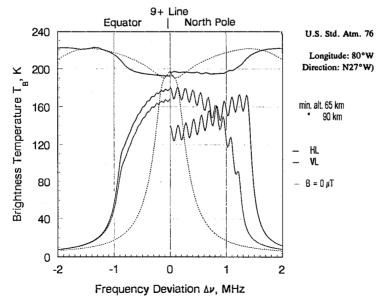
F3-8
POLARIZED TRANSMISSION AND EMISSION IN THE
MIDDLE ATMOSPHERE (25 - 150 km) NEAR THE
ZEEMAN-SPLIT MICROWAVE O₂ LINES
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Gaseous atmospheric attenuation and delay properties in the 50 to 75-GHz band and near 119 GHz are dominated by the fine structure lines of molecular oxygen. With increasing altitude the complex refractivity displays spectral patterns which change from an unstructured band to the isolated shapes of about 35 lines. The earth's magnetic field (22 - 65 μ Tesla) splits each line into many sublines. The anisotropic nature of these Zeeman components causes polarization discrimination and Faraday rotation.

The "Zeeman-effect" Propagation Model ZPM estimates the path transmission and emission that originates between 25 and 150 km near isolated $\rm O_2$ lines. Many aspects of how polarized, plane radio waves propagate through a spherically stratified model atmosphere are evaluated (i.e., complex refractivity tensor, path-specific attenuation rates, Faraday rotation, polarization, and optical depth). Related to the absorptive line properties is thermal emission. The example below shows emission near the 61.150-GHz line frequency, as seen by a linearly polarized (HL/VL) pencil-beam antenna looking from outer space into the atmosphere to minimum heights of 65 and 90 km above sea level at either the equator or the north pole. Radiometers of limb-sounding experiments conducted from space platforms (e.g., UARS-MLS) and ATLAS-MAS) are able to detect such emissions features, which reveal information on $\rm O_2$ density, ambient temperature, and pressure.



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National Radio Science Meeting 5-8 January 1993

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University of Colorado Boulder, Colorado U.S.A.

National Radio Science Meeting 5-8 January 1993 Condensed Technical Program

Monday, 4 January

2000-2400 USNC-URSI Meeting	Broker Inn
Tuesday, 5 January	
0835-1200 B-1 ANTENNAS	CR2-28
0855-1200 A-1 MICROWAVE MEASUREMENTS D-1 MICROWAVE, QUASI-OPTICAL, AND ELECTROOPTICAL DEVICES G-1 COORDINATED CAMPAIGNS AND ACTIVE EXPERIMENTS J/H-1 RADIO AND RADAR ASTRONOMY OF THE SOLAR SYSTEM	CR1-46 CR1-9 CR0-30 CR2-26
1335-1700 B-2 SCATTERING F-1 SENSING OF ATMOSPHERE AND OCEAN G-2 IONOSPHERIC PROPAGATION CHANNEL J/H-2 RADIO AND RADAR ASTRONOMY OF THE SOLAR SYSTEM	CR2-28 CR2-6 CR0-30 CR2-26
1355-1700 A-2 EM FIELD MEASUREMENTS D-2 OPTOELECTRONICS DEVICES AND APPLICATION	CR1-46 CR1-9
1700-1800 Commission A Business Meeting Commission C Business Meeting Commission D Business Meeting Commission G Business Meeting	CR1-46 CR1-40 CR1-9 CR0-30
Wednesday, 6 January	
0815-1200 PLENARY SESSION	MATH 100
1335-1700 B-3 EM THEORY D-3 MICROWAVE AND MILLIMETER AND RELATED DEVICES F-2 PROPAGATION MODELING AND SCATTERING H-1 PLASMA WAVES IN THE IONOSPHERE AND THE MAGNETOSPHERE J-1 FIBER OPTICS IN RADIO ASTRONOMY	CR2-28 CR1-9 CR2-6 CR1-42 CR2-26
1355-1700 E-1 HIGH POWER ELECTROMAGNETICS (HPE) AND INTERFERENCE PROBLEMS	CR1-40

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PROGRAM AND ABSTRACTS

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