2	DCHA028
PXPR=PXPR+PXPR2	DCHA029
PXPTH=PXPTH+X2*C/TEMP-PXPR2*EARTHE	DCHA030
50 IF (PERT.NE.0.) CALL ELECT1	DCHA031
RETURN	DCHA032
END	DCHA033-



## INPUT PARAMETER FORM FOR SUBROUTINE LINEAR

An ionospheric electron density model consisting of a linear layer

$$N = 0 \quad \text{for } h \leq h_{\min}$$

$$N = A(h - h_{\min}) \quad \text{for } h > h_{\min}$$

The ray will penetrate if  $h > h_{\max}$ .

Specify:

A = \_\_\_\_\_ electrons/cm<sup>3</sup>/ km (W101)

$h_{\max}$  = \_\_\_\_\_ km (W102)

$h_{\min}$  = \_\_\_\_\_ km (W103)

	SUBROUTINE LINEAR	LINE001
C	LINEAR ELECTRON DENSITY MODEL	LINE002
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)	LINE003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	LINE004
	COMMON R(6) /WW/ ID(10),W0,W(400)	LINE005
	EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FACT,W(101)),(HM,W(102)),	LINE006
	1 (HMIN,W(103)),(PERT,W(150))	LINE007
	REAL K	LINE008
	DATA (MODX(1)=6HLINEAR)	LINE009
	ENTRY ELECTX	LINE010
	H=R(1)-EARTH	LINE011
	HMAX=HM	LINE012
	X=PXPR*0.	LINE013
	IF (H.LE.HMIN) GO TO 50	LINE014
	PXPR=K*FACT/F**2	LINE015
	X=PXPR*(H-HMIN)	LINE016
50	IF (PERT.NE.0.) CALL ELECT1	LINE017
	RETURN	LINE018
	END	LINE019-

## INPUT PARAMETER FORM FOR SUBROUTINE QPARAB

An ionospheric electron density model consisting of a parabolic or a quasi-parabolic layer (concentric)

$$f_N^2 = f_c^2 \left[ 1 - \frac{h-h_{\max}}{Y_m} \cdot C^2 \right] \quad \text{if } f_N^2 > 0.$$

$$f_N^2 = 0, \quad \text{otherwise.}$$

$C = 1$ , for a parabolic layer

$$C = \frac{R_o + h_{\max} - Y_m}{R_o + h} \quad \text{for a quasi-parabolic layer}$$

where  $R_o$  is the radius of the earth.

Specify:

Critical frequency,  $f_c =$  \_\_\_\_\_ Mc/s (W101)

Height of maximum electron density,  $h_{\max} =$  \_\_\_\_\_ km. (W102)

Semi-thickness,  $Y_m =$  \_\_\_\_\_ km. (W103)

Type of profile:

Plain parabolic \_\_\_\_\_ (W104 = 0.)

Quasi-parabolic \_\_\_\_\_ (W104 = 1.)

	SUBROUTINE QPARAB	PARA001
C	PLAIN PARABOLIC OR QUASI-PARABOLIC PROFILE	PARA002
C	W(104) = 0. FOR A PLAIN PARABOLIC PROFILE	PARA003
C	= 1. FOR A QUASI-PARABOLIC PROFILE	PARA004
	COMMON /XX/ MODX(2),X,PXPR,PXPTR,PXPPH,PXPT,HMAX	PARA005
	COMMON R(6) /WW/ ID(10),W0,W(400)	PARA006
	EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FC,W(101)),(HM,W(102)),	PARA007
	1 (YM,W(103)),(QUASI,W(104)),(PERT,W(150))	PARA008
	DATA (MODX(1)=6HQPAB)	PARA009
	ENTRY ELECTX	PARA010
	HMAX=HM	PARA011
	PXPR=0.	PARA012
	H=R(1)-EARTH	PARA013
	FCF2=(FC/F)**2	PARA014
	CONST=1.	PARA015
	IF (QUASI.EQ.1.) CONST=(EARTH+HM-YM)/R(1)	PARA016
	Z=(H-HM)/YM*CONST	PARA017
	X=MAX1F(0.,FCF2*(1.-Z**2))	PARA018
	IF (X.EQ.0.) GO TO 50	PARA019
	IF (QUASI.EQ.1.) CONST=(EARTH+HM)*(EARTH+HM-YM)/R(1)**2	PARA020
	PXPR=-2.*Z*FCF2/YM*CONST	PARA021
50	IF (PERT.NE.0.) CALL ELECT1	PARA022
	RETURN	PARA023
	END	PARA024-

## INPUT PARAMETER FORM FOR SUBROUTINE BULGE

An analytic ionospheric electron density model which represents the general latitude variation of the equatorial ionosphere (afternoon, equinox, sunspot maximum) - see the center panel of figure 3.18b, page 133 of Davies (1965).

The model is an alpha Chapman layer with parameters which vary with geomagnetic latitude.

$$f_N^2 = f_c^2 e^{\frac{1}{2}(1-z-e^{-z})}$$

$$\text{where } z = \frac{h - h_{\max}}{H}$$

$f_N$  is the plasma frequency

$f_c$  is the critical frequency

$h_{\max}$  is the height of the maximum electron density

$H$  is the scale height

$h$  is height

---

$f_c$ ,  $h_{\max}$ ,  $H$  vary with geomagnetic latitude in the following way:

if  $h < 100$  km,  $h_{\max} = 350$  km,  $f_c = 15$  Mc/s

---

For  $h \geq 100$  km,

$h_{\max} = 350$  if  $\lambda \geq 24^\circ$

$h_{\max} = 430 + 80 \cos\left(\frac{180}{24} \lambda\right)$  if  $\lambda < 24^\circ$

$\lambda$  is the geomagnetic latitude in degrees

---

$$f_c = \sqrt{50 \left(\frac{\lambda}{8}\right)^2 \exp\left(2 - \left|\frac{\lambda}{8}\right|\right) + 40}$$


---

In all cases  $H$  is determined by the constraint that

$$f_N = 2 \text{ Mc/s at } 100 \text{ km.}$$


---

	SUBROUTINE BULGE	BULG001
C	ANALYTICAL MODEL OF THE VARIATION OF THE EQUATORIAL F2 LAYER	BULG002
C	IN GEOMAGNETIC LATITUDE (EQUATORIAL BULGE AND ANOMALY)	BULG003
C	SEE FIGURE 3.18B, PAGE 133 IN DAVIES (1965).	BULG004
C	THIS MODEL HAS NO VARIATION IN GEOMAGNETIC LONGITUDE.	BULG005
	COMMON /CONST/ PI,PID2,PID2,DUM(5)	BULG006
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	BULG007
	COMMON R(6) /WW/ ID(10),W0,W(400)	BULG008
	EQUIVALENCE (EARTH,R(2)),(F,W(6)),(PERT,W(150))	BULG009
	DATA (MODX(1)=6H BULGE)	BULG010
	ENTRY ELECTX	BULG011
	H=R(1)-EARTH	BULG012
	PHMPH=PFC2PTH=0.	BULG013
	HMAX=350.	BULG014
	FC2=225.	BULG015
	IF(H.LT.100.) GO TO 2	BULG016
C	EQUATORIAL BULGE	BULG017
	BULLAT=7.5*(PID2-R(2))	BULG018
	IF(ABS(BUL LAT).GE.PI) GO TO 1	BULG019
	HMAX=430.+80.*COS(BULLAT)	BULG020
	PHMPH=600.*SIN(BUL LAT)	BULG021
C	EQUATORIAL ANOMALY	BULG022
1	ANMLAT=22.5*(PID2-R(2))/PI	BULG023
	POW=2.-ABS(ANM LAT)	BULG024
	FC2=50.*ANM LAT**2*EXP( POW ) + 40.	BULG025
	PFC2PTH=-1125./PI*POW*ANMLAT*EXP(POW)	BULG026
C	FORCING PLASMA FREQ AT 100 KM TO BE 2 MHZ IN ORDER TO CALCULATE SH	BULG027
2	ALPHA=2.*ALOG(FC2/4.)+1.	BULG028
	Z100=-ALOG(ALPHA)	BULG029
	DO 3 I=1,5	BULG030
3	Z100=-ALOG(ALPHA-Z100)	BULG031
	SH=(100.-HMAX)/Z100	BULG032
	Z=(H-HMAX)/SH	BULG033
	EXZ=1.-EXP(-Z)	BULG034
	X=FC2*EXP(.5*(EXZ-Z))/F**2	BULG035
	PXPR=-0.5*X*EXZ/SH	BULG036
	PXPTH=-PXPR*(1.-Z/Z100)*PHMPH+(1.-Z*EXZ/(Z100*(1.-EXP(-Z100))))	BULG037
1	*X/FC2*PFC2PTH	BULG038
	IF (PERT.NE.0.) CALL ELECT1	BULG039
	RETURN	BULG040
	END	BULG041-

## INPUT PARAMETER FORM FOR SUBROUTINE EXPX

An exponential electron density profile

$$N = N_0 e^{a(h-h_0)}$$

$h$  is the height above the ground.

Specify:

the electron density at the height  $h_0$ ,  $N_0 =$  \_\_\_\_\_  $\text{cm}^{-3}$  (W101)

the reference height,  $h_0 =$  \_\_\_\_\_ km (W102)

the exponential increase of  $N$  with height,  $a =$  \_\_\_\_\_  $\text{km}^{-1}$  (W103)

	SUBROUTINE EXPX	EXPX001
C	EXPONENTIAL ELECTRON DENSITY MODEL	EXPX002
	COMMON /CONST/ PI,PIT2,PID2,DEGS,RAD,K,DUM(2)	EXPX003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	EXPX004
	COMMON R(6) /WW/ ID(10),W0,W(400)	EXPX005
	EQUIVALENCE (EARTH,W(2)),(F,W(6)),	EXPX006
	1 (N0,W(101)),(H0,W(102)),(A,W(103)),(PERT,W(150))	EXPX007
	REAL N, NO *K	EXPX008
	DATA (MODX(1)=4HEXPX),(HMAX=350.)	EXPX009
	ENTRY ELECTX	EXPX010
	H=R(1)-EARTH	EXPX011
	N=NO * EXP(A*(H-H0))	EXPX012
	X=K*N/F**2	EXPX013
	PXPR=A*X	EXPX014
	IF (PERT.NE.0.) CALL ELECT1	EXPX015
	RETURN	EXPX016
	END	EXPX017-

APPENDIX 4. PERTURBATIONS TO ELECTRON DENSITY MODELS WITH INPUT PARAMETER FORMS

The following perturbations to electron density models (irregularities) are available. The input parameter forms, which describe the perturbation, and the subroutine listings are given on the pages shown.

- a. Do-nothing perturbation (ELECT1) 126
- b. East-west irregularity with an elliptical cross-section above the equator (TORUS) 127
- c. Two east-west irregularities with elliptical cross-sections above the equator (DTORUS) 129
- d. Increase in electron density at any latitude (TROUGH) 131
- e. Increase in electron density produced by a shock wave (SHOCK) 132
- f. "Gravity-wave" irregularity (WAVE) 134
- g. "Gravity-wave" irregularity (WAVE2) 136
- h. Height profile of time derivative of electron density for calculating Doppler shift (DOPPLER) 138

To add other perturbations to electron density models the user must write a subroutine to modify the normalized electron density (X) and its gradient ( $\partial X/\partial r$ ,  $\partial X/\partial \theta$ ,  $\partial X/\partial \varphi$ ) as a function of position in spherical polar coordinates (r,  $\theta$ ,  $\varphi$ ).

The restrictions on electron density models also apply to perturbations. Again, the coordinates r,  $\theta$ ,  $\varphi$  refer to the computational coordinate system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine is through blank common (see Table 3) for the position (r,  $\theta$ ,  $\varphi$ ) and through common block /XX/ (see Table 8) for the unperturbed electron density and its gradient. The output is through common block /XX/. It is useful if the name of the subroutine suggests the perturbation model to which it corresponds. It should have an entry point ELECT1 so that it may be called by an electron density subroutine. Any parameters needed by the subroutine should be input into W151 through W199 of the W array. (See Table 2.)

If no perturbation is wanted, the following subroutine should be used.

```
C  SUBROUTINE ELECT1                                ELEC001
    USE WHEN AN ELECTRON DENSITY PERTURBATION IS NOT WANTED ELEC002
    COMMON /XX/ MODX(2),X(6)                        ELEC003
    COMMON /WW/ ID(10),W0,W(400)                   ELEC004
    EQUIVALENCE (PERT,W(150))                      ELEC005
    DATA (MODX(2)=6H NONE )                       ELEC006
    PERT=0.                                         ELEC007
    RETURN                                          ELEC008
    END                                             ELEC009-
```

## INPUT PARAMETER FORM FOR SUBROUTINE TORUS

A perturbation to an ionospheric electron density model consisting of an East-West irregularity with an elliptical cross section above the equator

$$N = N_0 (1 + \Delta)$$

$$\Delta = C_0 \exp \left\{ - \left[ \frac{(R_0 + H_0)(\theta - \pi/2) \cos \beta + (R - R_0 - H_0) \sin \beta}{A} \right]^2 - \left[ \frac{(R - R_0 - H_0) \cos \beta - (R_0 + H_0)(\theta - \pi/2) \sin \beta}{B} \right]^2 \right\}$$

$R_0$  is the radius of the earth.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

$N_0(R, \theta, \varphi)$  is any ionospheric electron density model.

Specify:

$C_0 =$  \_\_\_\_\_ . (W151)

Semi-major axis of ellipse,  $A =$  \_\_\_\_\_ km (W152)

Semi-minor axis of ellipse,  $B =$  \_\_\_\_\_ km (W153)

Tilt of ellipse,  $\beta =$  \_\_\_\_\_ degrees (W154)

Height of torus from ground,  $H_0 =$  \_\_\_\_\_ km (W155)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)



```

SUBROUTINE TORUS
COMMON /CONST/ PI,PIT2,PID2,DUM(5)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(C0,W(151)),(A,W(152)),(B,W(153)),
1 (BETA,W(154)),(H0,W(155))
REAL LAMBDA
DATA (PDPP=0.), (MODX(2)=6H TORUS)
ENTRY ELECT1
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN
IF (C0.EQ.0.) RETURN
R0=EARTH+H0
Z=R(1)-R0
LAMBDA=R0*(R(2)-PID2)
SINBET=SIN(BETA)
COSBET=COS(BETA)
P=LAMBDA*COSBET+Z*SINBET
Y=Z*COSBET-LAMBDA*SINBET
DELTA=C0*EXP(-(P/A)**2-(Y/B)**2)
DEL1=DELTA+1.
PDPR=-2.*DELTA*(P*SINBET/A**2+Y*COSBET/B**2)
PDPT=-2.*DELTA*(P*R0*COSBET/A**2-Y*R0*SINBET/B**2)
PXPR=PXPR*DEL1+X*PDPR
PXPTH=PXPTH*DEL1+X*PDPT
PXPPH=PXPPH*DEL1+X*PDPP
X=X*DEL1
RETURN
END
TOR 001
TOR 002
TOR 003
TOR 004
TOR 005
TOR 006
TOR 007
TOR 008
TOR 009
TOR 010
TOR 011
TOR 012
TOR 013
TOR 014
TOR 015
TOR 016
TOR 017
TOR 018
TOR 019
TOR 020
TOR 021
TOR 022
TOR 023
TOR 024
TOR 025
TOR 026
TOR 027
TOR 028-

```

## INPUT PARAMETER FORM FOR SUBROUTINE DTORUS

A perturbation to an ionospheric electron density model consisting of two east-west irregularities with elliptical cross sections above the equator. Since the model is expressed in spherical coordinates and does not depend on longitude, the perturbation is actually a torus circling the earth above the equator.

$$\begin{aligned}
 N &= N_0 (1 + \Delta) \\
 \Delta &= C_1 \exp \left\{ - \left[ \frac{(r_0 + H_1)(\theta - \pi/2) \cos \beta + (r - r_0 - H_1) \sin \beta}{A_1} \right]^2 \right. \\
 &\quad \left. - \left[ \frac{(r - r_0 - H_1) \cos \beta - (r_0 + H_1)(\theta - \pi/2) \sin \beta}{B_1} \right]^2 \right\} \\
 &+ C_2 \exp \left\{ - \left[ \frac{(r_0 + H_2)(\theta - \pi/2 + \delta\theta) \cos \beta + (r - r_0 - H_2) \sin \beta}{A_2} \right] \right. \\
 &\quad \left. - \left[ \frac{(r - r_0 - H_2) \cos \beta - (r_0 + H_2)(\theta - \pi/2 + \delta\theta) \sin \beta}{B_2} \right]^2 \right\}
 \end{aligned}$$

$\delta\theta$  = Northward angular displacement of the lower blob from the upper one

$$= \frac{H_1 - H_2}{\tan \beta (r_0 + H_2)}$$

$r_0$  is the radius of the earth.

$r, \theta, \varphi$  are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \varphi)$  is any electron density model.

Specify:

use perturbation \_\_\_\_\_ (W150 = 1.)

ignore perturbation \_\_\_\_\_ (W150 = 0.)

Fractional perturbation electron density at the center of the upper blob,  $C_1 =$  \_\_\_\_\_ (W151).

That of the lower blob,  $C_2 =$  \_\_\_\_\_ (W156).

Height (above ground) of the center of the upper blob,  $H_1 =$  \_\_\_\_\_ km (W155).

That of the lower blob,  $H_2 =$  \_\_\_\_\_ km (W159).

Angle (with a horizontal southward vector) of the line joining the  
blob centers,  $\theta =$  \_\_\_\_\_  $\frac{\text{rad}}{\text{deg}}$  (W154).

Semi-axis of the upper blob, to the  $1/e$  perturbation contour, in the  
direction of the line joining the blobs,  $A_1 =$  \_\_\_\_\_ km (W152).

That of the lower blob,  $A_2 =$  \_\_\_\_\_ km (W157).

Semi-axis of the upper blob in the direction normal to the line joining  
the blobs,  $B_1 =$  \_\_\_\_\_ km (W153).

That of the lower blob,  $B_2 =$  \_\_\_\_\_ km (W158).

```
SUBROUTINE DTORUS
COMMON /CONST/ PI,PIT2,PID2,DUM(5)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(C1,W(151)),(A1,W(152)),(B1,W(153)),
1 (BETA,W(154)),(H1,W(155)),(C2,W(156)),(A2,W(157)),(B2,W(158)),
2 (H2,W(159))
REAL LAMBDA1,LAMBDA2
DATA (MODX(2)=6HDTORUS),(PDPP=0.)
ENTRY ELECT1
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN
IF (C1.EQ.0.) RETURN
R1=EARTH+H1
R2=EARTH+H2
Z1=R(1)-R1
Z2=R(1)-R2
LAMBDA1=R1*(R(2)-PID2)
LAMBDA2=R2*(R(2)-PID2+(H1-H2)/R2/TANF(BETA))
SINBET=SIN(BETA)
COSBET=COS(BETA)
P1=LAMBDA1*COSBET+Z1*SINBET
P2=LAMBDA2*COSBET+Z2*SINBET
Y1=Z1*COSBET-LAMBDA1*SINBET
Y2=Z2*COSBET-LAMBDA2*SINBET
DELTA1=C1*EXP(-(P1/A1)**2-(Y1/B1)**2)
DELTA2=C2*EXP(-(P2/A2)**2-(Y2/B2)**2)
DEL1=1.+DELTA1+DELTA2
PDPR1=-2.*DELTA1*(P1*SINBET/A1**2+Y1*COSBET/B1**2)
PDPR2=-2.*DELTA2*(P2*SINBET/A2**2+Y2*COSBET/B2**2)
PDPT1=-2.*DELTA1*(P1*R1*COSBET/A1**2-Y1*R1*SINBET/B1**2)
PDPT2=-2.*DELTA2*(P2*R2*COSBET/A2**2-Y2*R2*SINBET/B2**2)
PXPR=PXPR*DEL1+X*(PDPR1+PDPR2)
PXPTH=PXPTH*DEL1+X*(PDPT1+PDPT2)
PXPPH=PXPPH*DEL1*PDPP
X=X*DEL1
RETURN
END
```

DTOR001  
DTOR002  
DTOR003  
DTOR004  
DTOR005  
DTOR006  
DTOR007  
DTOR008  
DTOR009  
DTOR010  
DTOR011  
DTOR012  
DTOR013  
DTOR014  
DTOR015  
DTOR016  
DTOR017  
DTOR018  
DTOR019  
DTOR020  
DTOR021  
DTOR022  
DTOR023  
DTOR024  
DTOR025  
DTOR026  
DTOR027  
DTOR028  
DTOR029  
DTOR030  
DTOR031  
DTOR032  
DTOR033  
DTOR034  
DTOR035  
DTOR036  
DTOR037-

## INPUT PARAMETER FORM FOR SUBROUTINE TROUGH

A perturbation to an ionospheric electron density model consisting of an increase in electron density near any latitude

$$N = (1 + \Delta) N_0 (R, \theta, \varphi) \qquad W = B \text{ for } \frac{\pi}{2} - \theta - \lambda \geq 0$$

$$\Delta = A \exp\left(-\left(\frac{\pi/2 - \theta - \lambda}{W}\right)^2\right) \qquad W = B \times C \text{ for } \frac{\pi}{2} - \theta - \lambda < 0$$

$N_0 (R, \theta, \varphi)$  is any ionospheric electron density model.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

Specify:

Amplitude of the perturbation,  $A =$  \_\_\_\_\_ (W151)

half width of the perturbation,  $B =$  \_\_\_\_\_ degrees (W152)

latitude of the perturbation,  $\lambda =$  \_\_\_\_\_ degrees (W153)

width factor for South of trough,  $C =$  \_\_\_\_\_ (W154)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

C	<pre> SUBROUTINE TROUGH       A PERTURBATION TO AN ELECTRON DENSITY MODEL COMMON /CONST/ PI,PIT2,PID2,DUM(5) COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX COMMON R(6) /WW/ ID(10),W0,W(400) EQUIVALENCE (A,W(151)),(B,W(152)),(ALAT,W(153)),(FACTOR,W(154)) DATA (MODX(2)=6HTROUGH) ENTRY ELECT1 IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN IF (A.EQ.0.) RETURN ANGLE=R(2)+ALAT-PID2 WIDTH=B IF (ANGLE.GT.0.) WIDTH=FACTOR*B ANGLE=ANGLE/WIDTH DELTA=A*EXP(-ANGLE**2) DEL1=DELTA+1. PXPR=PXPR*DEL1 XPPTH=XPPTH*DEL1-2.*X*ANGLE*DELTA/WIDTH PXPPH=PXPPH*DEL1 X=X*DEL1 RETURN       END                 </pre>	<pre> TROU001 TROU002 TROU003 TROU004 TROU005 TROU006 TROU007 TROU008 TROU009 TROU010 TROU011 TROU012 TROU013 TROU014 TROU015 TROU016 TROU017 TROU018 TROU019 TROU020 TROU021 TROU022-                 </pre>
---	--	---

## INPUT PARAMETER FORM FOR SUBROUTINE SHOCK

A perturbation to an ionospheric electron density model consisting of an increase in electron density produced by a shock wave

$$N(R, \theta, \varphi) = N_0(R, \theta, \varphi) \left[ 1 + P \exp\left(-9 \left(\frac{\rho_c - \rho}{w}\right)^2\right) \right]$$

$$\rho_c = s(h - h_0) - w$$

$$\rho = R \left| \cos^{-1} \left[ \cos(\varphi - \varphi_0) \cos(\lambda - \lambda_0) \right] \right|$$

$N_0(R, \theta, \varphi)$  is the ambient electron density specified by any electron density model.

$R, \theta, \varphi$  give the position in spherical polar coordinates.

$h = R - a$  is the height above the surface of the earth.

$a$  is the radius of the earth.

$\lambda = \frac{\pi}{2} - \theta$  is the latitude.

Specify:

Relative increase in electron density,  $P =$  \_\_\_\_\_ (W151).

Width of the disturbance,  $w =$  \_\_\_\_\_ km (W152).

Latitude of the center of the disturbance,  $\lambda_0 =$  \_\_\_\_\_ radians or degrees (W153).

Longitude of the center of the disturbance,  $\varphi_0 =$  \_\_\_\_\_ radians or degrees (W154).

Slope measured from vertical - rate of increase of  $\rho_c$  with height,  $s =$  \_\_\_\_\_ (W155).

Height to the bottom of the disturbance,  $h_0 =$  \_\_\_\_\_ km (W156).

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

```

SUBROUTINE SHOCK
C A PERTURBATION TO AN ELECTRON DENSITY MODEL SIMULATING A SHOCK WAVE
COMMON /CONST/ PI,PID2,PID2,DUM(5)
COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX
COMMON R(6) /WW/ ID(10),W0,W(400)
EQUIVALENCE (EARTH,W(2)),(P,W(151)),(WW,W(152)),(ALAT,W(153)),
I (ALON,W(154)),(S,W(155)),(H0,W(156))
REAL LAT,LON
DATA (MODX(2)=6H SHOCK)
ENTRY ELECT1
IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN
IF (P.EQ.0..OR.WW.EQ.0.) RETURN
H=R(1)-EARTH
RHOC=S*(H-H0)-WW
LON=R(3)-ALON
LAT=PID2-R(2)-ALAT
COSLON=COS(LON)
COSLAT=COS(LAT)
U=COSLON*COSLAT
RHO=R(1)*ACOS(U)
DIF=RHOC-RHO
CON=-9./WW**2
CONS=P*EXP(CON*DIF**2)
CONST=1.+CONS
CON=2.*CON*CONS*DIF
PXPR=PXPR*CONST+X*CON*(S-RHO/R(1))
CONS=R(1)*(1./SQRT(1.-U**2))
PXPTH=PXPTH*CONST+X*CON*CONS*COSLON*SIN(LAT)
PXPPH=PXPPH*CONST-X*CON*CONS*COSLAT*SIN(LON)
X=X*CONST
RETURN
END

```

SHOC001  
SHOC002  
SHOC003  
SHOC004  
SHOC005  
SHOC006  
SHOC007  
SHOC008  
SHOC009  
SHOC010  
SHOC011  
SHOC012  
SHOC013  
SHOC014  
SHOC015  
SHOC016  
SHOC017  
SHOC018  
SHOC019  
SHOC020  
SHOC021  
SHOC022  
SHOC023  
SHOC024  
SHOC025  
SHOC026  
SHOC027  
SHOC028  
SHOC029  
SHOC030  
SHOC031  
SHOC032-

## INPUT PARAMETER FORM FOR SUBROUTINE WAVE

A perturbation to an ionospheric electron density model consisting of a "gravity-wave" irregularity traveling from north pole to south pole

$$N = N_0(1 + \Delta)$$

$$\Delta = \delta \exp\left\{-\left[(R - R_0 - z_0)/H\right]^2\right\} \cdot \cos\left\{2\pi\left[t' + (\pi/2 - \theta)\frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z\right]\right\}$$

$$\frac{\partial N}{\partial t} = \frac{-2\pi}{\lambda_x} V_x N_0 \delta \exp - \left[(R - R_0 - z_0)/H\right]^2 \cdot \sin 2\pi\left[t' + (\pi/2 - \theta)\frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z\right]$$

$R_0$  is the radius of the earth.

$R$ ,  $\theta$ ,  $\varphi$  are the spherical (earth-centered) polar coordinates  
( $\Delta$  is independent of  $\varphi$ ).

$N_0(R, \theta, \varphi)$  is any electron density model.

Specify:

the height of maximum wave amplitude,  $z_0 = \underline{\hspace{2cm}}$  km (W151)

wave-amplitude "scale height,"  $H = \underline{\hspace{2cm}}$  km (W152)

wave perturbation amplitude,  $\delta = \underline{\hspace{2cm}}$  [0. to 1.] (W153)

horizontal trace velocity,  $V_x = \underline{\hspace{2cm}}$  km/sec (W154)  
(needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x = \underline{\hspace{2cm}}$  km (W155)

vertical wavelength,  $\lambda_z = \underline{\hspace{2cm}}$  km (W156)

time in wave periods,  $t' = \underline{\hspace{2cm}}$  [0. to 1.] (W157)

(W150: = 1. to use perturbation, = 0. to ignore perturbation)

	SUBROUTINE WAVE	WAVE001
C	PERTURBATION TO AN ALPHA-CHAPMAN ELECTRON DENSITY MODEL	WAVE002
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	WAVE003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT,HMAX	WAVE004
	COMMON R(6) /WW/ ID(10),W0,W(400)	WAVE005
	EQUIVALENCE (EARTH,R,W(2)),(Z0,W(151)),(SH,W(152)),(DELTA,W(153)),	WAVE006
	1 (VSUBX,W(154)),(LAMBDAZ,W(155)),(LAMBDAZ,W(156)),(TP,W(157))	WAVE007
	REAL LAMBDAZ,LAMBDAZ	WAVE008
	DATA (MODX(2)=6H WAVE )	WAVE009
	ENTRY ELECT1	WAVE010
	IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN	WAVE011
	IF (DELTA.EQ.0..OR.SH.EQ.0.) RETURN	WAVE012
	H=R(1)-EARTH	WAVE013
	EXPQ=EXP(-((H-Z0)/SH)**2)	WAVE014
	TMP=PIT2*(TP+(PID2-R(2))*EARTH/LAMBDAZ+H/LAMBDAZ)	WAVE015
	SINW=SIN(TMP)	WAVE016
	COSW=SIN(PID2-TMP)	WAVE017
	CONS=1.0+DELTA*EXPQ*COSW	WAVE018
	IF (H.NE.0.) PXPR=PXPR*CONS-X*DELTA*EXPQ*(2.0/SH**2*(H-Z0)*COSW	WAVE019
	1 +PIT2/LAMBDAZ*SINW)	WAVE020
	PXPTH=PXPTH*CONS+X*DELTA*PIT2*EARTH/LAMBDAZ*SINW*EXPQ	WAVE021
	PXPPH=PXPPH*CONS	WAVE022
	PXPT=0.	WAVE023
	IF (VSUBX.NE.0.) PXPT=-PIT2*VSUBX/LAMBDAZ*X*DELTA*EXPQ*SINW	WAVE024
	X=X*CONS	WAVE025
	RETURN	WAVE026
	END	WAVE027-



INPUT PARAMETER FORM FOR SUBROUTINE WAVE 2  
 PERTURBATION TO AN IONOSPHERIC ELECTRON DENSITY MODEL

A "gravity-wave" irregularity traveling from north pole to south pole - same as WAVE 1, but with Gaussian amplitude variations in latitude and longitude, and provision for a horizontal "group velocity "

$$N = N_0 (1 + AC)$$

$$A = \delta \exp - \left( \frac{r - r_0 - z_0}{H} \right)^2 \cdot \exp - \left( \frac{\theta - \theta_0(t)}{\Theta} \right)^2 \cdot \exp - \left( \frac{\phi}{\Phi} \right)^2$$

$$C = \cos 2\pi \left[ t' + (\pi/2 - \theta) \frac{r_0}{\lambda_x} + (r - r_0)/\lambda_z \right]$$

$$\theta_0 = \theta_\infty + V_g t / r_0$$

$r_0$  is the radius of the earth.

$r, \theta, \phi$  are spherical (earth-centered) polar coordinates.

$N_0(r, \theta, \phi)$  is any electron density model.

Specify:

use perturbation \_\_\_\_\_ (W150 = 1.)

ignore perturbation \_\_\_\_\_ (W150 = 0.)

the height of maximum wave amplitude,  $z_0 =$  \_\_\_\_\_ km (W151)

wave-amplitude "scale height,"  $H =$  \_\_\_\_\_ km (W152)

wave perturbation amplitude,  $\delta =$  \_\_\_\_\_ (0 to 1) (W153)

horizontal trace velocity,  $V_x =$  \_\_\_\_\_ km/sec. (W154)  
 (needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x =$  \_\_\_\_\_ km (W155)

vertical wavelength,  $\lambda_z =$  \_\_\_\_\_ km (W156)

time in wave periods,  $t' =$  \_\_\_\_\_ (W157)

amplitude "scale distance" in latitude,  $\Theta =$  \_\_\_\_\_ degrees (W159)

amplitude "scale distance" in longitude,  $\Phi =$  \_\_\_\_\_ degrees (W160)

latitude of maximum amplitude at  $t = 0$ ,  $\theta_\infty =$  \_\_\_\_\_ degrees (W158)

southward group velocity,  $V_g =$  \_\_\_\_\_ km/sec (W161)  
 (needed even if Doppler shift is not calculated)

	SUBROUTINE WAVE2	WAV2001
C	PERTURBATION TO AN ANY ELECTRON DENSITY MODEL	WAV2002
	COMMON /CONST/ PI,PIT2,PID2,DUM(5)	WAV2003
	COMMON /XX/ MODX(2),X,PXPR,PXPTH,PXPPH,PXPT	WAV2004
	COMMON R(6) /WW/ ID(10),WO,W(400)	WAV2005
	EQUIVALENCE (EARTH,W(2)),(Z0,W(151)),(SH,W(152)),(DELTA,W(153)),	WAV2006
	1 (VSUBX,W(154)),(LAMBDA,W(155)),(LAMBDAZ,W(156)),(TP,W(157)),	WAV2007
	2 (TH0,W(158)),(THC,W(159)),(PHIC,W(160)),(VGX,W(161))	WAV2008
	REAL LAMBDA,W(155),LAMBDAZ	WAV2009
	DATA (MODX(2)=6H WAVE2)	WAV2010
	ENTRY ELECT1	WAV2011
	IF (X.EQ.0..AND.PXPR.EQ.0..AND.PXPTH.EQ.0..AND.PXPPH.EQ.0.) RETURN	WAV2012
	IF (DELTA.EQ.0..OR.SH.EQ.0.) RETURN	WAV2013
	H=R(1)-EARTH	WAV2014
	TH0=TH0+LAMBDA*TP*VGX/VSUBX/EARTH	WAV2015
	EXPR=EXP(-((H-Z0)/SH)**2)	WAV2016
	EXPTH=EXP(-((R(2)-TH0)/THC)**2)	WAV2017
	EXPPHI=EXP(-((R(3)-PHIC)**2)	WAV2018
	WW=PIT2*(TP+(PID2-R(2))*EARTH/LAMBDA+H/LAMBDAZ)	WAV2019
	SINW=SIN(WW)	WAV2020
	COSW=COS(WW)	WAV2021
	E=DELTA*EXPR*EXPTH*EXPPHI	WAV2022
	CONS=1.0+E*COSW	WAV2023
	PXPR=PXPR*CONS-X*E*2.*(COSW*(H-Z0)/SH**2+PI/LAMBDAZ*SINW)	WAV2024
	PXPTH=PXPTH*CONS+2.*E*(X*PI*EARTH*SINW/LAMBDA-(R(2)-TH0)/	WAV2025
	1 THC**2*COSW)	WAV2026
	PXPPH=PXPPH*CONS-X*2.*E*R(3)/PHIC**2*COSW	WAV2027
	PXPT=-PIT2*VSUBX*E/LAMBDA*SINW+2.0*E*VGX/EARTH*COSW*(R(2)-TH0	WAV2028
	1 -LAMBDA*TP/EARTH)/THC	WAV2029
	X=X*CONS	WAV2030
	RETURN	WAV2031
	END	WAV2032-



```

SUBROUTINE DOPPLER                                DOPP001
C   COMPUTES DN/DT                                FROM PROFILES HAVING THE SAME FORM DOPP002
C   AS THOSE USED BY SUBROUTINE TABLE X          DOPP003
C   MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS DOPP004
C   NEEDS SUBROUTINE GAUSEL                        DOPP005
DIMENSION HPC(250),FN2C(250),ALPHA(250),BETA(250),GAMMA(250),
1  DELTA(250),SLOPE(250),MAT(4,5)                DOPP006
COMMON /CONST/ PE,PIT2,PI02,DEGS,RAD,K,DUM(2)     DOPP008
COMMON /XX/ MODX(2),XDUM,PXPR,PXPTH,PXPPH,X,HMAX   DOPP009
COMMON R(6) /HW/ ID(10),W0,W(400)                 DOPP010
EQUIVALENC (EARTH, W(2)), (F, W(6)), (READFN, W(151)) DOPP011
REAL MAT, K                                        DOPP012
DATA (MODX(2)=7HDOPPLER)                          DOPP013
ENTRY ELECT1                                       DOPP014
IF (READFN.EQ.0.) GO TO 10                          DOPP015
READFN=0.                                          DOPP016
READ 1000, NOC, (HPC(I),FN2C(I),I=1,NOC)           DOPP017
1000 FORMAT (I4/(F8.2,E12.4))                      DOPP018
PRINT 1200, (HPC(I),FN2C(I),I=1,NOC)              DOPP019
1200 FORMAT(1H1,14X,6HHEIGHT,4X,16+ DN/DT /((1X,F20.10,E20.1))) DOPP020
A=0.                                               DOPP021
IF (FN2C(1).NE.0.) A=ALOG (FN2C(2)/FN2C(1)) / (HPC(2)-HPC(1)) DOPP022
FN2C(1)=K*FN2C(1)                                 DOPP023
FN2C(2)=K*FN2C(2)                                 DOPP024
SLOPE(1)=A*FN2C(1)                                DOPP025
SLOPE(NOC)=0.                                     DOPP026
VMAX=1                                             DOPP027
DO 6 I=2,NOC                                       DOPP028
IF (FN2C(I).GT.FN2C(NMAX)) NMAX=I                 DOPP029
IF (I.EQ.NOC) GO TO 4                              DOPP030
FN2C(I+1)=K*FN2C(I+1)                             DOPP031
DO 3 J=1,3                                         DOPP032
M=I+J-2                                            DOPP033
MAT(J,1)=1.                                        DOPP034
MAT(J,2)=HPC(M)                                    DOPP035
MAT(J,3)=HPC(M)**2                                DOPP036
MAT(J,4)=FN2C(M)                                  DOPP037
3  CALL GAUSEL (MAT,4,3,4,NRANK)                   DOPP038
IF (NRANK.LT.3) GO TO 60                           DOPP039
SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)              DOPP040
4  DO 5 J=1,2                                       DOPP041
M=I+J-2                                            DOPP042
MAT(J,1)=1.                                        DOPP043
MAT(J,2)=HPC(M)                                    DOPP044
MAT(J,3)=HPC(M)**2                                DOPP045
MAT(J,4)=HPC(M)**3                                DOPP046
MAT(J,5)=FN2C(M)                                  DOPP047
L=J+2                                              DOPP048
MAT(L,1)=0.                                        DOPP049
MAT(L,2)=1.                                        DOPP050
MAT(L,3)=2.*HPC(M)                                DOPP051
MAT(L,4)=3.*HPC(M)**2                             DOPP052
5  MAT(L,5)=SLOPE(M)                               DOPP053
CALL GAUSEL (MAT,4,4,5,NRANK)                     DOPP054
IF (NRANK.LT.4) GO TO 60                           DOPP055
ALPHA(I)=MAT(1,5)                                  DOPP056
BETA(I)=MAT(2,5)                                   DOPP057
GAMMA(I)=MAT(3,5)                                  DOPP058
6  DELTA(I)=MAT(4,5)                                DOPP059
HMAX=AMAX1(HMAX,HPC(NMAX))                          DOPP060
NH=2                                               DOPP061
10 H=R(1)-EARTH                                     DOPP062
F2=F*F                                             DOPP063
IF (H.GE.HPC(1)) GO TO 12                          DOPP064
11 NH=2                                             DOPP065

```

IF(FN2C(1).EQ.0.) GO TO 50	DOPP066
X=FN2C(1)*EXP(A*(H-HPC(1)))/F2	DOPP067
GO TO 50	DOPP068
12 IF (H.GE.HPC(NOC)) GO TO 18	DOPP069
NSTEP=1	DOPP070
IF (H.LT.HPC(NH-1)) NSTEP=-1	DOPP071
15 IF (HPC(NH-1).LE.H.AND.H.LT.HPC(NH)) GO TO 16	DOPP072
NH=NH+NSTEP	DOPP073
GO TO 15	DOPP074
16 X=(ALPHA(NH)+H*(BETA(NH)+H*(GAMMA(NH)+H*DELTA(NH)))/F2	DOPP075
GO TO 50	DOPP076
18 X=FN2C(NOC)/F2	DOPP077
50 CONTINUE	DOPP078
RETURN	DOPP079
60 PRINT 6000, I,HPC(I)	DOPP080
6000 FORMAT(4H THE,I4,55TH POINT IN THE DN/DT PROFILE HAS T	DOPP081
THE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)	DOPP082
CALL EXIT	DOPP083
END	DOPP084
	DOPP085

## APPENDIX 5. MODELS OF THE EARTH'S MAGNETIC FIELD WITH INPUT PARAMETER FORMS

The following models of the earth's magnetic field are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Constant dip and gyrofrequency (CONSTY)	142
b.	Earth-centered dipole (DIPOLY)	143
c.	Constant dip. Gyrofrequency varies as the inverse cube of the distance from the center of the earth (CUBEY)	144
d.	Spherical harmonic expansion (HARMONY)	145

To add other models of the earth's magnetic field the user must write a subroutine that will calculate the normalized strength and direction of the earth's magnetic field ( $Y$ ,  $Y_r$ ,  $Y_\theta$ ,  $Y_\varphi$ ) and their gradients ( $\partial Y/\partial r$ ,  $\partial Y/\partial \theta$ ,  $\partial Y/\partial \varphi$ ,  $\partial Y_r/\partial r$ ,  $\partial Y_r/\partial \theta$ ,  $\partial Y_r/\partial \varphi$ ,  $\partial Y_\theta/\partial r$ ,  $\partial Y_\theta/\partial \theta$ ,  $\partial Y_\theta/\partial \varphi$ ,  $\partial Y_\varphi/\partial r$ ,  $\partial Y_\varphi/\partial \theta$ ,  $\partial Y_\varphi/\partial \varphi$ ) as a function of position in spherical polar coordinates ( $r$ ,  $\theta$ ,  $\varphi$ ). ( $Y = f_H / f$ , where  $f_H$  is the electron gyrofrequency and  $f$  is the wave frequency.)

The restrictions on electron density models also apply to models of the earth's magnetic field. The coordinates  $r$ ,  $\theta$ ,  $\varphi$  refer to the computational coordinate system, which is not necessarily the same as geographic coordinates. W24 and W25 give the geographic latitude and longitude of the north pole of the computational coordinate system.

The input to the subroutine ( $r$ ,  $\theta$ ,  $\varphi$ ) is through blank common. (See Table 3.) The output is through common block /YY/. (See Table 9.) It is useful if the name of the subroutine suggests the model to which it corresponds. It should have an entry point MAGY so that other subroutines in the program can call it. Any parameters needed by the subroutine should be input into W201 through W249 of the W array. (See Table 2.) If the subroutine needs massive amounts of data, these should be read in by the subroutine following the example of subroutine HARMONY.







## INPUT PARAMETER FORM FOR SUBROUTINE CUBEY

A model of the earth's magnetic field consisting of a constant dip and a gyrofrequency which varies as the inverse cube of the distance from the center of the earth

This model has the same height variation as a dipole magnetic field.

The gyrofrequency is given by:

$$f_H = f_{H_0} \left( \frac{a}{r} \right)^3$$

a is the radius of the earth.

r is the distance from the center of the earth.

Specify:

gyrofrequency at the ground,  $f_{H_0} =$  \_\_\_\_\_ MHz (W201)

dip, I = \_\_\_\_\_ radians  
degrees (W202)

The magnetic meridian is defined by the geographic coordinates of the north magnetic pole:

latitude = \_\_\_\_\_ radians  
degrees north (W24)  
km

longitude = \_\_\_\_\_ radians  
degrees east (W25)  
km

	SUBROUTINE CUBEY	CUBE001
C	CONSTANT DIP.	CUBE002
C	GYROFREQ DECREASES AS CUBE OF DISTANCE FROM CENTER OF EARTH.	CUBE003
C	THIS MODEL HAS SAME HEIGHT VARIATION AS A DIPOLE FIELD.	CUBE004
	COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRC	CUBE005
	1,PYTPT,PYTTP,YPH,PYPPR,PYPPT,PYPPP	CUBE006
	COMMON R /WW/ ID(10),W0,W(400)	CUBE007
	EQUIVALENCE (EARTH,R,W(2)),(F,W(6)),(FH,W(201)),(DIP,W(202))	CUBE008
	DATA (MODY=5HCUBEY)	CUBE009
	ENTRY MAGY	CUBE010
	Y=(EARTH/R)**3 *FH/F	CUBE011
	YR= Y*SIN(DIP)	CUBE012
	YTH= Y*COS(DIP)	CUBE013
	PYPR=-3.*Y/R	CUBE014
	PYRPR=-3.*YR/R	CUBE015
	PYTPR=-3.*YTH/R	CUBE016
	RETURN	CUBE017
	END	CUBE018-

## INPUT PARAMETER FORM FOR SUBROUTINE HARMONY

A model of the earth's magnetic field based on a spherical harmonic expansion

The upward, southerly, and easterly components of the earth's magnetic field are given by:

$$H_r = - \sum_{n=0}^6 (n+1) \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n H_n^m(\theta) \left( g_n^m \cos m\varphi + h_n^m \sin m\varphi \right)$$

$$H_\theta = - \frac{1}{\sin \theta} \sum_{n=0}^6 \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n G_n^m(\theta) \left( g_n^m \cos m\varphi + h_n^m \sin m\varphi \right)$$

$$H_\varphi = \frac{1}{\sin \theta} \sum_{n=0}^6 \left(\frac{a}{r}\right)^{n+2} \sum_{m=0}^n m H_n^m(\theta) \left( h_n^m \cos m\varphi - g_n^m \sin m\varphi \right)$$

where

$a$  is the radius of the earth.

$r, \theta, \varphi$  are spherical (earth-centered) polar coordinates.

$$H_0^0(\theta) = 1$$

$$H_1^0(\theta) = \cos \theta$$

$$H_1^1(\theta) = \sin \theta$$

$$H_{m+1}^m(\theta) = H_m^m(\theta) \cos \theta$$

$$H_{m+1}^{m+1}(\theta) = H_m^m(\theta) \sin \theta$$

$$H_{n+2}^m(\theta) = H_{n+1}^m(\theta) \cos \theta - \frac{(n+m+1)(n-m+1)}{(2n+3)(2n+1)} H_n^m(\theta)$$

$$G_n^m(\theta) = - \frac{d}{d\theta} H_n^m(\theta) \sin \theta$$

$$G_m^m(\theta) = -mH_m^m(\theta) \cos\theta$$

$$G_{n+1}^m(\theta) = -(n+1)H_{n+1}^m(\theta)\cos\theta + \frac{(n+m+1)(n-m+1)}{2n+1} H_n^m(\theta)$$

The recursion formulas for calculating  $H_n^m(\theta)$  and  $G_n^m(\theta)$  are from Eckhouse (1964).

This subroutine uses coefficients  $g_n^m$  and  $h_n^m$  for Gauss normalization. Some coefficients are now being published for Schmidt normalization (e. g. Cain and Sweeney, 1970). The factors  $S_{n,m}$  used for converting the "Schmidt normalized" coefficients to the "Gauss normalized" coefficients are as follows (Cain, et. al., 1968, Chapman and Bartels, 1940):

$$S_{0,0} = -1$$

$$S_{n,0} = S_{n-1,0} \left[ \frac{2n-1}{n} \right]$$

$$S_{n,1} = S_{n,0} \sqrt{\frac{2n}{n+1}}$$

$$S_{n,m} = S_{n,m-1} \sqrt{\frac{(n-m+1)}{n+m}} \quad \text{for } m > 1$$

By convention, the "Gauss normalized" coefficient  $g_1^0$  is positive, whereas the "Schmidt normalized" coefficient  $g_1^0$  is negative. Coefficients based on more recent data on the earth's magnetic field including more satellite data are in the POGO 8/69 model.

Specify below the Gauss coefficients  $g_n^m$  and  $h_n^m$  in gauss.

	columns 2 → 10	columns 11 → 20	columns 21 → 30	columns 31 → 40	columns 41 → 50	columns 51 → 60	columns 61 → 70
1st card	$g_0^0 =$ _____						
2nd card	$g_1^0 =$ _____	$g_1^1 =$ _____					
3rd card	$g_2^0 =$ _____	$g_2^1 =$ _____	$g_2^2 =$ _____				
4th card	$g_3^0 =$ _____	$g_3^1 =$ _____	$g_3^2 =$ _____	$g_3^3 =$ _____			
5th card	$g_4^0 =$ _____	$g_4^1 =$ _____	$g_4^2 =$ _____	$g_4^3 =$ _____	$g_4^4 =$ _____		
6th card	$g_5^0 =$ _____	$g_5^1 =$ _____	$g_5^2 =$ _____	$g_5^3 =$ _____	$g_5^4 =$ _____	$g_5^5 =$ _____	
7th card	$g_6^0 =$ _____	$g_6^1 =$ _____	$g_6^2 =$ _____	$g_6^3 =$ _____	$g_6^4 =$ _____	$g_6^5 =$ _____	$g_6^6 =$ _____

8th card  $h_0^0 = \underline{\hspace{2cm}}$   
 9th card  $h_1^0 = \underline{\hspace{2cm}}$   $h_1^1 = \underline{\hspace{2cm}}$   
 10th card  $h_2^0 = \underline{\hspace{2cm}}$   $h_2^1 = \underline{\hspace{2cm}}$   $h_2^2 = \underline{\hspace{2cm}}$   
 11th card  $h_3^0 = \underline{\hspace{2cm}}$   $h_3^1 = \underline{\hspace{2cm}}$   $h_3^2 = \underline{\hspace{2cm}}$   $h_3^3 = \underline{\hspace{2cm}}$   
 12th card  $h_4^0 = \underline{\hspace{2cm}}$   $h_4^1 = \underline{\hspace{2cm}}$   $h_4^2 = \underline{\hspace{2cm}}$   $h_4^3 = \underline{\hspace{2cm}}$   $h_4^4 = \underline{\hspace{2cm}}$   
 13th card  $h_5^0 = \underline{\hspace{2cm}}$   $h_5^1 = \underline{\hspace{2cm}}$   $h_5^2 = \underline{\hspace{2cm}}$   $h_5^3 = \underline{\hspace{2cm}}$   $h_5^4 = \underline{\hspace{2cm}}$   $h_5^5 = \underline{\hspace{2cm}}$   
 14th card  $h_6^0 = \underline{\hspace{2cm}}$   $h_6^1 = \underline{\hspace{2cm}}$   $h_6^2 = \underline{\hspace{2cm}}$   $h_6^3 = \underline{\hspace{2cm}}$   $h_6^4 = \underline{\hspace{2cm}}$   $h_6^5 = \underline{\hspace{2cm}}$   $h_6^6 = \underline{\hspace{2cm}}$

Set W200 = 1. to read in a set of coefficients.

This subroutine represents:

$H_n^m(\theta)$  by  $H(m+1, n+1)$

$G_n^m(\theta)$  by  $G(m+1, n+1)$

$g_n^m$  by  $GG(m+1, n+1)$

$h_n^m$  by  $HH(m+1, n+1)$

```

SUBROUTINE HARMONY
C MODEL OF THE EARTH S MAGNETIC FIELD BASED ON A HARMONIC ANALYSIS
  DIMENSION PHPTH(7,7),PGPTH(7,7),A1(7,7),B1(7,7)
  DIMENSION H(7,7),G(7,7),GG(7,7),HH(7,7),SINP(7),COSP(7)
  COMMON /YY/ MODY,Y,PYPR,PYPTH,PYPPH,YR,PYRPR,PYRPT,PYRPP,YTH,PYTPRHARM005
  1,PYTPT,PYTPP,YPH,PYPPR,PYPPT,PYPPP
  COMMON R(6) /WW/ ID(10),W0,W(400)
  COMMON /CONST/ PI,PIT2,PID2,DUM(5)
  EQUIVALENCE (THETA,R(2)),(PHI,R(3))
  EQUIVALENCE (EARTH,R(W(2))),(F,W(6)),(READFH,W(200))
C RATIO OF CHARGE TO MASS FOR ELECTRON
  DATA(EOM=1.7589E7)
  DATA (SET=0.),(H=1.,48(0.)),(G=49(0.)),(PHPTH=49(0.))
  1,(PGPTH=49(0.)),(MODY=7HHARMONY)
  ENTRY MAGY
  IF(SET) GO TO 2
  DO 1 M=1,7
  DO 1 N=1,7
  B1(M,N)=(N+M-1)*(N-M+1)/(2*N-1.)
  1 A1(M,N)=B1(M,N)/(2*N+1)
  SET=1.
HARM001
HARM002
HARM003
HARM004
HARM006
HARM007
HARM008
HARM009
HARM010
HARM011
HARM012
HARM013
HARM014
HARM015
HARM016
HARM017
HARM018
HARM019
HARM020
HARM021

```

```

2 IF(READFH.EQ.0.) GO TO 3
READ 2000,GG,HH
2000 FORMAT(1X,F9.4,6F10.4)
PRINT 2100,GG
2100 FORMAT(1H1,10X,1H0,14X,1H1,14X,1H2,14X,1H3,14X,1H4,14X,1H5,14X,1H6
1 /9X,7(1HG,14X)/10X,7(1HN,14X)//(1X,7F15.6))
PRINT 2200,HH
2200 FORMAT(// 11X,1H0,14X,1H1,14X,1H2,14X,1H3,14X,1H4,14X,1H5,14X,1H6
1 /9X,7(1HH,14X)/10X,7(1HN,14X)//(1X,7F15.6))
READFH=0.
3 COSTHE=COS(THETA)
SINTHE=SIN(THETA)
AOR=EARTH/R(1)
PAORPR=-AOR/R(1)
CNST2=AOR
PCNSPR=PAORPR
FIN1=PFIN1R=PFIN1T=PFIN1P=0.
FIN2=PFIN2R=PFIN2T=PFIN2P=0.
FIN3=PFIN3R=PFIN3T=PFIN3P=0.
DO 4 M=1,7
SINP(M)=SIN((M-1)*PHI)
4 COSP(M)=COS((M-1)*PHI)
H(1,2)=COSTHE
H(2,2)=SINTHE
DO 5 M=1,5
H(M+1,M+2)=COSTHE*H(M+1,M+1)
H(M+2,M+2)=SINTHE*H(M+1,M+1)
DO 5 N=M,5
5 H(M,N+2)=COSTHE*H(M,N+1)-A1(M,N)*H(M,N)
DO 6 M=1,6
G(M+1,M+1)=-M*COSTHE*H(M+1,M+1)
PHPTH(M+1,M+1)=-G(M+1,M+1)/SINTHE
PGPTH(M+1,M+1)=M*SINTHE*H(M+1,M+1)-M*COSTHE*PHPTH(M+1,M+1)
DO 6 N=M,6
G(M,N+1)=-N*COSTHE*H(M,N+1)+B1(M,N)*H(M,N)
PHPTH(M,N+1)=-G(M,N+1)/SINTHE
6 PGPTH(M,N+1)=N*SINTHE*H(M,N+1)-N*COSTHE*PHPTH(M,N+1)+B1(M,N)*PHPTH
1 (M,N)
DO 8 N=1,7
CR=PCRPTH=PCRPPH=0.
CTH=PCTHPT=PCTHPP=0.
CPH=PCPHPT=PCPHPP=0.
DO 7 M=1,N
TEMP1=GG(M,N)*COSP(M)+HH(M,N)*SINP(M)
TEMP2=(M-1)*(HH(M,N)*COSP(M)-GG(M,N)*SINP(M))
CR =CR +H(M,N)*TEMP1
PCRPTH=PCRPTH+PHPTH(M,N)*TEMP1
PCRPPH=PCRPPH+H(M,N)*TEMP2
CTH =CTH +G(M,N)*TEMP1
PCTHPT=PCTHPT+PGPTH(M,N)*TEMP1
PCTHPP=PCTHPP+G(M,N)*TEMP2
CPH =CPH +H(M,N)*TEMP2
PCPHPT=PCPHPT+PHPTH(M,N)*TEMP2
7 PCPHPP=PCPHPP-H(M,N)*(M-1)**2*TEMP1
CNST2=CNST2*AOR
PCNSPR=CNST2*PAORPR+AOR*PCNSPR
FIN1=FIN1+N*CNST2*CR
PFIN1R=PFIN1R+N*PCNSPR*CR
PFIN1T=PFIN1T+N*CNST2*PCRPTH
PFIN1P=PFIN1P+N*CNST2*PCRPPH
FIN2=FIN2+CNST2*CTH
PFIN2R=PFIN2R+PCNSPR*CTH
PFIN2T=PFIN2T+CNST2*PCTHPT
PFIN2P=PFIN2P+CNST2*PCTHPP
FIN3=FIN3+CNST2*CPH

```

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HARM022
HARM023
HARM024
HARM025
HARM026
HARM027
HARM028
HARM029
HARM030
HARM031
HARM032
HARM033
HARM034
HARM035
HARM036
HARM037
HARM038
HARM039
HARM040
HARM041
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HARM075
HARM076
HARM077
HARM078
HARM079
HARM080
HARM081
HARM082
HARM083
HARM084
HARM085
HARM086

```

```

      PFIN3R=PFIN3R+PCNSPR*CPH
      PFIN3T=PFIN3T+CNST2*PCPHPT
      PFIN3P=PFIN3P+CNST2*PCPHP>
      4THETA=-FIN2/SINTE
      HPHI=FIN3/SINTE
C***** CONVERT FROM MAG FIELD IN GAUSS TO GYROFREQ IN MHZ
      CONST=-EOM/PIT2*1.E-6/F
      YR=-CONST*FIN1
      YTH=CONST*HTHETA
      YPH=CONST*HPHI
      Y=SQRT(YR**2+YTH**2+YPH**2)
      PYRPR=-CONST*PFIN1R
      PYTPR=-CONST*PFIN2R/SINTE
      PYPPR=CONST*PFIN3R/SINTE
      PYPR=(YR*PYRPR+YTH*PYTPR+YPH*PYPPR)/Y
      PYRPT=-CONST*PFIN1T
      PYTPT=-CONST*(PFIN2T/SINTE+HTHETA*COSTHE/SINTE)
      PYPPT=CONST*(PFIN3T/SINTE-HPHI*COSTHE/SINTE)
      PYPTH=(YR*PYRPT+YTH*PYTPT+YPH*PYPPT)/Y
      PYRPP=-CONST*PFIN1P
      PYTPP=-CONST*PFIN2P/SINTE
      PYPPP=CONST*PFIN3P/SINTE
      PYPHP=(YR*PYRPP+YTH*PYTPP+YPH*PYPHP)/Y
      RETURN
C COEFFICIENTS IN GAUSSIAN UNITS FROM JONES AND MELOTTE (1953).
C THE FOLLOWING 14 CARDS CAN BE USED AS DATA CARDS FOR THIS SUBROUTINE
C 0.
C .3039 .0218
C .0176 -.0509 -.0135
C -.0255 .0515 -.0236 -.0074
C -.0393 -.0397 -.0238 .0087 -.0018
C .0293 -.0329 -.0130 .0031 .0034 .0005
C -.0211 -.0073 -.0007 .0210 .0017 -.0004 .0006
C 0.
C -.0555
C .0260 -.0044
C .0190 -.0033 -.0001
C -.0139 .0076 .0019 .0010
C .0057 -.0018 .0003 .0032 -.0004
C -.0026 -.0204 .0018 .0009 .0004 .0002
C THE FOLLOWING SET OF GAUSS NORMALIZED COEFFICIENTS WERE CONVERTED
C FROM THE SCHMIDT NORMALIZED COEFFICIENTS CALCULATED BY LINEARLY
C EXTRAPOLATING TO EPOCH 1974 THE COEFFICIENTS PUBLISHED FOR EPOCH
C 1960 BY CAIN AND SWEENEY (1970). (USES EARTH RADIUS = 6371.2)
C .000000
C+.300953 +.020298
C+.028106 -.05214 -.014435
C-.0308 +.06560 -.025252 -.006952
C-.041243 -.043956 -.016897 +.008021 -.002525
C+.014742 -.037078 -.018906 +.002819 +.003656 +.000036
C-.006713 -.012234 -.004364 +.02137 +.001593 -.000072 +.00068
C .000000
C .000000 -.057886
C .000000 +.035942 +.001129
C .000000 +.011084 -.004421 +.001180
C .000000 -.010299 +.008794 -.000086 +.002256
C .000000 -.003849 -.012615 +.007845 +.002207 -.000328
C .000000 +.003157 -.012670 -.009281 +.002286 -.000135 +.000243
      END

```

```

HARM087
HARM088
HARM089
HARM090
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HARM143
HARM144
HARM145-

```



## APPENDIX 6. COLLISION FREQUENCY MODELS WITH INPUT PARAMETER FORMS

The following collision frequency models are available. The input parameter forms, which describe the model, and the subroutine listings are given on the pages shown.

a.	Tabular profiles (TABLEZ)	152
b.	Constant collision frequency (CONSTZ)	155
c.	Exponential profile (EXPZ)	156
d.	Combination of two exponential profiles (EXPZ2)	157

To add other collision frequency models the user must write a subroutine that will calculate the normalized collision frequency ( $Z$ ) and its gradient ( $\partial Z/\partial r$ ,  $\partial Z/\partial \theta$ ,  $\partial Z/\partial \varphi$ ) as a function of position in spherical polar coordinates ( $r$ ,  $\theta$ ,  $\varphi$ ). ( $Z = \nu/2\pi f$ , where  $\nu$  is the collision frequency between electrons and neutral air molecules and  $f$  is the wave frequency. If the Sen-Wyller formula for refractive index is used, then  $Z = \nu_m/2\pi f$ , where  $\nu_m$  is the mean collision frequency.)

The restrictions on electron density models also apply to collision frequency models. The coordinates  $r$ ,  $\theta$ ,  $\varphi$  refer to the computational coordinates system, which may not be the same as geographic coordinates. In particular, they are geomagnetic coordinates when the earth-centered dipole model of the earth's magnetic field is used.

The input to the subroutine ( $r$ ,  $\theta$ ,  $\varphi$ ) is through blank common. (See Table 3.) The output is through common block /ZZ/. (See Table 10.) It is useful if the name of the subroutine suggests the model to which it corresponds. It should have an entry point COLFRZ so that other subroutines in the program can call it. Any parameter needed by the subroutine should be input into W251 through W299 of the W array. (See Table 2.) If the model needs massive amounts of data, these should be read in by the subroutine following the example of subroutine TABLEZ.





	SUBROUTINE TABLEZ	TABZ001
C	CALCULATES COLLISION FREQUENCY AND ITS GRADIENT FROM PROFILES	TABZ002
C	HAVING THE SAME FORM AS THOSE USED BY CROFTS RAY TRACING PROGRAM	TABZ003
C	MAKES AN EXPONENTIAL EXTRAPOLATION DOWN USING THE BOTTOM TWO POINTS	TABZ004
C	NEEDS SJBRROUTINE GAUSEL	TABZ005
	DIMENSION HPC(100),FN2C(100),ALPHA(100),BETA(100),	TABZ006
	1 GAMMA(100),DELTA(100),MAT(4,5),SLOPE(100)	TABZ007
	COMMON /CONST/ PI,PIT2,PID2,DJM(5)	TABZ008
	COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH	TABZ009
	COMMON R(6) /HW/ IO(10),W0,W(400)	TABZ010
	EQUIVALENCE (EARTH, W(2)), (F, W(6)), (READNU, W(250))	TABZ011
	REAL MAT	TABZ012
	DATA (MODZ=6HTABLEZ)	TABZ013
	ENTRY COLFRZ	TABZ014
	IF (.NOT.READNU) GO TO 10	TABZ015
	READNU=0.	TABZ016
	READ 2, NOC, (HPC(I), FN2C(I), I=1, NOC)	TABZ017
2	FORMAT(I4/(F8.2, E12.4))	TABZ018
	PRINT 1200, (HPC(I), FN2C(I), I=1, NOC)	TABZ019
1200	FORMAT(1H1, 14X, 6HHEIGHT, 4X, 20HCOLLISION FREQUENCY	TABZ020
	1(1X, F20.10, E20.10))	TABZ021
	A=0.	TABZ022
	IF (FN2C(1).NE.0.) A=ALOG(FN2C(2)/FN2C(1))/(HPC(2)-HPC(1))	TABZ023
	FN2C(1)=FN2C(1)/PIT2*1.E-6	TABZ024
	FN2C(2)=FN2C(2)/PIT2*1.E-6	TABZ025
	SLOPE(1)=A*FN2C(1)	TABZ026
	SLOPE(NOC)=0.	TABZ027
	DO 5 I=2, NOC	TABZ028
	IF (I.EQ.NOC) GO TO 6	TABZ029
	FN2C(I+1)= FN2C(I+1)/PIT2*1.E-6	TABZ030
	DO 3 J=1, 3	TABZ031
	M=I+J-2	TABZ032
	MAT(J,1)=1.	TABZ033
	MAT(J,2)=HPC(M)	TABZ034
	MAT(J,3)=HPC(M)**2	TABZ035
3	MAT(J,4)=FN2C(M)	TABZ036
	CALL GAUSEL (MAT,4,3,4,NRANK)	TABZ037
	IF (NRANK.LT.3) GO TO 20	TABZ038
	SLOPE(I)=MAT(2,4)+2.*MAT(3,4)*HPC(I)	TABZ039
5	CONTINUE	TABZ040
	DO 4 J=1, 2	TABZ041
	M=I+J-2	TABZ042
	MAT(J,1)=1.	TABZ043
	MAT(J,2)=HPC(M)	TABZ044
	MAT(J,3)=HPC(M)**2	TABZ045
	MAT(J,4)=HPC(M)**3	TABZ046
	MAT(J,5)=FN2C(M)	TABZ047
	L=J+2	TABZ048
	MAT(L,1)=0.	TABZ049
	MAT(L,2)=1.	TABZ050
	MAT(L,3)=2.*HPC(M)	TABZ051
	MAT(L,4)=3.*HPC(M)**2	TABZ052
4	MAT(L,5)=SLOPE(M)	TABZ053
	CALL GAUSEL (MAT,4,4,5,NRANK)	TABZ054
	IF (NRANK.LT.4) GO TO 20	TABZ055
	ALPHA(I)=MAT(1,5)	TABZ056
	BETA(I)=MAT(2,5)	TABZ057
	GAMMA(I)=MAT(3,5)	TABZ058
5	DELTA(I)=MAT(4,5)	TABZ059
	JUP=2	TABZ060
10	H=R(1)-EARTH	TABZ061
	IF (H.GE.HPC(1)) GO TO 12	TABZ062
11	JUP=2	TABZ063
	Z=FN2C(1)*EXP(A*(H-HPC(1)))/F	TABZ064
	PZPR=A*Z	TABZ065

RETURN	TABZ066
12 IF (H.GE.HPC(NOC)) GO TO 18	TABZ067
NSTEP=1	TABZ068
IF (H.LT.HPC(JUP-1)) NSTEP=-1	TABZ069
15 IF (HPC(JUP-1).GT.H.OR.H.GE.HPC(JUP)) GO TO 16	TABZ070
Z=(ALPHA(JUP)+H*(BETA(JUP)+H*(GAMMA(JUP)+H*DELTA(JUP))))/F	TABZ071
PZPR=(BETA(JUP)+H*(2.*GAMMA(JUP)+H*3.*DELTA(JUP)))/F	TABZ072
RETURN	TABZ073
16 JUP=JUP+NSTEP	TABZ074
IF (JUP.LT.2) GO TO 11	TABZ075
IF (JUP.LT.NOC) GO TO 15	TABZ076
18 JUP=NOC	TABZ077
Z=FN2C(NOC)/F	TABZ078
PZPR=0.	TABZ079
RETURN	TABZ080
20 PRINT 21, I,HPC(I)	TABZ081
21 FORMAT(4H THE,I4,58H TH POINT IN THE COLLISION FREQUENCY PROFILE HATA	TABZ082
15 THE HEIGHT,F8.2,40H KM, WHICH IS THE SAME AS ANOTHER POINT.)	TABZ083
CALL FXIT	TABZ084
END	TABZ085

## INPUT PARAMETER FORM FOR SUBROUTINE CONSTZ

An ionospheric collision frequency model consisting of a constant collision frequency

$$\nu = 0 \quad \text{for } h \leq h_{\min}$$

$$\nu = \nu_0 \quad \text{for } h > h_{\min}$$

Specify:

$\nu_0$  = \_\_\_\_\_ collisions per second (W251)

$h_{\min}$  = \_\_\_\_\_ km (W252)

C	<pre> SUBROUTINE CONSTZ CONSTANT COLLISION FREQUENCY COMMON /CONST/ PI,PIT2,PID2,DUM(5) COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH COMMON R(6) /WW/ ID(10),W0,W(400) EQUIVALENCE (EARTH,W(2)),(F,W(6)),(NU,W(251)),(HMIN,W(252)) REAL NU DATA (MODZ=6HCONSTZ) ENTRY COLFRZ H=R(1)-EARTH Z=0. IF (H.GT.HMIN) Z=NU/(PIT2*F)*1.E-6 RETURN END </pre>	<pre> CONZ001 CONZ002 CONZ003 CONZ004 CONZ005 CONZ006 CONZ007 CONZ008 CONZ009 CONZ010 CONZ011 CONZ012 CONZ013 CONZ014- </pre>
---	--	---

## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ

An ionospheric collision frequency model consisting of an exponential profile

$$\nu = \nu_0 e^{-a(h-h_0)}$$

$h$  is the height above the ground

Specify:

The collision frequency at the height  $h_0$ ,  $\nu_0 =$  \_\_\_\_\_  
collisions per second (W251)

The reference height,  $h_0 =$  \_\_\_\_\_ km (W252)

The exponential decrease of  $\nu$  with height,  $a =$  \_\_\_\_\_  $\text{km}^{-1}$   
(W253)

C	<pre> SUBROUTINE EXPZ   EXPONENTIAL COLLISION FREQUENCY MODEL   COMMON /CONST/ PI,PIT2,PID2,DUM(5)   COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH   COMMON R(6) /WW/ ID(10),W0,W(400)   REAL NU,NU0   EQUIVALENCE (EARTH,R(2)),(F,W(6)),(NU0,W(251)),(H0,W(252)), 1 (A,W(253))   DATA (MODZ=6H EXPZ )   ENTRY COLFRZ   H=R(1)-EARTH   NU=NU0/EXP (A*(H-H0))   Z=NU/(PIT2*F*1.E6)   PZPR  =-A*Z   RETURN   END                 </pre>	<pre> EXPZ001 EXPZ002 EXPZ003 EXPZ004 EXPZ005 EXPZ006 EXPZ007 EXPZ008 EXPZ009 EXPZ010 EXPZ011 EXPZ012 EXPZ013 EXPZ014 EXPZ015 EXPZ016-                 </pre>
---	---	---

## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ2

An ionospheric collision frequency model consisting of a combination of two exponential profiles

$$\nu = \nu_1 e^{-a_1(h-h_1)} + \nu_2 e^{-a_2(h-h_2)}$$

where  $h$  is the height above the ground.

Specify for the first exponential:

Collision frequency at height  $h_1$ ,  $\nu_1 = \frac{\text{_____}}{\text{per second (W251)}}$  collisions

Reference height,  $h_1 = \text{_____}$  km (W252)

Exponential decrease of  $\nu$  with height,  $a_1 = \text{_____}$  km<sup>-1</sup> (W253)

Specify for the second exponential:

Collision frequency at height  $h_2$ ,  $\nu_2 = \frac{\text{_____}}{\text{per second (W254)}}$  collisions

Reference height,  $h_2 = \text{_____}$  km (W255)

Exponential decrease of  $\nu$  with height,  $a_2 = \text{_____}$  km<sup>-1</sup> (W256)

C	<pre> SUBROUTINE EXPZ2   COLLISION FREQUENCY PROFILE FROM TWO EXPONENTIALS   COMMON /CONST/ PI,PIT2,PID2,DUM(5)   COMMON /ZZ/ MODZ,Z,PZPR,PZPTH,PZPPH   COMMON R(6) /WW/ ID(10),W0,W(400)   EQUIVALENCE (EARTH,W(2)),(F,W(6)),(NU1,W(251)),(H1,W(252)), 1 (A1,W(253)),(NU2,W(254)),(H2,W(255)),(A2,W(256))   RFAL NU1,NU2   DATA (MODZ=6H EXPZ?)   ENTRY COLFRZ   H=R(1)-EARTH   EXP1= NU1* EXP(-A1*(H-H1))   EXP2= NU2* EXP(-A2*(H-H2))   Z=(EXP1+EXP2)/(PIT2*F*1.E6)   PZPR=(-A1*EXP1-A2*EXP2)/(PIT2*F*1.F6)   RETURN   END                 </pre>	<pre> XPZ2001 XPZ2002 XPZ2003 XPZ2004 XPZ2005 XPZ2006 XPZ2007 XPZ2008 XPZ2009 XPZ2010 XPZ2011 XPZ2012 XPZ2013 XPZ2014 XPZ2015 XPZ2016 XPZ2017-                 </pre>
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## APPENDIX 7. CDC 250 PLOT PACKAGE

This appendix describes the plotting routines used by the Three-Dimensional Ray Tracing Program. The information was taken from "User's Guide to Cathode Ray Plotter Subroutines," ESSA Technical Memorandum ERLTM-ORSS 5, by L. David Lewis, January, 1970, and is printed with the permission of the author.

If you have access to a plotter, you may obtain plots by converting the following plotting commands to comparable commands on your system.

The CDC-250 Microfilm Recorder, under control of the NOAA Boulder CDC-3800 computer, plots data on the face of a high resolution cathode ray tube, which is photographed onto standard sized perforated, 35 mm film.

The plotting area, called a frame, is a square. Plotting positions are described in rectangular coordinates. Coordinate values are integers in the range 0 - 1023; (0, 0) is the "lower left hand corner".

Plotting specifications are transmitted to the plot routines via the following COMMON.

```
COMMON /DD/ IN, IOR, IT, IS, IC, ICC, IX, IY
```

The usage of each of the eight variables is listed below, followed by an explanation of the subroutine calls.

IN	Intensity. IN=0 specifies normal intensity. IN=1 specifies high intensity.
IOR	Orientation. IOR=0 specifies upright orientation. IOR=1 specifies rotated orientation (90° counter-clockwise).
IT	Italics (Font). IT=0 specifies non-Italic (Roman) symbols. IT=1 specifies Italic symbols.



IS      Symbol size.  
           IS=0 specifies miniature size.  
           IS=1 specifies small size.  
           IS=2 specifies medium size.  
           IS=3 specifies large size.

IC      Symbol case.  
           IC=0 specifies upper case.  
           IC=1 specifies lower case.

ICC     Character code, 0-63 (R1 format).  
           ICC and IC together specify the symbol plotted.

IX      X-coordinate, 0-1023.

IY      Y-coordinate, 0-1023.

CALL DDINIT (N, ID) is required to initialize the plotting process.

CALL DDBP      defines a vector origin at position IX, IY.

CALL DDVC      plots a vector (straight line), with intensity IN, from the vector origin defined by the previous DDBP or DDVC call, to the vector end position at IX, IY. A single call to DDBP followed by successive calls to DDVC (with changing IX and IY) plots connected vectors.

CALL DDTAB     initializes tabular plotting.

CALL DDTEXT (N, NT) plots a given array in a tabular mode, after initiating tabular plotting via DDTAB, as described above. NT is an array of length N, containing "text" for tabular plotting. Text consists of character codes, packed 8 per word (A8 Format). Text characters are plotted as tabular symbols until the command character  $\neq$  (octal code 14, card code 4, 8, or the alphabetic shift counterpart of the = on the keypunch) occurs. The command character is not plotted. DDTEXT interprets the next character as a command; and after the command is processed, tabular plotting resumes until  $\neq$  is again encountered.  $\neq$  . means end of text: DDTEXT returns to the calling routine.

CALL DDFR      causes a frame advance operation. Plotting on the current frame is completed, and the film advances to the next frame.

## APPENDIX 8. SAMPLE CASE

A sample case is included with the description of the program for three reasons. First, it demonstrates the use of the program. Second, it illustrates the three types of output available (printout, punched cards, and ray path plots). Finally, it serves as a test case to verify that the user's copy of the program is running correctly. This last point is especially important if the user has had to make many modifications in converting the program to run on a computer other than a CDC 3800.

Although the ionospheric models in the sample case demonstrate the use of the program, they don't give realistic absorption for the radio waves. The absorption in the sample case is too low for two reasons. First, although the Chapman layer has a realistic electron density for the F region, it has much too low an electron density for the D region, where most of the absorption occurs. Second, the collision frequency profile in the sample case is designed for use with the Sen-Wyller formula for refractive index rather than the Appleton-Hartree formula used in the sample case. Multiplying the collision frequency profile in the sample case by 2.5 gives an effective collision frequency profile for use with the Appleton-Hartree formula that will give nearly the correct absorption for HF radio waves (Davies, 1965, p. 89).

### Appendix 8a. Input Parameter Forms for the Sample Case

Filled-out input parameter forms are included to describe the sample case (i. e., show what ray paths are requested for which ionospheric models and what type of output is wanted). Furthermore, comparing them with Appendix 8b illustrates the relationship between the forms and the input data cards.



INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON A VERTICAL PLANE

Coordinates of the left edge of the graph:

Latitude = 40.  $\frac{\text{rad}}{\text{deg}}$  north (W83)  
km

Longitude = -105.  $\frac{\text{rad}}{\text{deg}}$  east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = 52.12  $\frac{\text{rad}}{\text{deg}}$  north (W85)  
km

Longitude = -81.8  $\frac{\text{rad}}{\text{deg}}$  east (W86)  
km

Height above the ground of the bottom of the graph = 0. km (W88)

Distance between tic marks = 100.  $\frac{\text{rad}}{\text{deg}}$  (W87)  
km

(W81 = 1.)

INPUT PARAMETER FORM FOR PLOTTING THE PROJECTION  
OF THE RAY PATH ON THE GROUND

Coordinates of the left edge of the graph:

Latitude = 40.  $\frac{\text{rad}}{\text{deg}}$  north (W83)  
km

Longitude = -105.  $\frac{\text{rad}}{\text{deg}}$  east (W84)  
km

Coordinates of the right edge of the graph:

Latitude = 52.12  $\frac{\text{rad}}{\text{deg}}$  north (W85)  
km

Longitude = -81.8  $\frac{\text{rad}}{\text{deg}}$  east (W86)  
km

Factor to expand lateral deviation scale by = 200. (W82)

Distance between tic marks on range scale = 100.  $\frac{\text{rad}}{\text{deg}}$  (W87)  
km

(W81 = 2.)

## INPUT PARAMETER FORM FOR SUBROUTINE CHAPX

An ionospheric electron density model consisting of a Chapman layer with tilts, ripples, and gradients

$$f_N^2 = f_c^2 \exp(\alpha(1-z-e^{-z}))$$

$$z = \frac{h - h_{\max}}{H}$$

$$f_c^2 = f_{c0}^2 \left( 1 + A \sin\left(2\pi\left(\theta - \frac{\pi}{2}\right)/B\right) + C\left(\theta - \frac{\pi}{2}\right) \right)$$

$$h_{\max} = h_{\max 0} + E\left(\theta - \frac{\pi}{2}\right) R_0$$

$f_N$  is the plasma frequency

$h$  is the height above the ground

$R_0$  is the radius of the earth in km

and  $\theta$  is the colatitude in radians.

Specify:

Critical frequency at the equator,  $f_{c0} = \underline{6.5}$  MHz (W101)

Height of the maximum electron density at the equator,  $h_{\max 0} = \underline{300}$  km (W102)

Scale height,  $H = \underline{62.}$  km (W103)

$\alpha = \underline{0.5}$  (W104, 0.5 for an  $\alpha$  Chapman layer, 1.0 for a  $\beta$  Chapman layer)

Amplitude of periodic variation of  $f_c^2$  with latitude,  $A = \underline{0.}$  (W105)

Period of variation of  $f_c^2$  with latitude,  $B = \underline{0.}$   $\frac{\text{rad}}{\text{deg}}$  (W106)  
km

Coefficient of linear variation of  $f_c^2$  with latitude,  $C = \underline{0.}$   $\text{rad}^{-1}$  (W107)

Tilt of the layer,  $E = \underline{0.}$   $\frac{\text{rad}}{\text{deg}}$  (W108)

## INPUT PARAMETER FORM FOR SUBROUTINE WAVE

A perturbation to an ionospheric electron density model consisting of a "gravity-wave" irregularity traveling from north pole to south pole

$$N = N_0(1 + \Delta)$$

$$\Delta = \delta \exp - [(R - R_0 - z_0)/H]^2 .$$

$$\cos 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z \right]$$

$$\frac{\partial N}{\partial t} = \frac{-2\pi}{\lambda_x} V_x N_0 \delta \exp - [(R - R_0 - z_0)/H]^2 .$$

$$\sin 2\pi \left[ t' + (\pi/2 - \theta) \frac{R_0}{\lambda_x} + (R - R_0)/\lambda_z \right]$$

$R_0$  is the radius of the earth.

$R$ ,  $\theta$ ,  $\varphi$  are the spherical (earth-centered) polar coordinates  
( $\Delta$  is independent of  $\varphi$ ).

$N_0(R, \theta, \varphi)$  is any electron density model.

Specify:

the height of maximum wave amplitude,  $z_0 = \underline{250.}$  km (W151)

wave-amplitude "scale height,"  $H = \underline{100.}$  km (W152)

wave perturbation amplitude,  $\delta = \underline{0.1}$  [0. to 1.] (W153)

horizontal trace velocity,  $V_x = \underline{-}$  km/sec (W154)  
(needed only if Doppler shift is calculated)

horizontal wavelength,  $\lambda_x = \underline{100.}$  km (W155)

vertical wavelength,  $\lambda_z = \underline{100.}$  km (W156)

time in wave periods,  $t' = \underline{0.}$  [0. to 1.] (W157)

## INPUT PARAMETER FORM FOR SUBROUTINE DIPOLY

An ionospheric model of the earth's magnetic field consisting of an earth centered dipole

The gyrofrequency is given by:

$$f_H = f_{H_0} \left( \frac{R_0+h}{R_0} \right)^3 \left( 1 + 3 \cos^2 \lambda \right)^{\frac{1}{2}}$$

The magnetic dip angle, I, is given by

$$\tan I = 2 \cot \lambda$$

h is the height above the ground

$R_0$  is the radius of the earth

$\lambda$  is the geomagnetic colatitude

Specify:

the gyrofrequency at the equator on the ground,  $f_{H_0} = \underline{0.8}$  MHz (W201)

the geographic coordinates of the north magnetic pole

latitude = 78.5 <sup>radians</sup> degrees north (W24)

longitude = 291. <sup>radians</sup> degrees east (W25)



## INPUT PARAMETER FORM FOR SUBROUTINE EXPZ 2

An ionospheric collision frequency model consisting of a double exponential profile

$$\nu = \nu_1 e^{-a_1(h-h_1)} + \nu_2 e^{-a_2(h-h_2)}$$

where  $h$  is the height above the ground.

Specify for the first exponential:

Collision frequency at height  $h_1$ ,  $\nu_1 = \frac{3.65 \times 10^4}{\text{per second (W251)}}$  collisions

Reference height,  $h_1 = \underline{100}$  km (W252)

Exponential decrease of  $\nu$  with height,  $a_1 = \underline{0.148}$  km<sup>-1</sup> (W253)

Specify for the second exponential:

Collision frequency at height  $h_2$ ,  $\nu_2 = \frac{30}{\text{per second (W254)}}$  collisions

Reference height,  $h_2 = \underline{140}$  km (W255)

Exponential decrease of  $\nu$  with height,  $a_2 = \underline{0.0183}$  km<sup>-1</sup> (W256)

## Appendix 8b. Listing of Input Cards for the Sample Case

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.....
X01 TEST CASE
1 0. OF DUPLICATE W CARDS, THE LAST ONE DOMINATES
1 -1. EXTRAORDINARY RAY
3 0. TRANSMITTER HEIGHT, KM
4 40. 1 TRANSMITTER LATITUDE, DEG NORTH
5 -105. 1 TRANSMITTER LONGITUDE, DEG EAST
7 6.0 INITIAL FREQUENCY, MC/S
9 0. DONT STEP FREQUENCY
11 45.0 1 INITIAL AZIMUTH ANGLE, DEGS CLOCKWISE FROM NORTH POLE
13 0. DONT STEP AZIMUTH ANGLE
15 0. 1 INITIAL ELEVATION ANGLE, DEG
16 90.0 1 FINAL ELEVATION ANGLE, DEG
17 15.0 1 STEP IN ELEVATION ANGLE, DEG
20 200. RECEIVER HEIGHT ABOVE THE EARTH, KM
22 3. NUMBER OF HOPS
57 2. INTEGRATE AND PRINT PHASE PATH
58 2. INTEGRATE AND PRINT ABSORPTION
71 5.0 NUMBER OF STEPS FOR EACH PRINTING
72 1. PUNCH RAYSETS
81 1. PLOT PROJECTION OF RAY PATH ON A VERTICAL PLANE
83 40.0 1 LEFT LATITUDE OF PLOT, DEG
84 -105. 1 LEFT LONGITUDE OF PLOT, DEG
85 52.12 1 RIGHT LATITUDE OF PLOT, DEG
86 -81.8 1 RIGHT LONGITUDE OF PLOT, DEG
87 100.0 1 DISTANCE BETWEEN TIC MARKS, KM
101 6.5 CRITICAL FREQUENCY, MC/S
102 300.0 HMAX, KM
103 62. SCALE HEIGHT, KM
104 0.5 ALPHA CHAPMAN LAYER
150 1. CALL PERTURBATION SUBROUTINE
151 250. Z0, KM
152 100. SH, SCALE HEIGHT, KM
153 0.1 DELTA
155 100. LAMBDA X, HORIZONTAL WAVELENGTH, KM
156 100. LAMBDA Z, VERTICAL WAVELENGTH, KM
201 0.8 GYROFREQUENCY ON THE GROUND AT THE EQUATOR, MHZ
24 78.5 1 ACCEPTED STANDARD LAT. OF NORTH MAGNETIC POLE, DEG NORTH
25 291. 1 ACCEPTED STANDARD LONG. OF NORTH MAGNETIC POLE, DEG EAST
251 3.65 E4 COLLISION FREQUENCY AT H1, /SEC
252 100.0 H1, REFERENCE HEIGHT, KM
253 .148 A1, EXPONENTIAL DECREASE OF NU WITH HEIGHT, /KM
254 30. COLLISION FREQUENCY AT H2, /SEC
255 140. H2, REFERENCE HEIGHT, KM
256 .0183 A2, EXPONENTIAL DECREASE OF NU WITH HEIGHT, /KM
(A BLANK IN COL. 1-3 ENDS THE CURRENT W ARRAY)

X01 TEST CASE
71 0. NO PERIODIC PRINTOUT
72 0. DO NOT PUNCH RAYSETS
81 2. PLOT PROJECTION OF RAY PATH ON THE GROUND
82 10.0 LATERAL DEVIATION EXPANSION FACTOR
(A BLANK IN COL. 1-3 ENDS THE CURRENT W ARRAY)
.....

```

Col. 1-3 Identification number

Col. 4-17 Data in E14.6 format

Col. 18 A 1 indicates an angle in degrees

Col. 19 A 1 indicates a central earth angle in kilometers

Col. 20 A 1 indicates a distance in nautical miles

Col. 21 A 1 indicates a distance in feet

Col. 22-24 Left for other conversions

Col. 25-80 Description of the data

X81 TEST CASE  
CHAPX WAVE D1POLY EXPZ2

11/05/74  
APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

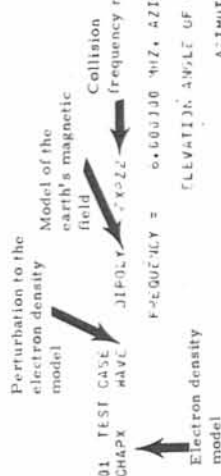
INITIAL VALUES FOR THE W ARRAY -- ALL ANGLES IN RADIANs, ONLY NONZERO VALUES PRINTED

1 -1.000000000000+000  
2 6.370000500000+003  
4 6.381317008003-001  
5 -1.83253571460+000  
7 6.000000000000+000  
11 7.85398153390-001  
16 1.57079632679+000  
17 2.61799387791-001  
20 2.000000000000+002  
22 3.000000000000+000  
23 1.000000000000+003  
24 1.37008346260+000  
25 5.07890812331+000  
41 3.000000000000+000  
42 1.000000000000-004  
43 5.000000000000+001  
44 1.000000000000+000  
45 1.000000000000+002  
46 1.000000000000-008  
47 5.000000000000-001  
57 2.000000000000+000  
58 2.000000000000+000  
71 5.000000000000+000  
72 1.000000000000+000  
81 1.000000000000+000  
83 6.981317008003-001  
84 -1.83253571460+000  
85 9.0965606124-001  
86 -1.42767932613+000  
87 1.55985871273-002  
101 6.500000000000+000  
102 3.000000000000+002  
103 6.139999999995+001  
104 5.000000000000-001  
150 1.000000000000+000  
151 2.500000000000+002  
152 1.000000000000+002  
153 1.000000000000-001  
155 1.000000000000+002  
156 1.000000000000+002  
201 8.000000000000-001  
251 3.650000000002+004  
252 1.000000000000+002  
253 1.460000000001-001  
254 3.000000000000+001  
255 1.399999999999+002  
256 1.830000000002-002

11/05/74  
 APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

F-COEFFICIENCY = 0.000330 W/MZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 3.000000 DEG



X01 TEST CASE CHAPX WAVE	HEIGHT KM	XTR IGN	RAJUE KY	AZIMUTH		ELEVATION		POLARIZATION		GROUP PATH KH	PHASE PATH KH	ABSORPTION DB
				XMR LOCAL DEG	LOCAL DEG	XMR LOCAL DEG	LOCAL DEG	REAL	IMAG			
0+000	0.0000	0.0000	0.0000	0.000	0.000	0.000	0.000	3.000	-1.000	0.0000	0.0000	0.0000
-3-011	52.9767	-18.7210	-16.6860	-0.000	-0.000	-0.000	7.364	3.597	-1.427	823.2593	823.2593	0.0000
0+000	59.4425	-16.6860	-16.6860	-0.000	-0.000	-0.000	7.797	3.513	-2.672	872.2593	872.2593	0.0000
-3-011	54.9916	56.0962	56.0962	-0.000	-0.000	-0.000	6.150	3.293	-2.386	912.2593	912.2593	0.0000
-3-011	7.7639	345.2593	345.2593	0.000	0.000	-0.000	6.502	0.141	-2.599	952.2593	952.2593	0.0000
-2-009	65.6755	1128.3470	1128.3470	0.000	-0.000	-0.000	9.345	0.021	-3.200	1048.2593	1048.2593	0.0004
-2-008	39.1450	1115.6700	1115.6700	0.000	-0.000	-0.000	10.039	0.000	-4.060	1128.2593	1128.2593	0.0014
-1-005	113.9441	1134.0737	1134.0737	0.000	-0.002	-0.002	10.683	0.001	-7.942	1208.2593	1208.2593	0.0034
-1-005	128.6479	1271.1132	1271.1132	0.000	-0.012	-0.012	10.999	0.001	-7.942	1288.2593	1288.2593	0.0058
-4-006	143.4758	1347.7703	1347.7703	0.001	0.012	-0.029	9.958	0.003	-5.844	1367.5112	1367.5112	0.0076
-7-016	153.6536	1416.6244	1416.6244	0.001	0.021	-0.073	6.649	0.003	-2.673	1440.2593	1440.2593	0.0089
-9-016	157.2033	1456.6624	1456.6624	0.001	0.029	-0.253	3.399	0.003	-2.083	1480.2593	1480.2593	0.0097
-9-016	158.1469	1491.1561	1491.1561	0.002	-0.039	-0.753	0.000	0.003	-1.720	1518.9022	1512.7768	0.0106
-3-011	156.1469	1491.1561	1491.1561	0.002	-0.039	-0.753	0.000	0.003	-1.720	1518.9022	1512.7768	0.0106
-6-011	150.1177	1494.9519	1494.9519	0.006	0.056	-1.756	-7.350	0.000	-1.365	1619.3022	1619.3022	0.0126
-2-005	133.2039	1659.7028	1659.7028	0.007	0.054	-2.782	-10.588	0.000	-1.318	1695.3022	1695.3022	0.0144
0+000	106.1392	1613.6452	1613.6452	0.006	0.016	-4.796	-10.863	0.000	-1.341	1855.9022	1855.9022	0.0191
0+000	90.8887	1559.0039	1559.0039	0.004	0.013	-6.536	-9.113	0.003	-1.403	2015.9022	2003.6590	0.0218
8-005	4.92235	2146.5251	2146.5251	0.006	0.010	-13.042	0.738	0.000	-1.000	2239.9022	2227.6590	0.0220
0+000	0.0000	2400.0492	2400.0492	-0.006	0.008	-15.566	7.401	-0.933	1.636	2955.4329	2943.2497	0.0220
0+000	2.23830	3540.6051	3540.6051	-0.003	0.008	-16.214	9.661	-0.933	1.705	3700.7307	3688.5475	0.0220
0+000	53.4935	3944.7384	3944.7384	-0.003	0.008	-16.214	9.661	-0.933	1.705	3700.7307	3688.5475	0.0220
-9-008	31.1154	3422.1351	3422.1351	-0.003	0.008	-16.435	10.475	-0.000	1.984	3957.7307	3945.5474	0.0227
-1-006	107.4072	3785.1994	3785.1994	-0.002	0.008	-16.435	10.475	-0.000	1.635	4053.7307	4041.5426	0.0247
-1-006	107.4072	3785.1994	3785.1994	-0.003	0.008	-16.435	10.475	-0.000	1.635	4053.7307	4041.5426	0.0247
-4-006	119.7975	4146.9652	4146.9652	-0.003	0.008	-16.578	10.307	-0.000	1.765	4117.7307	4105.5117	0.0266
-5-005	149.3069	4185.0204	4185.0204	-0.004	0.031	-16.914	9.243	-0.000	1.868	4261.7307	4248.4412	0.0303
-5-005	156.1106	4231.1682	4231.1682	-0.006	0.047	-17.169	4.188	-3.000	2.901	4341.7307	4325.8067	0.0320
0+000	157.9016	4305.0862	4305.0862	-3.107	0.076	-17.356	-0.000	-0.000	9.317	4388.0431	4369.8069	0.0330

THIS TABLE CALCULATION TOOK 12.382 SEC

Polarization = +) means the electric field vector is rotating counter clockwise when looking along the ray.  
 Angle of the wave normal with the local horizontal.  
 Elevation angle of current ray path point at the transmitter.  
 Azimuth angle of the wave normal in degrees clockwise from great circle between transmitter and ray point.  
 Azimuth angle of the direction of transmission in degrees clockwise from great circle between transmitter and ray point.  
 Great circle distance along the ground between the ray point and the transmitter.  
 Height of ray point above the ground.  
 v / Real part (n<sup>2</sup>)-1 where v is the magnitude of the wave normal index, and n is the complex phase refractive index.  
 This quantity would be zero if there were no errors in the numerical integration.

11/05/74

X01 TEST CASE CHAPX WAVE DIPOLY EXPZ2 APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG  
ELEVATION ANGLE OF TRANSMISSION = 15.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000	-0.000	-0.000	15.000	15.000	0.001	-1.000	0.0000	C.0000	0.0000
0+000 ENTR ION	52.9852	185.0044	-0.000	-0.000	15.000	16.073	3.113	-3.358	194.1401	194.1401	0.0000
0+000	63.7818	211.6249	-0.000	-0.000	15.000	16.903	2.323	-7.474	221.1401	221.1401	0.0000
0+000	66.6275	230.5627	-0.000	-0.000	15.000	17.074	1.297	-9.793	241.1401	241.1401	0.0000
-5-010	79.6832	272.1042	-0.000	-0.000	15.000	17.447	0.364	-18.918	285.1401	285.1401	0.0001
-7-009	91.7890	399.7220	-0.000	-0.000	15.000	17.765	0.353	-110.019	325.1401	325.1401	0.0004
-2-008	104.1175	347.1979	-0.000	-0.000	15.000	18.116	-0.015	29.112	365.1401	365.1401	0.0011
1-006	126.7952	414.2867	-0.000	-0.000	14.990	18.525	-0.001	10.310	437.0846	437.0846	0.0031
3-005	151.4725	433.2463	-0.000	-0.000	14.941	16.907	-0.001	459.192	517.1401	517.1401	0.0048
-9-006	161.8111	524.9617	-0.001	-0.002	14.599	13.612	0.000	-4.495	557.1401	557.1401	0.0057
-1-005	169.1102	561.6061	0.000	0.000	14.015	8.649	0.000	-2.160	597.1401	597.1401	0.0067
-7-005	172.1418	598.0253	0.015	0.155	13.155	1.378	0.000	-1.453	637.1401	637.1401	0.0079
-3-011 MIN DIST	172.1342	604.1034	0.015	0.184	12.975	0.000	0.000	-1.481	643.8528	632.4672	0.0081
-3-011 MIN DIST	172.1342	604.1034	0.015	0.184	12.975	0.000	0.000	-1.481	643.8528	632.4672	0.0081
2-010 WAVE REV	172.0659	607.7222	0.015	0.196	12.844	-0.320	0.000	-1.452	647.8528	635.9500	0.0083
-3-005	193.8276	726.9590	0.024	-0.078	8.942	-15.755	0.000	-1.173	756.8528	735.2639	0.0114
-3-005	105.1516	855.5089	0.037	-0.032	5.835	-18.205	0.000	-1.154	836.8528	813.9074	0.0133
-3-005	81.0547	730.6930	0.039	-0.030	3.033	-17.862	0.000	-1.162	916.8528	893.8080	0.0160
-3-005	48.4334	1036.8769	0.042	-0.027	1.752	-17.194	0.000	-1.174	996.8528	973.8062	0.0173
0+000 GRND REF	0.0000	1212.9251	0.046	-0.023	-1.983	-16.239	0.075	-1.000	1108.8528	1085.8062	0.0174
-3-011 ENTR ION	52.9855	1403.8408	0.045	-0.020	-5.455	-14.656	-0.002	1.885	1292.1934	1269.1468	0.0174
-3-011	54.1136	1406.8464	0.045	-0.020	-4.165	-16.366	-1.306	1.885	1490.3081	1467.2614	0.0174
-9-007	113.0072	1585.1016	0.052	-0.018	-3.213	-17.984	-1.305	2.105	1494.3081	1471.2614	0.0174
-1-006	134.3372	1559.5945	0.052	-0.018	-2.890	-18.138	-0.000	2.342	1683.3081	1660.2574	0.0191
-5-006	156.1139	1726.0642	0.051	-0.021	-2.688	-15.432	-0.000	3.024	1895.3081	1870.0923	0.0214
-1-005	165.1214	1762.7028	0.051	-0.021	-2.675	-11.173	-0.000	14.789	1875.3081	1846.1145	0.0237
-7-006	170.7131	1799.2958	0.054	-0.020	-2.779	5.562	0.000	-3.471	1915.3081	1884.8621	0.0248
0+000 MIN DIST	171.9566	1828.6204	0.059	-0.065	-2.959	-0.000	0.000	-1.952	1947.4107	1912.8923	0.0258

THIS RAY CALCULATION TOOK 11.2+7 SEC

X01 TEST CASE  
 CHAPX WAVE DIPOLY EXP22 APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS 11/05/74  
 FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG  
 ELEVATION ANGLE OF TRANSMISSION = 30.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
3-011 XMTR	0.0000	0.0000									
0+030 ENTR ION	52.4735	89.9010	-0.300	-0.000	30.000	30.000	-0.000	0.000	0.0000	0.0000	0.0000
-3-011	60.5634	102.6627	-0.300	-0.000	30.000	30.423	-0.213	1.259	104.6686	104.6686	0.0000
0+000	65.8141	111.1535	-0.000	-0.000	30.000	31.000	-0.110	1.614	119.6686	119.6686	0.0000
-2-010	78.2077	131.4761	-0.000	-0.000	30.000	31.163	-0.017	1.609	129.6686	129.6686	0.0000
-2-009	88.5857	148.3519	-0.300	-0.000	30.000	31.334	-0.004	1.592	153.6686	153.6686	0.0000
-1-007	107.3766	178.5314	-0.000	-0.000	30.000	31.600	-0.000	1.562	173.6686	173.6686	0.0002
-1-006	128.3934	211.9316	-0.800	-0.001	29.993	31.773	-0.000	1.539	208.6686	208.6686	0.0008
5-006	149.2460	245.1205	-0.300	-0.009	29.935	31.883	-0.000	1.574	249.6686	249.6686	0.0020
-2-005	168.5912	277.9003	-0.311	-0.068	29.657	27.599	-0.000	1.695	289.6686	289.6686	0.0029
2-005	144.0395	310.3408	-0.016	0.029	28.909	19.956	-0.000	7.331	329.6686	329.6686	0.0038
1-006	131.3569	342.2014	-0.027	0.444	27.306	5.952	0.000	-1.911	369.6686	369.6686	0.0052
0+000 MIN DIST	131.5641	354.9408	-1.023	0.219	26.398	0.000	0.000	-1.530	409.6686	390.0649	0.0073
0+030 MIN DIST	131.5641	354.9408	-1.023	0.219	26.398	0.000	0.000	-1.530	425.7924	400.9614	0.0082
6-010 WAVE REV	131.0066	358.1065	-0.920	0.137	26.154	-1.395	0.000	-1.675	429.7924	403.6919	0.0085
4-006	172.8604	426.8860	0.135	-0.735	19.783	-23.513	0.000	-1.113	518.7924	468.9259	0.0125
2-005	135.6960	493.7148	0.238	-0.521	12.986	-29.961	0.000	-1.076	598.7924	544.6800	0.0147
2-005	95.7861	561.7275	0.298	-0.455	7.074	-29.718	0.000	-1.080	678.7924	624.5459	0.0172
2-005	56.5015	630.5940	0.348	-0.405	2.257	-29.100	0.039	-1.052	754.7924	704.5457	0.0177
2-005	46.7356	644.4661	0.357	-0.397	1.406	-28.975	0.035	-1.018	770.7924	720.5457	0.0177
-3-011 GRND REF	3.0000	733.6080	0.405	-0.308	-3.299	28.173	0.000	1.000	872.6850	822.4383	0.0177
0+000 ENTR ION	52.3781	830.3382	0.446	-0.308	-0.104	29.043	-0.277	1.223	983.3256	933.0769	0.0177
-4-007	92.6830	933.1376	0.470	-0.284	1.779	29.671	-0.002	1.697	1068.3256	1018.0788	0.0180
-2-005	132.5904	968.2234	0.490	-0.266	3.349	30.046	-0.000	1.461	1148.3256	1093.9864	0.0201
-1-005	163.3870	1021.9170	0.501	-0.253	4.356	26.978	-0.000	1.628	1208.3256	1156.1125	0.0215
-2-005	179.0552	1055.0139	0.507	-0.452	4.735	20.718	-0.000	2.556	1248.3256	1191.4144	0.0228
-9-006	186.0742	1087.7704	0.511	0.102	4.744	9.433	0.000	-4.108	1288.3256	1222.7505	0.0245
0+000 MIN DIST	109.8217	1107.5272	0.511	0.454	4.541	0.000	0.000	-1.771	1312.8736	1240.1070	0.0259

THIS RAY CALCULATION TOOK 8.642 SEC

X01 TEST CASE  
 CHAPX WAVE

DIPOLY EXPZ2

11/05/74  
 APFLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 45.000000 DEG

174

		HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
				XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000	XMTR	0.0000	0.0000									
0+000	ENTR ION	52.9759	52.3246	-0.000	-0.000	45.000	45.471	-0.113	1.100	0.0000	0.0000	0.0000
3-011		62.2439	61.3530	-0.500	-0.000	45.000	45.552	-0.059	1.207	74.6128	74.6128	0.0000
0+000		69.3926	68.2802	0.000	0.000	45.000	45.614	-0.021	1.219	87.6128	87.6128	0.0000
-3-009		65.1331	63.4659	-0.000	-0.000	45.000	45.751	-0.002	1.217	97.6128	97.6128	0.0001
-4-008		39.4743	97.2070	-0.000	-0.000	45.000	45.873	-0.000	1.213	119.6128	119.6128	0.0004
-1-007		111.8433	110.8668	-0.000	-0.000	45.000	45.886	-0.000	1.209	139.6128	139.6125	0.0009
5-005		142.5506	138.0458	-0.002	0.003	44.980	45.857	-0.000	1.209	159.6128	159.6091	0.0021
-2-005		173.1037	164.7826	-0.011	-0.106	44.789	43.268	-0.000	1.260	199.6128	199.4233	0.0030
6-005		193.2703	190.6257	-0.071	0.369	44.104	34.673	-0.000	1.601	239.6128	237.3102	0.0048
-3-011	RCVR	200.0000	200.2014	-0.132	0.607	43.622	28.480	-0.000	2.293	279.6128	278.4398	0.0058
2-006		209.6843	233.2599	-0.069	-1.590	40.436	4.780	0.000	-1.576	295.1649	295.1649	0.0095
2-006	APOGEE	209.6634	235.8629	-0.040	-1.746	40.109	3.110	0.000	-1.831	348.1649	348.1649	0.0098
3-005	WAVE REV	209.4300	241.0969	0.021	-1.982	39.431	-0.266	0.000	-1.623	352.1649	352.1649	0.0103
3-007		205.9618	259.5199	0.244	-2.001	36.819	-12.892	0.000	-1.268	360.1649	360.1649	0.0122
3-011	RCVR	260.0000	274.3768	0.385	-1.415	34.427	-23.456	0.000	-1.142	388.1649	388.1649	0.0137
1-005		172.6329	311.6736	0.547	-0.183	27.233	-41.318	0.000	-1.041	410.8543	410.8543	0.0169
4-006		117.8520	366.4800	0.563	-0.033	16.021	-45.154	0.000	-1.031	467.8543	467.8543	0.0193
6-006		89.5472	394.2849	0.570	-0.030	10.932	-44.927	0.000	-1.632	547.8543	547.8543	0.0204
6-006		38.3028	444.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	587.8543	587.8543	0.0206
6-006	EXIT ION	38.3028	444.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	659.8543	659.8543	0.0206
0+000	GRND REF	0.0000	484.7060	0.575	-0.025	-2.180	44.114	-0.000	1.000	566.5130	566.5130	0.0206
0+000	ENTR ION	52.9777	538.6621	0.578	-0.022	3.168	44.599	-0.105	1.694	715.5626	715.5626	0.0206
-6-007		110.1068	595.3574	0.581	-0.020	7.705	45.103	-0.000	1.136	791.3392	791.3392	0.0214
-6-007		138.4237	622.9759	0.581	-0.016	9.586	45.095	-0.000	1.194	872.3392	872.3392	0.0226
-4-006		165.8750	650.1889	0.579	-0.068	11.199	42.989	-0.000	1.230	912.3392	912.3392	0.0236
7-005		183.5935	676.6832	0.572	0.057	12.376	35.811	-0.000	1.434	952.3392	952.3392	0.0252
-3-011	RCVR	200.0000	691.5163	0.546	0.796	12.769	26.353	-0.000	2.159	992.3392	992.3392	0.0266

THIS RAY CALCULATION TOOK 10.310 SEC

XJ1 TEST CASE  
 CHAPX WAVE DIPOLE EXPZ2 11/05/74  
 APPLETON-HAKTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 8.600000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 60.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000	-0.300	-0.300	60.000	60.000	-0.000	1.000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9788	30.2936	-0.000	-0.000	60.272	61.0908	-0.053	1.049	61.0908	61.0908	0.0000
0+000	59.9272	34.2235	-0.000	-0.000	60.308	69.0908	-0.035	1.092	69.0908	69.0908	0.0000
0+000	67.7470	38.6347	-0.000	-0.000	60.000	60.348	-0.012	1.103	78.0908	78.0908	0.0000
-7-009	85.1334	46.3990	-0.000	-0.000	60.000	60.435	-0.001	1.103	98.0908	98.0908	0.0001
-1-007	102.5337	58.1107	-0.000	-0.000	60.000	60.522	-0.003	1.101	118.0908	118.0908	0.0004
-4-007	119.9585	67.7596	-0.000	-0.000	59.939	60.592	-0.000	1.100	138.0908	138.0820	0.0011
-4-006	154.5342	86.6836	-0.000	-0.000	59.967	60.149	-0.000	1.103	178.0908	177.5563	0.0022
4-005	196.1632	105.2175	-0.094	0.088	59.597	55.830	-0.000	1.158	218.0908	211.9329	0.0037
0+000 RCVR	200.0000	114.4162	-0.157	-0.951	39.325	50.578	-0.003	1.250	239.2552	225.6343	0.0050
9-005	224.6062	137.6335	-0.234	0.797	37.441	24.698	-0.000	5.656	296.2552	248.7596	0.0090
-1-004	225.8382	146.6930	-0.354	5.402	35.872	-5.298	0.000	-2.002	324.2552	253.2681	0.0114
-1-004	221.4135	153.9714	-0.522	9.624	34.028	-33.727	0.000	-1.186	348.2552	277.0925	0.0135
-1-004	204.2335	166.0800	-1.210	16.099	49.687	-60.227	0.000	-1.030	388.2552	259.8039	0.0168
0+000 RCVR	200.0000	166.4639	-1.383	16.376	48.744	-62.499	0.000	-1.025	395.1225	273.3158	0.0173
-1-005	153.9530	185.8715	-3.075	16.994	38.450	-69.419	0.000	-1.013	452.1225	318.1391	0.0201
-4-005	116.7265	198.8446	-4.206	15.804	29.230	-69.701	0.000	-1.013	492.1225	357.6180	0.0213
-2-005	79.2276	212.0602	-5.211	14.796	19.415	-69.586	0.000	-1.013	532.1225	397.6124	0.0223
-2-005	20.7751	236.9198	-6.644	13.564	5.561	-69.411	0.000	-1.000	588.1225	453.6124	0.0223
0+000 GRND REF	0.0000	240.7118	-7.005	13.003	-1.003	60.320	-0.000	1.000	616.2330	482.2229	0.0223
-3-011	52.9846	260.0662	-7.392	12.021	10.239	69.499	-0.026	1.025	673.3376	538.8275	0.0223
-3-011	56.7354	261.4245	-8.356	11.957	11.019	69.511	-0.003	1.038	677.3376	542.8275	0.0223
-3-005	113.9101	281.9769	-9.739	11.075	20.549	69.695	-0.000	1.050	738.3376	603.8243	0.0232
6-005	151.2131	295.2551	-3.446	10.631	25.514	69.477	-0.000	1.051	778.3376	643.4278	0.0244
-4-005	185.7553	308.0424	-9.310	9.231	29.335	66.600	-0.000	1.071	818.3376	678.2033	0.0259
3+000 RCVR	200.0000	313.9111	-10.115	8.555	30.682	63.564	-0.000	1.099	837.7013	691.8808	0.0271

THIS RAY CALCULATION TOOK 8.004 SEC



X01 TEST CASE  
CHAPX WAVE

DIFOLY EXPZ2

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

11/05/74

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 75.000000 DEG

		HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
				XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000	XMTR	0.0000	0.0000									
0+000	ENTR ION	52.9846	14.0759	-0.000	-0.000	75.000	75.127	-0.024	1.023	0.0000	0.0000	0.0000
-3-011		69.7166	16.1160	-0.000	-0.000	75.000	75.145	-0.014	1.044	54.8374	54.8374	0.0000
-3-011		65.5499	17.3768	-0.000	-0.000	75.000	75.156	-0.003	1.047	67.8374	67.8374	0.0000
-2-006		87.7850	23.1903	-0.000	-0.000	75.000	75.209	-0.000	1.048	90.8374	90.8374	0.0001
-1-007		107.1237	28.2124	-0.000	-0.000	75.000	75.253	-0.000	1.047	110.8374	110.8365	0.0005
-9-007		126.4556	33.2025	-0.001	0.000	74.999	75.279	-0.000	1.046	130.8374	130.8161	0.0013
4-007		145.6921	38.1498	-0.007	-0.012	74.994	75.188	-0.000	1.047	150.8374	150.6143	0.0019
3-005		174.7737	46.8081	-0.062	-0.321	74.920	73.849	-0.000	1.057	186.8374	183.5011	0.0030
0+000	RCVR	206.0000	52.6875	-0.362	1.334	74.783	71.357	-0.000	1.087	213.9181	202.1284	0.0046
6-005		224.9440	60.6695	-1.077	1.370	74.379	52.074	-0.000	1.568	266.9181	217.2922	0.0091
1-004		230.7996	66.1625	-1.711	-7.396	73.432	15.569	0.000	-9.825	306.9181	220.0213	0.0127
1-004	APOGEE	230.9183	67.5564	-1.511	-8.853	73.111	8.237	0.000	-3.876	314.9181	220.4673	0.0134
9-005	WAVE REV	230.2593	70.7641	-0.922	-11.181	72.307	-4.474	0.000	-1.950	330.9181	221.5782	0.0147
5-005		222.0226	84.2620	1.343	-14.206	68.508	-29.337	0.000	-1.143	378.9181	228.5798	0.0187
0+000	RCVR	200.0000	106.6297	5.511	-13.543	61.084	-46.475	0.000	-1.024	434.9673	249.2747	0.0232
2-005		161.0693	134.6131	8.542	-11.278	49.256	-55.293	0.000	-1.007	491.9673	291.7804	0.0263
-2-005		129.0155	155.6964	10.092	-9.754	38.661	-56.407	0.000	-1.006	531.9673	330.5977	0.0275
-3-006		95.7312	177.1805	11.285	-8.562	27.405	-56.282	0.000	-1.006	571.9673	370.5601	0.0289
-1-006		62.4974	196.9508	12.221	-7.620	16.463	-56.084	0.001	-1.006	611.9673	410.5600	0.0292
-1-006	EXIT ION	29.3420	220.9903	12.392	-6.356	6.551	-55.885	0.000	-1.000	651.9673	450.5600	0.0292
0+000	GRNG REF	0.0000	240.7589	13.557	-6.291	-1.083	55.706	-0.000	1.000	687.4459	486.0386	0.0292
0+000	ENTR ION	52.9836	276.3463	14.370	-5.479	9.565	56.027	-0.069	1.064	751.4563	550.0489	0.0292
-8-003		67.0257	298.8465	14.784	-5.065	14.785	56.231	-0.001	1.134	792.4563	591.0489	0.0293
-6-006		120.3097	320.5864	15.127	-4.721	18.948	56.406	-0.000	1.130	832.4563	631.0396	0.0303
2-005		153.3776	342.0624	15.421	-4.425	22.353	55.904	-0.000	1.135	872.4563	670.5260	0.0314
4-005		183.3935	362.9372	15.675	-4.214	24.915	51.894	-0.000	1.195	912.4563	705.8795	0.0327
-3-011	RCVR	200.0000	375.3567	15.794	-3.644	25.997	45.376	-0.000	1.374	937.9577	722.8911	0.0343

THIS RAY CALCULATION TOOK 9.624 SEC

11/05/74

401 TEST CASE DRAPX WAVE DIPOLY \*K\*ZZ APPLETON-HAFTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.400000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 90.000000 DEG

MODE	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000	KMTR	0.0000									
0+000	ENTR ION	52.9908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0000	52.9908	0.0000	0.0000
0+000	MAX LAT	54.9908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0112	54.9908	0.0000	0.0000
0+000		56.9908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0114	56.9908	0.0000	0.0000
0+000		62.4908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0114	62.4908	0.0000	0.0000
0+000		64.9908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0114	64.9908	0.0000	0.0000
0+000		67.4908	0.0000	90.0000	90.0000	90.0000	-0.0000	1.0114	67.4908	0.0000	0.0000
0+000	WAVE REV	69.2408	0.0000	90.0000	90.0000	90.0000	-0.0002	1.0119	69.2408	0.0000	0.0000
0+000		71.7408	0.0000	90.0000	90.0000	90.0000	-0.0002	1.0113	71.7408	0.0000	0.0000
0+000		74.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0119	74.2408	0.0000	0.0000
0+000		76.7408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0119	76.7408	0.0000	0.0000
0+000		79.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0119	79.2408	0.0000	0.0000
0+000		81.7408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0119	81.7408	0.0000	0.0000
0+000		84.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0119	84.2408	0.0000	0.0000
0+000		86.7408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	86.7408	0.0000	0.0000
0+000		89.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	89.2408	0.0000	0.0000
0+000		91.7408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	91.7408	0.0000	0.0000
0+000		94.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	94.2408	0.0000	0.0000
0+000		96.7408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	96.7408	0.0000	0.0000
0+000		99.2408	0.0000	90.0000	90.0000	90.0000	-0.0001	1.0118	99.2408	0.0000	0.0000
0+000	WAVE REV	105.2339	0.0001	145.186	-179.920	30.0000	0.0000	1.0118	105.2339	0.0000	0.0000
0+000		110.2339	0.0001	145.276	-179.991	30.0000	0.0000	1.0118	110.2339	0.0000	0.0000
0+000		115.2339	0.0003	145.262	0.004	30.0000	0.0000	1.0118	115.2339	0.0000	0.0000
0+000		119.7339	0.0007	145.269	0.006	30.0000	0.0000	1.0118	119.7339	0.0000	0.0000
0+000		125.7339	0.0039	145.266	0.001	30.0000	0.0000	1.0118	125.7339	0.0000	0.0000
0+000		136.6856	0.0121	145.266	0.000	30.0000	0.0000	1.0118	136.6856	0.0000	0.0000
0+000		146.5779	0.0314	145.266	0.000	30.0000	0.0000	1.0118	146.5779	0.0000	0.0000
0+000		154.1272	0.0700	145.266	0.000	30.0000	0.0000	1.0118	154.1272	0.0000	0.0000
0+000		167.4214	0.1197	145.266	0.000	30.0000	0.0000	1.0118	167.4214	0.0000	0.0000
0+000	WAVE REV	175.1512	0.1283	145.266	-180.000	30.0000	0.0000	1.0118	175.1512	0.0000	0.0000
0+000		176.9406	0.2533	145.266	-180.000	30.0000	0.0000	1.0118	176.9406	0.0000	0.0000
0+000		182.0775	0.3058	145.266	-180.000	30.0000	0.0000	1.0118	182.0775	0.0000	0.0000
0+000	RCVR	200.0000	0.3932	145.266	-180.000	30.0000	0.0000	1.0118	200.0000	0.0000	0.0000
0+000		210.5537	0.7709	145.266	-180.000	30.0000	0.0000	1.0118	210.5537	0.0000	0.0000
0+000		227.9866	0.9818	145.266	0.000	30.0000	0.0000	1.0118	227.9866	0.0000	0.0000
0+000	WAVE REV	230.2036	2.6623	145.266	0.000	30.0000	0.0000	1.0118	230.2036	0.0000	0.0000
0+000		237.3654	3.6109	145.266	0.000	30.0000	0.0000	1.0118	237.3654	0.0000	0.0000
0+000	WAVE REV	236.2305	4.6933	145.266	0.000	30.0000	0.0000	1.0118	236.2305	0.0000	0.0000
0+000	APOGEE	233.0027	5.1809	145.266	0.000	30.0000	0.0000	1.0118	233.0027	0.0000	0.0000
0+000		236.0571	5.1809	145.266	0.000	30.0000	0.0000	1.0118	236.0571	0.0000	0.0000
0+000		221.7660	9.4861	145.266	0.000	30.0000	0.0000	1.0118	221.7660	0.0000	0.0000
0+000	RCVR	200.0000	15.3973	145.266	0.000	30.0000	0.0000	1.0118	200.0000	0.0000	0.0000
0+000		131.1549	24.7823	145.266	0.000	30.0000	0.0000	1.0118	131.1549	0.0000	0.0000
0+000		112.0674	32.1355	145.266	0.000	30.0000	0.0000	1.0118	112.0674	0.0000	0.0000
0+000	EXIT ION	72.7938	39.6085	145.266	0.000	30.0000	0.0000	1.0118	72.7938	0.0000	0.0000
0+000		25.6748	48.6937	145.266	0.000	30.0000	0.0000	1.0118	25.6748	0.0000	0.0000
0+000	GRND REF	0.0000	53.7019	145.266	0.000	30.0000	0.0000	1.0118	0.0000	0.0000	0.0000
0+000		52.1956	63.9364	145.266	0.000	30.0000	0.0000	1.0118	52.1956	0.0000	0.0000
0+000	ENTR ION	45.1535	70.2016	145.266	0.000	30.0000	0.0000	1.0118	45.1535	0.0000	0.0000
0+000		12.16623	77.6423	145.266	0.000	30.0000	0.0000	1.0118	12.16623	0.0000	0.0000
0+000		103.26623	86.9935	145.266	0.000	30.0000	0.0000	1.0118	103.26623	0.0000	0.0000
0+000		137.0502	92.0447	145.266	0.000	30.0000	0.0000	1.0118	137.0502	0.0000	0.0000
0+000	RCVR	200.0000	92.7166	145.266	0.000	30.0000	0.0000	1.0118	200.0000	0.0000	0.0000

THIS TABLE CALCULATION TOOK 13.937 SEC

1 -1.0J10J0G0000+00  
 2 5.3700000000+003  
 4 0.38131706603-001  
 5 -1.83259571460+000  
 7 5.0000000000+000  
 11 7.8239616330-001  
 16 1.57073632679+000  
 17 2.6179387791-001  
 20 2.0100000000+002  
 22 3.0000000000+000  
 23 1.0000000000+003  
 24 1.37006348283+000  
 25 5.37894812331+003  
 41 3.0000000000+000  
 42 1.0000000000-004  
 43 5.0000000000+001  
 44 1.0000000000+000  
 45 1.0000000000+002  
 46 1.0000000000-008  
 47 5.0000000000-001  
 57 2.0000000000+000  
 58 2.0000000000+000  
 71 1.0000000000+003  
 81 2.0000000000+000  
 82 1.0000000000+001  
 83 6.38131706603-001  
 84 -1.83259571460+000  
 85 9.09665606124-001  
 86 -1.42767932813+000  
 87 1.56945871273-002  
 101 6.5000000000+000  
 102 3.0000000000+002  
 103 6.19999999995+001  
 104 5.0000000000-001  
 150 1.0000000000+000  
 151 2.5000000000+002  
 152 1.0000000000+002  
 153 1.0000000000-001  
 155 1.0000000000+002  
 156 1.0000000000+002  
 201 8.0000000000-001  
 251 3.65000000002+004  
 252 1.0000000000+002  
 253 1.4000000000-001  
 254 3.0000000000+001  
 255 1.39999999999+002  
 256 1.33000000002-002

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG  
 ELEVATION ANGLE OF TRANSMISSION = 0.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000	0.000	0.000	0.000	0.000	0.000	-1.000	0.0000	0.0000	0.0000
-3-011 ENTR ION	52.9737	18.7210	0.000	0.000	-0.000	7.364	0.597	-1.427	823.2593	823.2593	0.0000
-3-011 MIN DIST	158.1469	1.91.1561	0.002	-0.029	-0.753	0.000	0.000	-1.720	1518.9022	1512.7786	0.0106
-3-011 ENTR ION	158.1469	1.91.1561	0.002	-0.029	-0.753	0.000	0.000	-1.720	1518.9022	1512.7786	0.0106
-6-011 WAVE REV	158.1177	1.90.9559	0.003	-0.039	-0.716	0.341	0.000	-1.695	1522.9022	1516.5809	0.0107
8-006 EXIT ION	49.2239	21.85.5251	0.003	0.013	-0.572	-7.138	0.205	-1.093	2239.9022	2227.6590	0.0220
0+000 GRNU REF	0.0000	2.00.0482	-0.000	0.010	-13.042	0.738	0.001	-1.000	2955.4929	2943.2497	0.0220
0+000 ENTR ION	52.9830	3.40.8051	-0.003	0.008	-15.566	7.401	-0.933	1.636	3700.7907	3688.5475	0.0220
-1-005 MAX LAT	107.9072	3.85.1998	-0.702	0.008	-16.445	10.475	-0.000	1.635	4053.7907	4041.5426	0.0247
-1-005 WAVE REV	107.9072	3.85.1998	-0.002	0.008	-16.445	10.475	-0.000	1.635	4053.7907	4041.5426	0.0247
0+000 MIN DIST	157.9016	4.05.0842	-0.007	0.076	-17.366	-0.000	-0.000	9.317	4388.0431	4369.8069	0.0330

THIS RAY CALCULATION TOOK 10.723 SEC

ELEVATION ANGLE OF TRANSMISSION = 15.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000	0.000	0.000	0.000	0.000	0.000	-1.000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9852	18.6.0744	-0.300	-0.000	15.000	16.673	3.119	-3.368	194.1401	194.1401	0.0000
-3-011 MIN DIST	172.1332	2.84.1034	0.015	0.184	12.975	0.000	0.000	-1.081	632.4672	632.4672	0.0081
-3-011 ENTR ION	172.1332	2.84.1034	0.015	0.184	12.975	0.000	0.000	-1.081	632.4672	632.4672	0.0081
2-010 WAVE REV	172.0659	5.07.7222	0.015	0.196	12.864	-0.820	0.000	-1.452	643.8528	635.9500	0.0083
-3-005 EXIT ION	48.8383	10.36.8769	0.042	-0.027	-1.983	-16.239	0.076	-1.037	1106.8528	1085.8928	0.0174
0+000 GRNU REF	0.0000	1.21.9251	0.046	-0.023	-5.495	14.656	-0.002	1.000	1292.1934	1269.1468	0.0174
-3-011 ENTR ION	52.9895	14.03.0468	0.049	-0.020	-4.165	16.366	-1.305	1.885	1490.3081	1467.2614	0.0174
0+000 MIN DIST	171.9560	1.62.6204	0.059	-0.065	-2.959	-0.000	0.000	-1.952	1947.4107	1912.3923	0.0258

THIS RAY CALCULATION TOOK 10.736 SEC

X01 TEST CASE  
CHAPX WAVE

DIPOLY EXPZ2

APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

11/05/74

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 30.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
3-011 XMTR	0.0000	0.0000			30.000	30.000	-0.000	1.000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9739	89.9010	-0.000	-0.000	30.000	30.000	-0.332	1.259	104.6686	104.6686	0.0000
0+000 MIN DIST	191.5641	354.9403	-0.023	0.219	26.398	0.000	0.000	-1.530	425.7924	400.9614	0.0082
0+000 MIN DIST	191.5641	354.9403	-0.023	0.219	26.398	0.000	0.000	-1.530	425.7924	400.9614	0.0082
6-010 WAVE REV	191.4066	358.1069	-0.020	0.137	26.154	-1.395	0.000	-1.475	429.7924	403.6919	0.0085
2-005 EXIT ION	48.7358	144.4661	0.357	-0.397	1.406	-28.975	0.035	-1.018	770.7924	720.5457	0.0177
-3-011 GRND REF	0.0000	733.6080	1.405	-0.349	-3.299	28.173	-0.000	1.000	872.6850	822.4383	0.0177
0+000 ENTR ION	52.9731	130.3382	0.448	-0.308	-0.104	29.043	-0.277	1.223	983.3256	933.0789	0.0177
0+000 MIN DIST	139.8217	1107.5272	1.514	0.454	4.581	0.000	0.000	-1.771	1312.8736	1240.1070	0.0259

THIS RAY CALCULATION TOOK 8.113 SEC

ELEVATION ANGLE OF TRANSMISSION = 45.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000			45.000	45.000	-0.000	1.000	0.0000	0.0000	0.0000
0+000 ENTR ION	52.9759	52.3246	-0.000	-0.000	45.000	45.471	-0.113	1.100	74.6128	74.6128	0.0000
-3-011 RCVR	200.0000	200.2014	-0.132	0.607	43.622	28.480	-0.000	2.293	295.1649	278.4398	0.0058
2-005 APOGEE	209.6834	235.8029	-0.040	-1.746	40.109	3.110	0.000	-1.831	352.1649	306.3131	0.0098
3-006 WAVE REV	209.4300	241.0969	0.021	-1.382	39.431	-0.266	0.000	-1.623	360.1649	310.1892	0.0103
3-011 RCVR	200.0000	274.3788	0.389	-1.415	34.427	-23.456	0.000	-1.142	410.8543	336.8372	0.0137
6-005 EXIT ION	38.3028	144.9524	0.574	-0.027	2.978	-44.472	0.004	-1.000	659.8543	566.5130	0.0206
0+000 GRND REF	0.0000	484.7060	1.576	-0.025	-2.110	44.114	-0.000	1.000	715.5626	622.2214	0.0206
0+000 ENTR ION	52.9777	131.6621	0.575	-0.022	3.158	44.599	-0.105	1.094	791.3392	697.9980	0.0206
-3-011 RCVR	200.0000	691.5163	0.548	0.796	12.769	26.953	-0.000	2.159	1015.3617	905.3752	0.0266

THIS RAY CALCULATION TOOK 9.813 SEC

X01 TEST CASE 11/05/74  
 CHAPX WAVE DIPOLY EXPZ2 APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS

FREQUENCY = 6.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG  
 ELEVATION ANGLE OF TRANSMISSION = 60.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000							0.0000	0.0000	0.0000
0+000 ENTR ION	52.9736	30.2935	-0.000	-0.000	60.000	60.272	-0.053	1.049	61.0908	61.0908	0.0000
0+000 RCVR	200.0000	114.4182	-0.157	-0.851	59.325	50.578	-0.000	1.250	239.2552	225.6343	0.0050
-1-004 APOGEE	225.8332	146.6930	-0.354	5.402	55.872	-5.298	0.000	-2.002	324.2552	253.2681	0.0114
-1-004 WAVE REV	225.8332	146.6930	-0.354	5.402	55.872	-5.298	0.000	-2.002	324.2552	253.2681	0.0114
0+000 RCVR	200.0000	168.1639	-1.383	16.376	+8.744	-62.499	0.000	-1.025	395.1225	273.3158	0.0173
-2-005 EXIT ION	26.7751	230.9196	-6.447	13.564	5.561	-59.411	0.000	-1.000	588.1225	453.6124	0.0223
0+000 GRND REF	0.0000	240.7118	-7.009	13.003	-1.043	69.320	-0.000	1.000	616.7330	482.2229	0.0223
-3-011 ENTR ION	52.9816	260.0662	-7.992	12.021	10.299	69.499	-0.026	1.025	673.3376	538.8275	0.0223
0+000 RCVR	200.0000	313.9111	-10.115	8.555	30.632	63.564	-0.000	1.099	837.7013	691.0808	0.0271

THIS RAY CALCULATION TOOK 7.518 SEC

ELEVATION ANGLE OF TRANSMISSION = 75.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	0.0000	0.0000							0.0000	0.0000	0.0000
0+000 ENTR ION	52.9846	14.0759	-0.000	-0.000	75.000	75.127	-0.024	1.023	54.8374	54.8374	0.0000
0+000 RCVR	200.0000	52.6375	-0.302	1.334	74.783	71.357	-0.000	1.087	213.9181	202.1284	0.0046
1-004 APOGEE	230.9183	67.5504	-1.511	-8.853	73.111	8.237	0.000	-3.876	314.9181	220.4673	0.0134
9-005 WAVE REV	230.2533	70.7691	-0.322	-11.181	72.307	-4.474	0.000	-1.950	330.9181	221.5782	0.0147
0+000 RCVR	200.0000	106.6297	5.511	-13.543	61.034	-46.475	0.000	-1.024	434.9673	249.2747	0.0232
-1-006 EXIT ION	29.3420	220.9308	12.792	-6.856	6.551	-55.885	0.000	-1.000	651.9673	450.5600	0.0292
0+000 GRND REF	0.0000	240.7589	13.557	-6.291	-1.043	55.706	-0.001	1.000	687.4459	486.0386	0.0292
0+000 ENTR ION	52.9836	276.3483	14.370	-5.479	9.565	56.027	-0.069	1.064	751.4563	550.0489	0.0292
-3-011 RCVR	200.0000	375.3507	15.794	-3.644	25.987	45.376	-0.000	1.374	937.9577	722.8911	0.0343

THIS RAY CALCULATION TOOK 9.049 SEC

X01 TEST CASE CHAPX WAVE DIPOLY EXPZZ APPLETON-HARTREE FORMULA EXTRAORDINARY WITH COLLISIONS 11/05/74

FREQUENCY = 9.000000 MHZ, AZIMUTH ANGLE OF TRANSMISSION = 45.000000 DEG

ELEVATION ANGLE OF TRANSMISSION = 90.000000 DEG

	HEIGHT KM	RANGE KM	AZIMUTH DEVIATION		ELEVATION		POLARIZATION		GROUP PATH KM	PHASE PATH KM	ABSORPTION DB
			XMTR DEG	LOCAL DEG	XMTR DEG	LOCAL DEG	REAL	IMAG			
0+000 XMTR	3.0000	0.0000							0.0000	0.0000	0.0000
-3-011 ENTR ION	52.9908	0.0000			30.000	90.000	-0.000	1.000	52.9908	52.9908	0.0000
-3-011 MAX LAT	54.9908	0.0000			30.000	90.000	-0.003	1.012	54.9908	54.9908	0.0000
-6-011 WAVE REV	69.2408	0.0000			30.000	90.000	-0.002	1.019	69.2408	69.2408	0.0000
-3-009 WAVE REV	119.7305	0.0003	-145.262	0.004	30.000	90.000	-0.000	1.018	119.7408	119.7331	0.0010
2-007 WAVE REV	175.1532	0.1157	-145.266	-180.000	89.991	89.993	-0.003	1.021	176.7408	173.9276	0.0029
0+000 RCVR	200.0000	0.3028	-145.266	-180.000	99.911	88.779	-0.000	1.030	207.1658	194.9665	0.0009
4-006 WAVE REV	230.2036	0.9919	-145.266	0.000	99.747	89.207	-0.000	1.052	262.1658	213.5202	0.0009
-3-003 WAVE REV	235.2305	3.6109	-145.266	0.000	99.019	-11.150	-0.003	4.505	310.1658	215.7522	0.0127
-2-003 APOGEE	239.0000	6.0933	-145.266	-0.000	98.978	-33.291	0.003	-5.606	318.1658	215.8346	0.0134
-6-011 RCVR	243.0000	15.3973	-145.266	0.000	95.950	-75.213	0.003	-1.080	411.0570	238.1693	0.0205
2-005 EXIT ION	25.6746	48.6937	-145.266	0.000	27.535	-78.968	0.001	-1.000	586.0570	410.5365	0.0252
-3-011 GRND REF	0.0000	53.7019	-145.266	0.000	-0.242	78.918	-0.003	1.000	622.2176	436.6971	0.0252
-3-011 ENTR ION	52.9956	63.9944	-145.266	0.000	39.224	79.011	-0.003	1.003	676.2116	430.6911	0.0252
0+000 RCVR	200.0000	92.7180	-145.266	0.000	64.368	77.233	-0.003	1.006	833.2790	635.5670	0.0300

THIS RAY CALCULATION TOOK 13.161 SEC

Appendix 8d. Listing of Punched Card Output (ray sets)  
for Sample Case

```

X01 TEST CASE
CHAPX 6.500+000 3.000+002 6.200+001 5.000-001 0.000+000 0.000+000 0.000+000
WAVE 2.500+002 1.000+002 1.000-001 0.000+000 1.000+002 1.000+002 0.000+000
DIPOLY 8.000-001 0.000+000 0.000+000 0.000+000 0.000+000 0.000+000 0.000+000
EXPZ2 3.650+004 1.000+002 1.480-001 3.000+001 1.400+002 1.830-002 0.000+000
X01X 0 40000255000 2000000 60000 4500000 0 0 -1003T
1581469 14911561 2 -29 0 1514389 4513 -1610 11 0 0 -1721M
1581469 14911561 2 -29 0 1514389 4513 -1610 11 0 0 -1722M
1581469 29000482 -0 10 738 2875068 80425 68182 22 0 0 -1003G
1579016 43050842 -7 76 -0 4278561109482 91246 33 0 -0 9323M
X01X 0 40000255000 2000000 60000 4500000 1500000 0 -1003T
1721392 6041034 15 184 0 635731 8122 -3264 8 0 0 -1481M
1721392 6041034 15 184 0 635731 8122 -3264 8 0 0 -1482M
1721418 12129251 46 -23 14656 1211094 81100 58053 17 0 -0 1003G
1719566 18286204 59 -65 -0 1854769 92641 58123 26 0 0 -1953M
X01X 0 40000255000 2000000 60000 4500000 3000000 -0 1003T
1915641 3549408 -23 219 0 407964 17828 -7003 8 0 0 -1531M
1915641 3549408 -23 219 0 407964 17828 -7003 8 0 0 -1532M
1916346 7336080 405 -349 28173 733203139482 89236 18 0 -0 1003G
1898217 11075272 514 454 0 1138430174444101677 26 0 0 -1773M
X01X 0 40000255000 2000000 60000 4500000 4500000 -0 1003T
0 2002014 -132 607 28480 285194 9971 -6754 6 0 -0 2291R
2096843 2743788 389 -1415-23456 342980 67874 -6143 14 0 0 -1142R
2096843 4847060 576 -25 44114 484589230974137632 21 0 -0 1003G
0 6915163 548 796 26953 729880285982175495 27 0 -0 2163R
X01X 0 40000255000 2000000 60000 4500000 6000000 -0 1003T
0 1144182 -157 -851 50578 231305 7950 -5671 5 0 -0 1251R
2258382 1681639 -1383 16876-62499 262993132130 10323 17 0 0 -1032R
2258382 2407118 -7009 13003 69320 240697376036241525 22 0 -0 1003G
0 3139111-10115 8555 63564 376316461386314765 27 0 -0 1103R
X01X 0 40000255000 2000000 60000 4500000 7500000 -0 1003T
0 526875 -302 1334 71357 207034 6884 -4906 5 0 -0 1091R
2309183 1066297 5511-13543-46475 227435207533 21840 23 0 0 -1022R
2309183 2407589 13557 -6291 55706 240745446701245294 29 0 -0 1003G
0 3753507 15794 -3644 45376 430430507528292462 34 0 -0 1373R
X01X 0 40000255000 2000000 60000 4500000 9000000 -0 1003T
0 3028214734180000 88779 200000 7166 -5034 5 0 -0 1031R
2382305 153973214734 0-75213 200610210447 37559 20 0 0 -1082R
2382305 537019214734 0 78918 53702568516382995 25 0 -0 1003G
0 927180214734 0 77233 221057612222414510 30 0 -0 1013R

```

- The first card is the title card.
- The second card contains the name of the electron density model plus parameters W101-W107.
- The third card contains the name of the perturbation model plus parameters W151-W157.
- The fourth card contains the name of the magnetic field model plus parameters W201-W207.
- The fifth card contains the name of the collision frequency model plus parameters W251-W257.

For description of remaining cards, see figures 1 and 2.



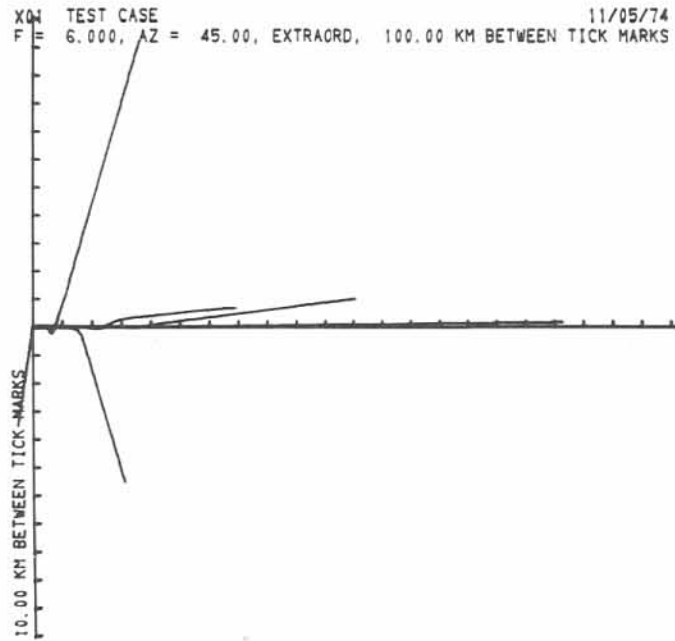
Appendix 8e. Ray Path Plots for Sample Case

Projection of raypath on vertical plane

X01 TEST CASE 11/05/74  
F = 6.000, AZ = 45.00, EXTRAORD, 100.00 KM BETWEEN TICK MARKS



Projection of raypath on ground for sample case





## BIBLIOGRAPHIC DATA SHEET

1. PUBLICATION OR REPORT NO. OTR 75-76		2. Gov't Accession No.	3. Recipient's Accession No.
4. TITLE AND SUBTITLE A versatile three-dimensional ray tracing computer program for radio waves in the ionosphere		5. Publication Date October 1975	
7. AUTHOR(S) R. Michael Jones and Judith J. Stephenson		6. Performing Organization Code OT/ITS, Div. 1	
8. PERFORMING ORGANIZATION NAME AND ADDRESS U. S. Department of Commerce Office of Telecommunications Institute for Telecommunication Sciences Boulder, Colorado 80302		9. Project/Task/Work Unit No.	
11. Sponsoring Organization Name and Address U. S. Department of Commerce Office of Telecommunications Institute for Telecommunication Sciences Boulder, Colorado 80302		10. Contract/Grant No.	
14. SUPPLEMENTARY NOTES		12. Type of Report and Period Covered Technical Report	
15. ABSTRACT (A 200-word or less factual summary of most significant information. If document includes a significant bibliography of literature survey, mention it here.) This report describes an accurate, versatile FORTRAN computer program for tracing rays through an anisotropic medium whose index of refraction varies continuously in three dimensions. Although developed to calculate the propagation of radio waves in the ionosphere, the program can be easily modified to do other types of ray tracing because of its organization into subroutine. The program can represent the refractive index by either the Appleton-Hartree or the Sen-Wyller formula, and has several ionospheric models for electron density perturbations to the electron density (irregularities), the earth's magnetic field and electron collision frequency. For each path, the program can calculate group path length, phase path length, absorption, Doppler shift due to a time-varying ionosphere, and geometrical path length. In addition to printing these parameters and the direction of the wave normal at various points along the ray path, the program can plot the projection of the ray path on any vertical plane or on the ground and punch the main characteristics of each ray path on cards. The documentation includes equations, flow charts, program listings with comments, definitions of program variables, deck set-ups, description of input and output, and a sample case.		13.	
KEY WORDS: Appleton-Hartree formula; computer program; ionosphere; radio waves; ray tracing; Sen-Wyller formula; three-dimensional.			
17. AVAILABILITY STATEMENT <input checked="" type="checkbox"/> UNLIMITED. <input type="checkbox"/> FOR OFFICIAL DISTRIBUTION.		18. Security Class (This report) UNCLASSIFIED	20. Number of pages 197
		19. Security Class (This page) UNCLASSIFIED	21. Price:

OT REPORT 75-76

A VERSATILE THREE-DIMENSIONAL RAY TRACING COMPUTER PROGRAM  
FOR RADIO WAVES IN THE IONOSPHERE

by R. Michael Jones and Judith J. Stephenson

The 8th line from the bottom of page 9 should read:

(1) the dispersion relation cannot be exactly satisfied, or

The first line in the 3rd complete paragraph on page 12 should read:

Similarly, the AHWFNC (Appleton-Hartree, with field, no colli-

Line PRIN125 in SUBROUTINE PRINTR on page 80 should read:

RANGE=EARTH\*ATAN2(RCE,EARTH+EPS+XMTRH) PRIN125

Line BQNC020 in SUBROUTINE BQWFNC on page 100 should read:

REAL N2, NNP, LPOLAR, LPOLRI, KR, KTH, KPH, K2, KDOTY, K4, KDOTY2, BQNC020

Line TABX064 in SUBROUTINE TABLEX on page 112 should read:

PXPR=PXPTH=PXPPH=0. TABX064

Following line CHAP024 in SUBROUTINE CHAPX on page 116, insert the line:

PXPPH=0. CHAP0245

Line VCHA010 in SUBROUTINE VCHAPX on page 117 should read:

X=PXPR=PXPTH=PXPPH=0. VCHA010

Line DCHA014 in SUBROUTINE DCHAPT on page 119 should read:

X=PXPR=PXPTH=PXPPH=0. DCHA014

Line LINE013 in SUBROUTINE LINEAR on page 120 should read:

X=PXPR=PXPTH=PXPPH=0. LINE013

Line PARA012 in SUBROUTINE QPARAB on page 121 should read:

X=PXPR=PXPTH=PXPPH=0. PARA012

Following line BULG038 in SUBROUTINE BULGE on page 123, insert the line:

PXPPH=0. BULG0385

Following line EXPX014 in SUBROUTINE EXPX on page 124, insert the line:

PXPTH=PXPPH=0. EXPX0145

The equation for the gyrofrequency near the top of page 143 should read:

$$F_H = F_{H_0} (R_0 / (R_0 + h))^3 (1 + 3 \cos^2 \theta)^{1/2}$$

where  $\theta$  is the geomagnetic colatitude.

Line TABZ015 in SUBROUTINE TABLEZ on page 153 should read:

IF (READNU.EQ.0.) GO TO 10 TABZ015